AIRCRAFT VORTEX WAKE DESCENT AND DECAY UNDER REAL ATMOSPHERIC EFFECTS
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$\qquad$
AeroVironment, Incorporated

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 entration, wake descent and stability, and atmorphertc dynamics are considered.

Operational equations for ercouster hazard, wake generation, and atrespheric dynamics are given, includng \& brief description of possible automatic meteorological system to provide atmos. pheric data for in airport wake fniecasting progran.

A new analysis for Ccow Instability in umbient turbulrnee is given, expressing time-to-linkage as an explicit function of the tests although only limited dita is available,

Hake ciescent in a stracified inviscid fluid is studied analylically providing new results for thie problen. According to the present theory, the vortex span reduces upon descent into a stably present theory, the vortex span reduces upon toscent ified flow, causing the rate of tescent to increase. Exact stiutions jre derived for vortex cell shapes in a uniformly sheared crosswind, showing that the upwind cell is greatly increased in size. It is belicvod that this may partly account for the observed unsymatrical behavior (banking, etc,) in crosswinds.

A discussion of core bursting ind turbulent wake entrainfent during cescent is given, with some tentative formulations for the lattar. Fuli understanding of these two aspects must still be considered inconplete.

Finaliy, an assescment of the remaintms problems is given, with recommendations for further analytical and flight test re: search.
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There exists a massive body of literature on aircraft wake vortex research. As an aid for further work, a new bibliography has been prepared. The base of this list has been taken from the excellent listing given by McCormick, 1971: Aircraft Wakes: A Survey of the Problem (FAA Symposium on Turbulence, Washington, D. C.). This base has been complemented and updated by AeroVironment and is believed to be very comprehensive. The actual report refers to only a portion of these papers, and contains other references which are not directly concerned with vortex wakes. Thus, for ease of reading, a separate reference list has been prepared, containing only those papers quoted in this report.

It is pointed out that thiz is a first report, and by the natur 3 of this particular subject, contains hypotheses and suggestions for which conclusive experimental data is still lacking. A second report will appear later in 1973 by which time more analytical and experimental data will be available. It is believed that this new iesearch should settle some of the crucial questions identified in this report.

## TABLE OF CONTENTS

1. INTRODUCTION ..... 1-1
2. LIFT AND VORTEX. GENERATION ..... 2-1
3. THE ENCOUNTER HAZARD ..... 3-1
4. VORTEX TRANSPORT ..... 4-1
4.1 General ..... 4-1
4.2 Neutral Irrotational Field ..... 4-6
4.3 Neutral Rotational Field ..... 4-18
5. 4 Wake Transport in a Stratified Field ..... 4-33
4.4.1 Experimental Observations ..... 4-33
4.4.2 Previous Theoretical Models ..... 4-39
4.4.3 New Theory for Trailing Vortices in a Stably Stratified Atmosphere ..... 4-50
6. VORTEX DECAY ..... 5-1
5.1 General ..... 5-1
5.2 Linking Instability ..... 5-3
5.3 Core Bursting ..... 5-23
7. METEOROLOGICAL ASPECTS ..... 6-1
6.1 General ..... 6-1
8. 2 The Quantities of Interest ..... 6-4
6.3 Boundary Layer Relationships ..... 6-8
6.4 A Review of Operational Equations ..... 6-30
6.5 An Approach to an Operational Data System ..... 6-35
9. SUMMARY ..... 7-1
10. CONCLUSIONS ..... 8-1
11. REFERENCES ..... 9-1
i0. BIBLIOGRAPHY ..... 10-1

## LIST OF ILLUSTRATIONS

Figure Page
2-1 Stages of Wake Vortex Development ..... 2-8
4-1 Viscous Effeces on Vortex Cell ..... 4-8
4-2 Vortex Cell Stages ..... 4-10
4-3 Non- Dimensional Vortex Transport Curves ..... 4-16
4-4 Coordinate System and Flow Geometry for Vortex-Pair in Uniform Cross-Flow ..... 4-20
4-5 Streamlines in the Vicinity of a Vortex Pair with no Shear. Shear Parameter, $\sigma=0$. ..... 4-24
4-6 Streamlines in the Vicinity of a Vortex Pair in Light Shear. Value of the Shear Param- eter, $\sigma=1.0$ ..... 4-25
4-7 Streamlines in the Vicinity of a Vortex Pair in Moderate Shear. Value of the Shear Parameter, $\sigma=2.0$ ..... 4-26
4-84-9Streamlines in the Immediate Vicinity of theTop Stagnation Point for a Vortex Pair inHeavy Shear. Value of the Shear Param-eter $\sigma=3.0$4-28
$4-10$ Trajectory of a Vortey Pair $n$ a Linearly Density Stratified Medium ..... 4-36
4-11 Formation of Drift Behind a Vortex Pair ..... 4-55
4-12 Calculation of Drift ..... 4-56
4-13 The Dimensionless Drift, $\eta(\xi)$ ..... 4-59
4-14 Density Field Above the Wake ..... 4-61

## I. ${ }^{\text {SST OF }}$ ILLUSTRATIONS (Cont.)

| Figure |  | Page |
| :---: | :---: | :---: |
| 4-15 | Profiles of the Buoyant Upwash | 4-64 |
| 4-16 | Induced Field of the Buoyant Upwash | 4-65 |
| 4-17 | Framework for Calculating the BuoyancyInduced Field at a Vortex Core . . . . . . . . . . . | 4-67 |
| 4-18 | Wake Trajectory for the Case $\Gamma=9000 \mathrm{ft}{ }^{2} / \mathrm{sec}$. . $2 S_{0}=110 \mathrm{ft} ., N=0.035 \mathrm{sec}^{-1} \ldots . . . . .$. | 4-83 |
| 4-19 | Flow Within the Contracting Recirculatio. Cell . | 4-85 |
| 5-1 | Geometry of an Oscillating Vortex Pair | 5-7 |
| 5-2 | Non-Dimensional Vortex Pair "Lifetime" to Linking, Eqn. 5-32 ................... | 5-17 |
| 5-3 | Time to Vortex Pair Linking as a Function of Atmospheric Turbulent Dissipation Rate, $\epsilon$, for take off . | 5-18 |
| 5-4 | Comparison of Theoretical Prediction of Time-to-Linking with Experiment . . . . . . . . . . . . . . | 5-20 |
| 5-5 | Vortex Wake Streamlines and Core Geometry | 5-26 |
| 5-6 | Shed Vorticity in Transverse Plane | 5-28 |
| 5-7 | Pressure and Velocities Near Representative <br> Vortex . | 5-33 |
| $5-8$ | Estimated Vortex Breakdown Diagram | 5-38 |
| 6-1 | Universal Function $\Psi$ for the Integrated Wind Profile . . . . . . . . . . . . . . . . . . . . . . . . . | 6-14 |
| - - 2 | 1/L as a Function of Pasquill Classes and zo | 6-17 |

Table Page
4-1 The Ordering of the Response for Four Cases of Similarity Solutions ..... 4-14
4-2 Comparison of Theoretical Mudels for Descent of a Vortex Wake in a Stabīy Statified Atmosphere ..... $4-39$
6-1 $R_{i}-z / L^{\prime}$ Relationship ..... 6-11
6-2 Relation of Turbulence Types to Weather Condi- tions ..... 6-16
6-3 Data Summary ..... 6-16

## 1. INTRODUCTION

The hazards posed to following aircraft by the vortex wake system of a large airplane are v/ell-known. In order quantitatively to assess the hazard, as an aid in designing wake avoidance systems, it is evidently necessary to know where the wake is, what its strength is, and what the danger to an encounter aircraft is. These three aspects are addressed in this report.

The hazard determination depends upon the strength of the wake and the response/control characteristics of the encounter aircraft. For a given wake this becomes a problem in aircraft dynamics, and the specialized aspects of this are not treated here, but in Section 3 sinple danger indices are derived. The technology is vailable for the refinement of these indices, but in view of the uncertainties of wake strength and position, such extra precision may not be required.

The main thrust of this report is directed at the determination of wake strength and position. These characteristics are fundamentally connected with the initial parameters of the wake. Section 2 describes the mechanism of vortex sheet generation and roll-up. It is believed that this aspect, essentially an aerodynamic problem, is reasonably well understood and quantified.

After roll-up, we are concerned with the descent and decay of the wake in a real atmosphere. This atmosphere contains omnipresent turbulence, buoyancy, wind shears and wind gusts, Effects of these dynamic characteristics of the atmosphere on the vortex wake are very poorly understood and experimental data is limited. The major portion of this report, Sections 4 and 5, addresses itself to the principal elements of of this interaction, with the goal of arriving at quantitative expressions for these effects so that operational predictive equations can be derived.

Finally, to specify the atmospheric inputs, it is necessary to be able to infer or predict the turbulence, shear and other atmospheric dynamic properties from the basic meteorological observables. Section 4 shows how this may be done by using various earth boundary layer models and the large body of statistical meteorological data.

## 2. LIFT AND VORTEX GENERATION

The process of lift generation and vortex shedding on a continuous wing is well understood. Thwaites ( 1960 ) gives an excellent review of modern wing theory. We note that the determination of the spanwise and chorduise loading on a wing is not of concern here, but that a vast volume of literature treating this subject exists anf medern lifting suríace theory (as employed by all the major ai faf: $=0 \mathrm{~m}$ panies) gives a prediction of the wing characteristics ardjerformances which correlates very well with experimental tests. For cur pirpose we are concerned with the way in which the trailing vortex core system is connected to the bound vorticity on the wing. For the moment, we will discuss a wing in attached flow with a continuous trailing edge; that is, where there are no special disturbances due to fiaps or propulsion or fuselage attachments.

The vorticity on the wing itself is principally of f . $2 n w i s e$ orientation and associated with the wing boundary layer. if the wing is $=$. zero lift this vorticity is shed oriented approximately parallel to the trailing edge and results in a low energy planar wake in which the velocity perturbations are mainly in the direction of the free stream, and represent a reduction in the free stream velocity associated with the viscous drag of the wing. The momentum loss in this wake can be directly connected to that in the boundary layer and is represented in an integrated form in the profile drag coefficient of the wing. $\mathrm{C}_{\mathrm{Do}_{0}}$. The downstream development in the non-lifting case consists of entrainment of outer flow air and a general coalescence of the wake crosssection from a sheet to an ellipse and finally a circle. However, these details are not important since we are interested in the lifting case, when the downstream development is dominated by the lift induced effects.

If the wing is lifting, an additional vorticity component is generated at the trailing edge. This vorticity is shed parallel to the local
velocity vertor, that is, approximately in the direction of the free stream. It can be shown that the vorticity shed at each spanwise station is exactly equal to the gradient ir spanwise lift at that station. It is this vorticity which rolls up to form the trailing vortex pair ard with which the induced drag $D_{i}$ is associated. This vorticity creates (or is created by) the sidewash and downwash fields downstream of the wing and does not directly produce streamwise flows; however, as is shown later, viscous effects control the rollup of this vortex system and ultimately cause streamwise velocity perturbations which couple with the profile drag vorticity to create axial flows in the vortex core.

Continuing the aralysis of classical wing theory, a number of significant (ard exact) results can be developed for the relationships between spanwise loading, irduced drag, and the final vortex position.

We write the spanwise circulation loading as

$$
\Gamma=2 \mathrm{Ub} \sum_{\mathrm{l}}^{\mathrm{m}} \mathrm{~A}_{\mathrm{n}} \operatorname{Sin} n \theta \quad, \quad n \text { odd }
$$

where $\cos \theta=r$ is the normalized span coordinate, $U$ is the fight speed, and b the wing span. For symmetrical loadings only odd $n$ is required. Then the center section circulation $\Gamma_{0}$ is given by

$$
\Gamma_{0}=2 U b \sum_{1}^{p}(-1)^{\frac{n-1}{2}} A_{n}
$$

while the lift $L$ and induced drag are given by the formulas

$$
\begin{aligned}
& L=q b^{3} \pi A_{1} \\
& D_{i}=q b^{2} \pi \sum_{1}^{\infty} n A_{n}^{a}
\end{aligned}
$$

where $q$ is the dynamic pressure $1 / 2 \rho U^{2}$. A significant term in the downstream core development is the vortex span $b_{v}$. This car be shown to be equal to the centroid of the shed vorticity (for conservation of lift), and thus we obtain

$$
b_{v}=\frac{\pi b}{4} \frac{A_{1}}{\Sigma(-1)^{\frac{n-1}{2}} A_{n}}=\frac{4}{\pi} \cdot \frac{1}{0 U T_{0}}
$$

These simple results are usually derived from lifting line theory, but it is worth noting thar they are exact for all planar wings (even of low aspect ratio) providing that vorticity is shed only from the trailing edge.

It is usual to consider only elliptically leaded wirss irsplying $A_{n}=0, n>1$. We remark that for any continuous loading the additional coufficients may be found by Fourier analysis and the more precise results determined. However, deviations from the elliptical loading for aircraft wings in the clean condition are usually small. As an illustration we quote resuits from Thwaites for a number of different untwisted planforms. We exhibit these as factors for comparison with the elliptical results. Thus for an untwisted elliptical wing we get

$$
D_{i}=\frac{L^{2}}{q \pi b^{2}} \quad, \quad b_{v}=\frac{\pi b}{4}
$$

while for the other wings we write

$$
D_{i}=K_{D} \frac{L^{2}}{q \pi b^{2}} \quad, \quad b_{v}=K_{v} \frac{\pi b}{4}
$$

Ty pical extreme cases are listed below for a wing of aspect ratio 6 .

Straight Constant Chord Wing

| $\therefore \mathrm{D}$ | $\mathrm{k}_{\mathrm{v}}$ |
| :---: | :---: |
| 2.05 | 1.11 |
| 1.14 | . 82 |
| 1.08 | 1. 12 |

We note that these cases giva quite different spanwise vorticity loadings. Compared with the elliptical, the pointea wing has increased inboardi loadiag. The swept wing has inboard 'cading significantly lower than that at abou. iof span (the so-called saddleback loading), while the rectangula.: wing has celatively tigh tip loading. However, even in these cxtreme cases (within normal wing proportiona) there is not very much difference in the final vorte ispan or the induced drag.

It should also be resalled that most aircraft wings are aerodynamic; $f^{f} \dot{f}$ sted so that the distribution at cruise will be nearly elliptical ... thus they will not exhibit such extreme loading distributions as the untwisted planforms noted above.

On these grounds we note that the elliptical assumption is certainly adequate for most practical wings for our purposes, especially in view of the mush cruder atsumptions mherent in analysis of the vortex wake development. However, hif requirat, the precise determination of the global vorter propsrties can readily be made by the formulae above.

When tie wing is in the dirty condition, with flaps sxtended, the remarks about elliptical loading no longer apply. However, the chree equations for $L, D_{i}$ and $b_{v}$ still obtain and the calculations can still be conducted. With part span flaps the circization distribution contains gradiente which are 1 erawithmically singular at the discontinuity, which means that many terms in the sinc expansion must be used for an accurate determination of $D_{i}$ and $b_{v}$. This can be avoided by introducing
an acditional circulation function capable of exactly matching the discontinuities in airfcil effective angle at the flap extremities. Using this stratagem a very arcurate solution can be obtained with only a few terms. A modern development of this technique applicable to chord and twist discontinuities is given by Lissaman (1973). This method has been programmed and is available at NASA (Iangley). It represents an exact solution for the problem discussed here.

The effect of inboard flaps is greatly to increase $k_{D}$ and to reduce $k_{v}$. These changes are such that the elliptical estimate is certainly no longer valid, and each flap type circulation loading should be analyzed as described above.

As an illustration of this effect, take a wing of aspect ratio 5 with $30 \%$ chord half-span inboard flaps deflected at $8^{\circ}$. Taking the lift distribution from a similar arrangement given by Bilanin and Wianall (1973) we obtain:

Straight Constant Chord Flapped Wing

| $k_{D}$ | $k_{v}$ |
| :--- | :--- |
| 1.31 | .757 |

The circulation loading distribution of a wing defines all the profurties of the inviscid vortex development. However, as pointed out above, for most vortex development purposes only a few characteristics of this distribution are required; the zero moment for the lift, and the center section bound vorticity for the centroid of lift. In addition, we require the induced drag which theoretically involves all moments and is actually quite strongly dependent on highe: terms in the sine expansion

One further moment will prove useful; the moment of momentum, that is. the angular momentum of the fluid rotating around the vortex sheet. It has been shown by Betz (1933) that this can be conveniently evaluated by taking the second soment of the shed vorticity about any
spanwise position. A convenient position is the centroid of this vorticity. This term $M_{m}$ can also be expressed compactly in the Fourier coefficients as

$$
M_{m}=\rho b^{3} U^{3} K_{m}
$$

and

$$
K_{m}=\sum_{1}^{\infty} \frac{(-1)^{\frac{n+1}{2}} A_{n}}{n^{3}-4}-\frac{\pi^{3}}{64} \frac{A_{1}^{2}}{\sum_{1}^{\infty}(-1)^{\frac{n_{1}-1}{2}} A_{n}}
$$

Physically, $M_{m}$ has the dimensions of a moment, and represents the moment applied to the air about the centroid of vorticity by one sife of the wing. It can quite easily be evaluated mechanically if the spanwise loaring is known by taking the moment of this load about the centro: $\exists$ of vorticity.

We now not: that we have developed expressions for the six major parameters of the trailing vortex sheet. These are the lift, the induced drag, the moment of mom sntum, the vortex span, the total bound vorticity, and the p.ofle drag, The first five of these can be expressed directly from an analysis of the spanwise loading curve. Only four of these are actually irciependent in that $\Gamma_{0}$ can be determined from $b_{V}$ and $L$. The profile dray $D_{0}$ can readily be determinod from standard drag determination nettends, The total cirag $D_{T}=D_{o}+D_{i}$.

Thus we can consider the vorticity sheet shed at the wing tr:ining ecige to be determined (in a global sense) by $L, D_{T}, M_{m}, b_{v}, \Gamma_{0}$.

## Rollup and Development of the Vortex Sheet

After leaving the wing, the shed vorticity undergoes a series of complicated convolutions until it is finally diffused into the atmosphere. We can arbitrarily define four stages to this process, in which fairly distinct interactions occur. These are shown in Figure 2.1 and can be thought of as the Planar Sheet Stage, the Rollup Stage, the Viscous Vortex pair stage, and the Distributed Impulse stage. Our main interest is in the latter two. They are, of course, controlled by the initial conditions, which come from the Planar Stage discussed in the previous section.

After leaving the wing, various mutual inductions cause the vortex sheet to curl up at the tips and to form a scroll-like shape and initiate the rollup stage. Within about 20 spans this has developed into a viscous vortex pair of strength and span $\Gamma_{0}, b_{v}$ and having a finite core within which the vorticity is contained. Many authors (see Bibliography) have attempted to describe this rollup process but the proper analysis is still highly controversial. We will not specifically discuss these analyses here except to state that the results of numerical free vortex schemes should be regarded with caution. While these techniques in principal exactly treat the self-induced development of the trailing vortex sheet, the necessary discretization for computer operation may be expected to introduce instabilities and motions reIated more to the discretization and computer scheme than to the actual deformation of the vortex sheet. This is particularly true if the vortex sheet is modeled by a set of discrete line vortices. For our purposes we are not concerned with the details of the rollup process because this occurs so close to the wing.

The next stage is the viscous vortex development, which is discussed in more detail in following chapter, and finally the decay to


Figure 2-1. Stages of Wake Vortex Development
the distributed impulse (also considered in more detail in the Transport Section, 4).

To assist in the discussion of this development, we describe the various global invariants in the process. Evidently the lift and drag must be conserved throughout the development. Now we note the other terms; shed vorticity and moment of momentum are actually null for the entire wing and can be defined only for one half wing. However, provided there is no viscous interaction between the two sides (left and right) of the vortex system, then even in a viscous fluid we will find that $M_{m}$ and $\Gamma_{0}$ are conserved. We note by the kinematic condition of the solenoidal quality of vorticity, that even though the vorticity is diffused into originally irr otational flow, $\Gamma_{0}$ must be conserved until vorticity from the one core interacts with the other on the center line. The moment of momentum is conserved according to the same arguments. A physical interpretation of this is that no external torque can be applied to the vortex core and its surrounding flow system until appreciable vorticity has reached the center line. We comment here that as long as the flow on the center line is symmetrical even introducing viscous transfer in the equations will produce no shear along the center line since velocity gradients there are zero.

Thus during the rollup and the viscous pair stage a good as sumption may be to conserve $M_{m}$ and $\Gamma_{0}$. In the final decaying impulse stage we will expect $M_{m}$ and $\Gamma_{0}$ to decay according to transfer processes, both of molecular viscosity and turbulence.

We have not discussed the variation of $H_{0}=p / p+\frac{1}{2} v^{2}$, the total head of the flow. This provides a local flow property (as opposed to global properties) which is useful to us. We note that at the wing trailing edge there is already a sheet of low energy air ahed from the boundary
layer, having dimensions scaled by the span and the trailing edge boundary layer displacement thickness, and containing a total head deficit relating to the boundary layer momentum thickness. This wake will entrain external flow in the way normally associated with a non-lifting velocity deficit wake.

Of more interest is the entrainment and head loss due to the rotational flow associated with the sidewash and downwash flows connected to the induced drag. This effect features in the swirling vis cous interaction which develops the axial flow in the vortex core. Far convenience we define $H_{L}$ as the head loss connected with the lifting flow. It seems that this term becomes significant only in the vortex pair and decay stages.

It is noted that the global force and local energy relationships are not sufficient to solve for the details of the vortex development. However, they provide a useful insight into the processes occuring and can be very helpful in assessing different theoretical analyses, most of which contain direct or implicit statements relating to assumptions concerning these quantities.

## 3. THE ENCOUNTER HAZARD

Evidently, if conditions are such that an aircraft encounters a vortex wake while the wake is atill in the organized state, consisting of two near-axial viscous vorticss, the hazard will be a function of some parameters describing the strength and geometry of the wake, coupled with further factors relating to the control power and design load factors of the encounter aircraft. Other parameters in this situation are the flight path relative to the wake of the generating aircraft, and, of course, the proximity to the ground of the encounter.

Initially, we shall consider the case of the encounter aircraft entering the eye of one of the trailing vortices on a flight path approximately parallel to the vortex axis. In these circumstances a severe rolling moment will be induced on the encounter aircraft. It would at first appear that the magnitude of the rolling moment would be closely related to the core development, that is, the vortex age. However, when one assumes that the core will be less than $2 / 10$ of the generating span, and that outside this the flow very closely approaches that of a horseshoe vortex, we see that much of the encounter aircraft wing for an encounter span of about $1 / 2$ the vortex span, will be in the inviscid portion of the flow. Coupling this with the fact that the contributions to rolling moment vary directly with the spanwise position, we may expect that the actual size of the core will not be too important providing it is less than a certain fraction of the encounter span. On this basis, Crow (1971) has shown that the induced rolling moment will be proportional to $\Gamma_{g} / b_{e}^{a}$, where $\Gamma_{g}$ is the strength of the generator aircraft vortex and $b_{e}$ the span of the encounter aircraft. If it is then assumed that there is a certain maximum aileron induced roll rate of the encounter aircraft, $P_{m a x}$, one can assert that hazardous conditions will be obtained when she ratio of vortex induced roll to aileron rolling power exceeds a ceitain value. Writing the trailing vortex
strength in terms of lift coefficient, $C_{L}$, span $b$, and aspect ratio, $A$ we define this ratio below.


#### Abstract

If the critical rolling rate of the encounter aircraft is $P_{\text {max }}$ (in radians persecond), the non-dimensional danger factor $D$, expressing the hazard, can be written as


$$
D=\left(b \cup C_{L} / \pi A\right)_{g} /\left(b^{2} P_{\max }\right)_{e}
$$

where the suffice $g$ applies to the generator aircraft and $e$ to the encounter aircraft. It has been estimated by Crow from Boeing tests by Condit and Tracy (1971) that a value for $D$ of about unity is a good measure of critical hazard. If $D$ exceeds this value, dangerous bank attitudes may be expected. At low altituies a smaller $D$ value will be appropriate since recovery time is reduced.

The danger factor has a very compact formulation and it will be noted that it contains no term relating to the distance behind the generator aircrafi. This is a significant simplification and relates to the slow growth of the core and the invariance of the total circulation. A more detailed but still idealized analysis of the induced rolling moment conducted by Condit and Tracy (1971) shows the separation distance dependence. Taking the results of their numerical calculations, we see Shat the danger factor is in fact almost invariant with separation distance. For example, using dimensions appropriate to a B707 we finc that the more precise analytical calculations will give a normalized reduction in induced rolling moment of about $2.5 \%$ per nautical mice behind the generator aircraf. Thesc numbers are substantiated by flig'it tests on Boeing aircraft reported in the can. paper.

It is noted that at first this appears to contradict the results of Bisgood, Mallby and Dee (ig71) who use Squire's Theory to compute
wake dissipation and exhibit a theoretical curve showing a rather marked reduction in rolling rate with vortex age. This curve is for an assumed eddy viscosity of $=004 \Gamma_{0}$ ( $\Gamma_{0}$ being the vortex strength). However, the paper shows eddy viscosity inferred from flight measurement of core velocity peaks, indicating that the effective eddy viscosity according to this theory is more like . 0002 $\Gamma_{0}$ with some readings as low as. $00018 \Gamma_{0_{0}}$ Using the value of $.0002 \mathrm{~F}_{\mathrm{o}}$ and conducting a similar analysis would show the reduction in rolling rate to be about $5 \%$ per nautical mile, which is more in line with the more modern and comprehensive Boeing tests.

Now, we should consider the other real factors associated with vortex entry in a near axial direction. These relate to contro! -eversals as the encou:ter aircraft traverses the vortex laterally or enters the other vortex, as well as the coupled yaw-roll dynamics associated with the effect of the vertical stabilizer. Such excursions due to random oblique encounters will introduce hazard elements of more significance than those computable by simple axial vortex decay and thus justify ignoring the latter effect.

On thege grounds it is believed that the danger factor $L$ represents a useful criterion for the vortex hazard due to nesr axial encounterg, and the assumption that $D$ in invariant with separation distance is a significant and valid simplification. The critical value of $D$ will depend on the aircraft ground proximisy.

It must be remerked here that while the danger lactor is invariant with vortex age or separation distance, its existence depends upon there being an organized viscous vortex; thus once the regular metion has been disrupted by sinuous or core instability the danger factor whll change. This illustrates the importance of the understanding and prediction of the catastroparc inctabilities.

The other critical moce, the ancounter normal to the trating vortices: is somewhat tese simple to handle. In prixiple it involves the
same factors. The disturbance system will be described in terms of the peak vertical velocities and their spacing, while the response system now relates to the longitudinal transfer function of the encounter aircraft. Confining the analysis to the encounter with only one core of the vortex pair, it has been shown by Houbolt (1971) that the maximum normal acceleration $N$ of the encounter aircraft can be written in a form involving the wing loadings and speeds of both aircraft, the aspect ratio of the generating aircraft and the tranofer function of the encounter aircraft. This transfer funcion ( $k_{t}$ ) is of course related to relative mass of tine encounter aircraft as well as the ratio of its chord to the vortex core. A curve for this is given by Houbolt, which shows fairly weak variation with relative mass and core radius. We write Joubolt's equation in the same terms as the Crow axial danger factor, and assume a relatively high value of $k_{g}$ of .85 (to round off numbers), then obtain a transverse danger factor $D_{T}$ as

$$
\mathfrak{N}_{\mathrm{T}}=20\left(U C_{\mathrm{L}} / \pi \mathrm{A}_{\mathrm{g}}, \quad \mathrm{VUC}_{\mathrm{L}}\right)_{\mathrm{e}}
$$

where $N$ is the limit normal acceleration factor of the encounter wircraft. On this basis, a value of $D_{T}$ exceeding unity may be regarded as haza ráous.

We now noie that the rativ of these terms wili sndicate the retative severity of axial or nommal encomater. Thus we obain

$$
D / D_{T}=b_{g}\left(N U C_{L} / 20 b^{3} P_{n a x}\right)_{c}
$$

 faincd, we cafo wite this as

$$
\mathrm{D} / \mathrm{D}_{\mathrm{T}}=0.35\left(\mathrm{NC} C_{\mathrm{L}}\right)_{\mathrm{e}} \mathrm{~b}_{\mathrm{g}} / \mathrm{b}
$$

Now, taking representative transport aircraft values of N. 4.5, $C_{L}-.6$ we see that the ratio of the danger factors is approximately equal to the span ratio, showing that in all significant cases (omalier encounter aireraft) $D>D_{T}$ so that the axial encounter can be expected to be more hazaridous.

## 4. VORTEX TRANSPORT

### 4.1 GENERNL

The trailing vortex pair creates an induced field which causes the vortex system to convect downwards. This vertical speed is relatively small compared to the flight speed. Making the classical assumptions of elliptical loading and inviscid rollup to a finite vortex pair, we obtain for the downwash velocity $U *$

$$
U^{*} / U=4 C_{L} / \pi^{3} A
$$

We note for a typical wing of aspect ratio 6 that the downward motion may be about $1 \%$ of the flight cruise speed, increasing to 3 or $4 \%$ at the landing state.

This can be compared with the same ratio before the sheet has rolled up, which is about $\left(C_{L} / \pi\right)\{1 / A+1 / \pi\}$ at the trailing edge, taking into account the bound vortex term, and about $2 C_{L} / \pi A$ a few spans downstream. Thus for the first few spans the downwash may be about four times that of the final state; however, our interest is with the motion after the vortex pair has developed. Field observations have indicated that apparently in all cases the vortex dee $2=$ nt rate is approximately constant for $30-40$ seconds, then begins to reduce. Condit and Tracy (1971) report that in no cases have descents of more than about 1000 ft been observed, while the recent experiments of Tombach (1972) show a distinct reduction in transport velocity. However, we stress that definitive measurements of vortex descent rate are lacking. We discuss this further in the vortex decay section. It appears clear that after about 60 seconds the descent rate is not adequately described by the classical equations.

If the pair is inmersed in a uniform crosswind, standard calculations indicate that the pair should simply be laterally convected with the crosswind, with no change in cell shape or vortex dynamics. This appears to be the case.

For a vertically non-uniform crosswind the situation is less simple. The pair is now immersed in a field containing a directed or ganized vorticity and one would expect variGus unsymmetrical effects such as vortex tilting to occur.

The simplest case, that of a uniformly sheared crosswind (linear wind p.ofilo), can be solved exactly in the inviscid case. This poss ses a steady solution in a frame of reference fixed in the vortex cores, and for the inviscid model gives no banking or change in trans ort speed. This analysis is developed in a later section. It is simply noted here that the reason this problem can be solved is that aithough the external vorticity is redistributed by the vortex fieid, this does not introduce ny unstendiness in the flow since the redistribution does not change the uniform vorticity field in which the cell is immersed.

If one consirt res a field having a shear gradient, then even the inviscid problem becomes unsteady because of the redistribution of the non-uniform ambient vorticity. Ail unsieady model was utilized br Burnham (1972) in attempting to model the ground effect case with a non-uniform crosswind. A finite set of discrete vortices was introduced into the ambient field and permitted to convect with the flow. This model is an atcempt to explain the observed rising of one of the vortices near the ground. It is believed that this model would give different quantitative vortex paths for different or more detailed models of the same crosswind field. We note too, that Bisgood, Maltby, and Dee (1971) modeled the vortex tiajectory in ground effect with a non-uniform crosswind simply by computing the convection at each height and taking no account
of vorticity redistribution, This simple kinematic model gave excellent corretation with the otserved vortex paths, at least for the few reported tests.

We note hers hat a large proportion of the fests conducted by Tombach exhibited unsymmetrical vortex response under crosswinds. This was normally observed in the vortex pair banking, ard occasionally in the collapse of one vortex due to core bursting. Such behavior might be expected under the ambient conditions described, where interaction between the ambient and core vorticity could cause stiong anti-symmetrie effects both with respect to vortax transport and core development. Changes in core development could accelerate the instability of one element of the pair.

One of the more important effects of nisymmetrical core bursting might be expected to be the development of the solitary vortex. In a number of tests reported by Tombach it was observed thet when only one vortex remained (or at least was visible) this vortex appeared to have a much longer ife than usual, and cid not show any signs of sinuous Crow Instability.

It appears that if these unsymmetrical effects are due to wind shear, then they shculd certainly be primarily attributable to the uniform shear case and that it should not be necessary to postulate a non-uniform shear to describe them. The basia for this is twofold, the first reason being that these effects have been observed at altituHes at which the uniform ahear is $w$ ank and the shear gradient even weaker; the second that any viscous model will indeed give antisymmetric behavior when immersed in a unform shear field.

Thus while an inviscid uniform shear model will not give any vortex tilting, it is belicved that the uniform shear is the dominant
effect and that a combined uniform shear/viscous decay model should be sufficient to account for yortex behavior in a general crosswind. This is described in Section 4.3

A further factor complicating the descent dynamics is the effect of flow stratification. The previous paragraphs have described the situation in a homogeneous fluid with the main effects due to entrainment or ambient organized vorticity. If the atmosphere is stratified as is frequently the case, one would certainly expect this to have some effect on the descent rate. At the moment, both the sense of the effect and its magnitude are controversial. Contrary to intuition, it appears that descen does not necessarily slow down in a stable stratification; that is, the cell is not necessarily retarded on penetrating the denser medium. Physically, this anomalous behavior is accounted for by the contraction of the vortex span -- the problem is whether the span dres in fact expand or contract and at what rate. In a later section we develop a new theory which shows that an inviscid, stratified model does :- iy a span contraction and consequent acceleration of the downward descent. More details and an evaluation of these as sumptions are made in the section.
ii shuuid dways be recalled that in practically all cases the organized vortex pair motion is terminated by sinuous instability or by core bursting. This is the element which makes experimental data on descent rate so incomplete.

In the following sections we assume thai an organized vortex pair motion exists and discuss in more detail three important cases: the descent in a homogeneous irrotational fluid, the descent in a homogeneous fluid of uniform rotation, and the descent in an irrotational fluid of uniform stratification.

Thus, if we consider the descent rate to be characterized by $U^{*}=U *(Z, I / T)$ where $Z$ is the ambient vorticity and $T$ the BruntVaisala time (a measure of the stratification) we attempt to describe the problems

$$
\begin{aligned}
& U^{*}=U^{*}(0,0) \\
& U^{*}=U^{*}(Z, 0) \\
& U^{*}=U^{*}(0,1 / T)
\end{aligned}
$$

Evidently a full analysis of the general problem cannot be achieved until the special cases above are understood and quantified. However, we note that a validation of the above cases may be difficult because it must depend on one of the atmospheric parameters approaching zero in the field test.

$$
4-5
$$

4.2 NEUTRAL IRROTATIONAL FIELD

After the wing vortex sheet has rolled up,the trailing $\mathrm{sff}_{\mathrm{f}} \mathrm{s}$ tem consists of a pair of vortices of finite rotational core area, but with the core radius a relatively small fraction of the vortex span. If this vortex pair is immersed in a still homogeneous inviscid flow, a well known classical solution exists under which the pair is convected downwards at a velocity $\Gamma_{0} / 2 \pi \mathrm{~b}$. The classical analysis shows that there is a closed recirculating mass of air, of roughly oval propor.. tions, associated with the concentrated vortex pair, and that this cell is convected downwards at a uniform speed. Flow exterior to this cell never enters it. On this basis a long vortex pair, which may be regarded as substantially two-dimensional, will move downwards in an unbounded fluid with constant velocity for all times.
in real flows this situation does not persist indefinitely, and most experiments appear to show the rate of descent to reduce and finally approach zero. This is certainly due to diffusion of core vorticity by some combination of laminar and turbulent viscosity and will occur even in homogeneous (unstratified) slows. Much effort has gone into explaining and quantifying this effect, but the subject still remains controversial. Recently a new rational interpretation of this, coupled with careful observations, has been put forward by Maxworthy (1972) which greatly assists explanation of the effect.

Maxworthy conduct experiments with vortex rings in water, using various visualization techniques to identify where the flow went. When the vorticity was relatively well distributed in the ring he observed that the outer flow was entrained into the back of the cell, causing an increase in the cell volume. At the same time a portion of the cell vorticity was shed into the wake, removing both vorticity and momentum from the cell. The consbined effect of this is to increase the cell size and to reduce its propagation velocity.

The mechanism of entrainment is important for our further development. Figure 4-1 shows a sketch of this flow field, in coordinates fixed at the core centers, so that the outer flow is represented by a uniform but unsteady flow from below. We note that the cell has a well defined stagnation point, $A$, and over the front portion, a well defined cell boundary, A-B. Across this boundary the pressure and velocity fields of inner and outer flows are continuous, the only discontinuity being between the inner vortical fluid and the outer irrotational flow. Due to both laminar and turbulent effects, this vorticity is transferred to the outer flow, and as a consequence the total head of the flow is reduced. Thus, after passing the maximum velocity point near B , the outer flow, contained approximately by the stream tube $C D$, is unable to recover sufficient velocity to rejoin the outer flow at the rear, but remains as part of the stationary cell. Thus the cell size is increased. At the same time, a neighboring stream tube EF acquires a smaller amount of vorticity and suffers less head reduction, so that is does depart from the cell at the rear, but at a lower than free stream total head. This portion develops into a wake behind the cell.

Thus the same process causes entrainment of the outer flow into the cell and a detrainment (removal) of some of the cell vorticity and momentum. A further process occurs on the centerline of the cell, AX. Here vorticity is annihilated by diffusion from the left and right cells. Ti, us three vorticity transfer mechanisms occur and i:ie overall effect controls the cell dynamics.

Maxworthy showed that initially the vortex shedding to form the wake was extremely weak, since the cell vorticity at the boundary was quite weak. Thus, although the cell grew in size, it did not lose momentum, and the impulse I was conserved. In these circumstances the main vorticityloss occurred along the centerline and was small, and there was minimal wake momentum loss. During the later stages in growth history, when more vorticity is present near the boundary, the


Figure 4-1. Viscous Effects on Vortex Cell.
wake develops; thus the impulse in the cell reduces while the cell size increases. Both these effects contribute to the reduction in speed and final complete annihilation of the cell momentum.

It must be noted that Maxworthy's experiments were conducted with vortex rings at extremely low Reynolds Numbers, when the flow was certainly laminar. Maxworthy has indicated in a private - munication that further flow visualization tests with finite wings a 100 exhibited a detrained wake. These experiments were also at very low Reynolds Numbers. It is, however, possible that during the later stages of development of an aircraft trailing vortex systeri: that similar processes of mass entrainment and momentum detrainment occur. For laminar transfer the time scales would be too long to be of interest, but if the transfer is assumed turbulent, it may be possible to account for some of the observed effects. Thus it appears very probable that the later development of a vortex pair followa qualitatively the stages described by Maxworthy; with an additional initial stage which we will discuss. We therefore postulate three stages as shown in Figure 4-2.

Stage I - The Inviscid Cell. Here the vorticity : © confined to well within the cell boundary. On the boundary itself there will be no laminar transfer (since there is no distortion) and turbulent transfer will have no net effect, since both inner and outer flow have the same total head. In these circumstances the inviscid cell model will be a good representation of the dynamics and we find that the time rates of change of cell size, $\dot{b}_{v}$, and impulse $I$ are zero; so that the propagation velocity $\mathrm{U}^{*}$ is constant.

Stage II - The Entraining Cell. As the core vorticity diffuses and approaches the cell boundary, the first process (of mass entrainment) occurs, represented by $\dot{b}_{v}>0, \dot{i}=0$. We find that the propagation velocity reduces from the inviacid value.


Figure 4-2. Vortex Cell Stages

Stage III - The Decaying Cell. During the later stages, substantial monentum and vorticity shedding occur, causing a wake to develop and giving $\dot{b}_{v}>0, \dot{i}<0$. Of course, various catastrophic instabilities may have developed before the complete decay has occurred.

## 1. Ordering of the Process

Maxworthy has made a jaminar analysis of the last two stages for a ring vortex, and found excellent correlation with his experiments. He also includes an analysis for the two dimensional vortex pair for laminar diffusion. Here we continue this for turbulent diffusion.

To establish a curbulent transfer coefficient, we note that there will presumably be two turbulent scales at the cell boundary. During the initial stage we assume that ambient furbulence is the major term, since the inner flow near the cell boundary has very weak voricity. If this is the case, the only turbulent parameter is , the :urbulent dissipation, and we can define an rms turbulent velocity $V_{\varepsilon}$ for all scales amaller than the cell scale by by

$$
v_{e}-e^{2 / 3} b_{v}^{1 / 2}
$$

In a later staige, we expect mechanically-generated turbulence due to the interior cell motion to cominate. Nor we will take a turbulent velocity coupled to the celi velocity, bo define another turbulent scale velocity

$$
v_{m}-v^{*}
$$

a. The Entraining Cell. Here we assume vorticity is entrained linearly with time as the layer progresses from A to B. Thus the height $\delta_{e}$ is given by

$$
\delta_{e}=V_{\epsilon} t^{t}
$$

where $t$ is the time of travel. An estimate of $t$ is

$$
t^{\prime}=\pi b_{v} /(2 U *)
$$

for the inviscid cell proportions. Now the unit density mass flow through $B D$ is $2 U_{s}$, and taking the cell area to be about $m b^{2}$ and assuming all this mass is entrained, we obtain

$$
\dot{b}_{V} \sim v_{\epsilon}, \quad b_{V} \sim \epsilon^{2 / 2} t^{3 / 2}
$$

Taking the total impulse to be that of the inviscid eftl, $I=\Gamma \mathrm{b}$. and the propagation sped $U *=\Gamma / 2 n b$. Thus for I invariant we obtain

$$
U_{*}^{*}=I f\left(2 \pi \varepsilon \varepsilon^{3}\right)
$$

We compare this with Aaxwormy's resuft for lamitar diffigion of kinematic viscosity. . . Where he zives

$$
v_{v}=\left\{\left.1 u\right|^{1 / 4} t^{2!}\right.
$$

$$
u=-\left(1^{3}, y^{2} 1 / 5 e^{-1 / 3}\right.
$$


 case the drparture frow the inviscte ceit resuitk. at represented by the expoments, are nuch more tistinct.
b. The Decaying Cell. In thio case we admit both entrai ient of mass and detrainment of momectum; thus the impulse is no lon f $_{\text {f }}$ conserved. We assume a turbulent velocity acale $\mathrm{V}_{\mathrm{T}}$ which will be defined later. Then, as before, we obtain the cell growth equation

$$
\dot{b}_{v}=a v_{T}
$$

where $a$ is some constant.

Now, to find the change in impulse, we estimate the reduction in circulation. We ignore vorticity annihilation on the interior cell centerline but attempt to compute the vorticity transferred to the wake by turbuience through the cell boundz: $y$. Dimensional arguments suggest that the cell vonitity gres like $\mathrm{T} / \mathrm{b}_{\mathrm{v}}{ }^{2}$ and the turbulent flux of vorticity through the boundary then becomes of the arder (r/b ${ }^{2}$ ) $b_{V} V_{T}$. Thus we obsain

$$
\dot{r}=\theta r \dot{v}_{\mathrm{V}} v_{\mathrm{T}}
$$

where $\beta$ is another constant. Combining these equations with $\mathrm{s}^{1}=\mathrm{y}$ gives the result

$$
\dot{\Gamma} \Gamma=v b_{v} / b_{v}
$$

independent of $V_{T}$. implying that $-b_{Y}{ }^{-V}$.
If we now assume $V_{i}$ - Uo we obinin the reauls

$$
\begin{aligned}
& U^{*}-t^{-1 / 4 f(2+y)} \\
& \text { 8- } f^{(1-y l i z i z)}
\end{aligned}
$$

We obscrve that for the impulse to decay, $v>1$; howwer, the reduction of propagation velocity will be leos than that in the case of assuming turbulence coupled to the ambient dissipation. We note also an inherent contradiction in that if the mechanical turbulence is couplod to the $U^{*}$, then eventually it will reduce so that ambient turbulence again dominates.

## 2. Comparison of the Assumptions

We note that in this simple model we can assume that the impulse is conserved or not, and that the turbulent seale is either the ambient turbulent velocity or the propagation velocity. Thus we actually ubtain four cases for similarity solutions. The ordering of the response is show in the following table.

TABLE 4 : ThE こRERING OF THE RESPONSE Be FOUR ASES OF SIMILARITY SOLHT GNS

| Turbulent Mechansism |  | Ambient | Mechanical |
| :---: | :---: | :---: | :---: |
| No Wake | ${ }^{1}$ | $t^{5 / 2}$ | $8^{1 / 3}$ |
|  | U* | $5^{-3}$ | $:^{-2 / 3}$ |
|  | ! | ${ }^{\text {e }}$ | $0^{\circ}$ |
| Wa: | 3 | 13 | .$\left.^{1 / 7} y^{2} 2\right)$ |
|  | \% | $e^{-3\left(x^{+1}\right) r^{2}}$ |  |
|  | 1 | , -1/y-1/2 | $s^{-(y-1) /(y+2)}$ |

We note that even on this simple basis, four different similarity solutions can be developed. It is possible that the stages with time progress from the ambient/no wake to mechanical turbulence/no wake case through the mechanical turbulence/wake case to the ambient/ turiulence/wake case. However, this assumption, although sounding physically plausible, gives a non-monotonic variation in the propagation speed decay rate, which appears surprising. For example, if $\gamma=1+\delta$, we would find the propagation speed varied like $t^{-3}, t^{-2 / 3}, t^{-(2+\delta) /(3+\delta)}$, $t^{-3(1+2 \delta) / 2}$ as the different stages occurred.

We note here that, in fact, transfer must be due to some combination of the ambient and mechanically generated turbulence. A similarity solution does exist for the case $V_{T}=\epsilon^{1 / 3} b_{V}^{1 / 3}+\gamma U^{*}$ but only for $y=-4 / 3$ which gives an increasing impulse! It is clear from the above equation that if the cell size increases with time and the velocity reduces (as appears to be the case for a descending vortex pair), then the ambient and mechanically generated turbulence cannot maintain the same ratio. This illustrates that similarity solutions can only be expected when one form of transfer dominates.

## 3. Dimensional Analysis

If we attempt simply to model the problem dimensionally, we find that there are three initial parameters, $b_{0}, U_{0}=\Gamma_{0} / 2 \pi b{ }_{0}$ and $\varepsilon$. A functional equation connecting $U$ and $t$ may be written as

$$
U * / U_{0}=f_{m}\left(\varepsilon^{1 / 3} b_{0}^{1 / 3} / U_{0}, t / b_{0} U_{0}^{1}\right)
$$

where we have used a mechanical time scale to normalize $t$ and retain the parameter $n=\epsilon^{1 / 3} b_{o}^{1 / 3} U_{o}^{-1}$. This would provide a collapsing of data where ambient turbulence was insignificant. However, where ambient turbulence dominated, one would use an ambient time scale and write

$$
U * / U_{0}=f_{a}\left(\varepsilon^{1 / 3} b_{0}^{1 / 3} / U_{0}, t / b_{0}^{2 / 3} \epsilon^{-1 / 3}\right)
$$



Thus formulation will collapse data in regions where the ambient turbulence is important. A representation of this is shown in Figure 4-3 with the ambient/mechanical ratio $n$ as a parameter.

If data collapse of this sort is observed [rom proper analysis of flight tests, then the exponents of $t$ described in the previous sections can be determined.

We note in passing that there is yet another time scale, that of viscous transfer. According to Maxworthy's experiments, this has the form $\left(b_{0}^{3} / U_{o} v^{1 / 2}\right.$. It appears that this scale is too large to be of concern to us; that is, the time involved for the problem to become a purely viscous one is beyond our range of interest.

### 4.3 NEUTRAL ROTATIONAL FIELD

## Statement of the Problem

Various explanations of the motion of vortices in a crosswind with shear have been given !Burnham, 1972; Harvey and Perry, 1971;. These analyses were concerned with the vortex transport near the ground plane, where, typically, both vertical shear and ver.. tical shear gradients are present when there is a crosswind. The objective of these works was to analytically model the observed unsymmefrical motion of the vortex pair as they descended into the wind shear region. Instead of soth vortices leveling off at the expected height above the ground of $i=b / 2$, the upwind vortex sometimes drops below this value, and the downwind vortex ievels off and tren begins to rise to a distance greater than $b / 2$. Thus, a tilting of the vortex pair results, superimposed on the lateral motion of the vortices caused by the uniform component of the crosswind, and the proximity of the ground piane.

Such analysis near the ground plane is unsteady both because of the redistribution of vorticity in the shear gradient of the external flow and the change in separation of the aircraft vortex pair due to the presence of the ground plane.

However, vortex tilting has been observed at sufficient heights above the ground (Tombach 1972), that ground effect need not be considered. Under these conditions, the problem can be treated in the classical sense of a descending vortex pair in steady flow. Also, at altitude, any severe vertical crosswind shear gra lients would not be expected to be present, and the shear can be represented as a uniform vorticity field, again a steady flow situation. The problem may now be examined in a steady coordinate systcm, fixed to the vortex pair. The
solution to this problem will yield the streamlines of the external flow as well as the recirculating flow within the vortex cells. Since the vortex tilting phenomenon has been observed at altitude, such an analysis is expected to give a good inviscid representation of the effects of the flow asymmetry produced by a sheared crosswind interacting with the vortex pair and should form the basis for a more detailed viscous solution.

## Analysis

A vortex pair with circulation $\pm \Gamma_{c}$ are located at $y=0$ and $x= \pm b / 2$ respectively. A vertical component of velocity $v=\Gamma_{o} / 2 \pi b$ is directed upwardfrom below the vortex pair. The wind shear is represented as a horizontal velocity, $u=f(y)$ with $u=f(0)=0$ to remove the effects of the contribution of the uniform component of the crosswind (for the steady coordinate system chosen).

Assuming a linear crosswind shear profile, $u=f(y)=K y$, from which $f(0)=0$. This satisfies the condition of the removal of the uniw form component of the crosswind. The details of the combined flows described abrve are presented in Figure 4-4.

The stream functions for each of the flows may easily be written separately as:

$$
\Psi_{1}=-\frac{\Gamma_{0}}{4 \pi} \ln \left[\frac{(x-b / 2)^{3}+y^{3}}{(x+b / 2)^{3}+y^{3}}\right]
$$



Figure 4-4. Coordinate System and Flow Geometry for Vortex-Pair in Uniform Cross-Flow.

## uniform vertical flow:

$$
\Psi_{2}=-\frac{\Gamma_{0}}{2 \pi b} x
$$

and
shear flow:

$$
\Psi_{3}=\frac{K}{2} y^{3}
$$

from which the stream function $\psi$ for the complete flow field may be obtained.

By superposition:

$$
\Psi=\Psi_{1}+\Psi_{a}+\Psi_{3},
$$

and

$$
\begin{equation*}
\Psi=\frac{K}{2} y^{a}-\frac{\Gamma_{0}^{<}}{2 \pi b} x-\frac{\Gamma_{0}}{4 \pi} \ln \left[\frac{(x-b / 2)^{2}+y^{a}}{(x+b / 2)^{2}+y^{2}}\right] \tag{4-1}
\end{equation*}
$$

Introducing the following dimensionless parameters,

$$
\begin{aligned}
X & =\frac{x}{b / 2}, \\
Y & =\frac{y}{b / 2}, \\
\Psi * & =\frac{4 \pi}{\Gamma_{0}} \Psi, \\
\sigma & =\frac{K \pi b^{a}}{2 \Gamma_{0}},
\end{aligned}
$$

Eqn. 4-1 becomes:

$$
\begin{equation*}
\Psi *=\sigma y^{2}-x-\ln \left[\frac{(x-1)^{3}+y^{3}}{(x+1)^{3}+y^{2}}\right] \text {. } \tag{4-2}
\end{equation*}
$$

Here, $\sigma$ is a non-dimensional shear parameter which relates the scale of the wind shear to the vortex parameters, $\Gamma_{0}$ and $b$. The vortices are located at $Y=0, X= \pm 1$.

Before proceeding further it would be is eful to calculate the order of magnitude of the shear parameter, $\sigma$, for subsequent evaluation of the streamlines, $\psi^{*}$, from Eqn. : 2 .

For a 727 aircraft,

$$
\begin{aligned}
& \Gamma_{0}=34\left(0 \mathrm{ft}^{2} / \mathrm{sec}\right. \\
& \mathrm{b}_{0}=108 \mathrm{ft} \\
& \mathrm{~b}=(\pi / 4) \mathrm{b}_{0}=\pi(108) / 4=84.78 \mathrm{ft} \\
& \sigma=\frac{\mathrm{K}_{\pi} \mathrm{b}^{3}}{2 \Gamma_{0}}=3.31 \mathrm{~K}
\end{aligned}
$$

Typical values of shear near the ground (Zwieback, 1964) give a velocity of $18 \mathrm{ft} / \mathrm{sec}$ at 44 ft above erornd, so that $\mathrm{K}=\mathrm{u} / \mathrm{y}=18 / 44=.4 \mathrm{sec}^{-1}$. Stronger shears a: possible closer to the ground or in gusty situations, so that it is not unreasonable to assume a range of $0 \leq K \leq 1$. This gives an upper limit on $\sigma$ of $\sigma=3.31$. Eqn. 4-2 will be evaluated over f range of $\sigma$ rom zero to 3.

## Computation of

Equation 4-2 gives the coordinates $X, Y$ for the family of streamines $\psi_{n}^{* *}$ at constant $\sigma$. Although the equation is in closed form, it is not easy to solve for $X$ and $Y$ for any given $\Psi_{n}^{*}$, especially when the value of $\Psi_{s}^{*}$, the stagnation point streamline, which also describes the bounding streamline of the recirculating vortex cells, is not known apriori.

The expression for $\psi^{*}$ was programmed on a Hewlett-Pack .rd Model 9820A mini-computer using an $x-y$ plotter for data output. After some experimentation with various inputs of and $\sigma$, the pattern of the structure of the streamlines became clear. Figures $4^{\circ}-5,-6,-7$, and -8 show the geometry of the stagnation, or dividing streamline of the vortex recirculation cells as well as several streamlines of the exterior flow for values of $\sigma=0,1,2,3$ respectively. In addition, Figure 4-9 is a "magnified" view of the streamlines in the immediate vicinity of the stagnation point for the case $\sigma=3$.

## Discussion of Fesults

The obvious (and very striking) conclusion from Figures 4-5 through 4-8 is that as the strength of the wind shear increases, the size of the upwind vortex recirculation cell increases and the size of the downwind cell decreases. Presumably, for o large enough (a not very realistic case), the downwind cell would approach zero.

The problem treated here was assumed steady in all respects, including constant $\sigma$. That is, the vortex pair was generated in a crosswind shear field of constant o which remains constant as the vortex pair translates downwind. It is not clear from the present analysis what the


4.2:

 the shext Bitametor, $=2.0$. Crosswind from the Left.

Figure 4-8. Streamlines in the Vicinity of a Vortex Pair in Feavy Shear. Value of the Shear Parameter, $\sigma=3.0$. Crosswind from the Left.


Figure 4-9. Streamlines in the Immediate Vicinity of the Top Stagnation Point for a Vortex Pair in Heavy Shear. Value of the Shear Parameter, $\sigma=3.0$.
effects of changing $\sigma$ would be on the geometry of the flow, once established. However, it might be reasonable to speculate that Figures 4-5 through 4-8 give an instantaneous, steady view of the flow for each value of $\sigma$, and if $\sigma$ changes slowly, the flow field would change roughly in accordance with these figures. Thus, as $\sigma$ increases, the upwind cell would gradually increase in size and the downwind cell would gradually shrink.

Further analysis must be performed to verify this, but if the above assumption is reasonable, a method of predicting the flow geometry of a vortex pair descending into a crossflow with shear gradients would be possible.

It is not clear from the present analysis of the exact effects that a sheared crosswind has on vortex tilting. However, it is unquestionably the case that the wind shear produces asymmetry in the recirculating cells surrounding the two vortices. This asymmetry may explain the process whereby vortex tilting occurs. It should be remembered that the asymmetry always acts to increase the size of the upwind vortex recirculation cell and to reduce the size of the downwind cell, and that this effect is more pronounced as the strength of the wind shear is increased.

Section 4-2 describes the outward diffusion of the vortical core as the age of the vortex pair increases. The core diffusion is presumably related to vorticity gradients. Due to wind shear, the downwind vortex recirculation cell will be smaller than the upwind cell. The diffusing vorticity of the downwind vortex will thus reach the boundary of the inviscid cell before the similar situation occurs for the upwind vortex. Thus, vorticity associated with the downwind vortex will diffuse into the free stream flow and be swept away, reducing the vorticity and hence the circulation of that vortex. An asymmetry will then be produced
in the downwash velocities of one vortex with respect to the other.

Although the effect of these asymmerical induction velocities should take into account the exact distribuno of vorticity within the altered vortical cores, the sense in which the vorticen translate with respect to each other should be straightforward.

If the circulation of the downwind vortex is decreased over that of the upwind vortex due to detrainment $u$ vorticity, then its downwash velocity will be less than that of the upwind vortex. As a result, the upwind vortex will tanslate downward less rapidly than the downwind vortex. Tilting will occur with the sense that that the upwind vortex will rise with respect to the downwind vortex.

This result contradicts Burnham's observation that "the upvind vortex ueually drops to a lower altitude than the downwind vortex." However, Burnham's work, and the experimental observations which he cites were all performed in ground affect at relatively low crosswind velocities. He states, "Both experimental observations and calculations indicate that other phenomena occur at much higher crosswind speeds (e.g., 30 fps instead of the 10 fps considered here)."

Firture work needs to be performed to adequately describe the mes:anism which initiates vortex tilting. It seems reasonable from the p.esent analysis that the asymmerry in the flow fields surrounding the vortex cores forms an excellent point of departure for such further investigation. The assumption of aniform shear was sufficient to incicate this asymmetrical behavior. Including the effects of shear gradients in future work would undoubtedly refine the results somewhat, but their effects should probahly be small and would produce results in the same senue indicated by the present analysis.

For a sufficiently large vortex age, the core diffusion process would cortinue to a point $2 t$ which vorticity detrainment srom the downwind vortex (with shear flow) would result in sufficient reduction of its circulation that for all practicul pur ooses, it would cease to exist. This would not be the case for the upwind vortex, considering its larger cell size. This process would proceed more rapidly for larger $\sigma$. Indeed, for a locally large value of $\sigma$ such as might exist in a gust with high shear rate, the above process would occur caiastrophically -- a pussible expianation for core bursting. For both the case of the slower detrainment rate and the catast: ophic situat.on, the remaining vortex would be the upwind vortex.

## Concluding Remarks

The calculation of the family of strearilines $\psi_{n}^{2 / 2}$ for a given value of shear $\sigma$ is tedious. The only really necessary streamline to obtain is that which describes the boundary of the recirculation cells, $\Psi_{b}^{*}$.
 at the stagnation point $\Psi_{\delta}^{*}$. $\Psi_{s}^{*}$ may be solved in terms of $\sigma$ explicitiy and presented graphically or in tabular form. Substitution of $\psi_{s}^{*}=\Psi^{*}(\sigma)$ in Eqn. (4-2) will then give the shape of the recirculation cells for any desired value of $\sigma$.

$$
\begin{align*}
& \text { The conditions for the calc alation of } \Psi_{s}^{\mu} \text { are: } \\
& \underset{\substack{*}}{\underset{\sim}{*}}-f_{1}\left(X_{s}, Y_{s}, \sigma\right)=0 \quad,  \tag{4-3}\\
& u=\left.\frac{\partial \Psi *}{\partial Y}\right|_{s}=f_{a}\left(X_{s}, Y_{s}, \sigma\right)=0,  \tag{4-4}\\
& \mathrm{v}=-\left.\frac{\partial \psi *}{\partial \mathrm{X}}\right|_{\mathbf{s}}=f_{3}\left(\mathrm{X}_{\mathbf{s}}, \mathbf{Y}_{\mathbf{s}}\right)=0, \tag{4-5}
\end{align*}
$$

where the subscript $s$ implies that the quantities are to be evaluated at the otagnation point. Equation (4-3) is identical to Egn. (IV-2) evaluated at $s$.

The simultaneous solution of Eqns. (4-3, 4, 5) will yield the values of the three unknowns, $X_{8}, Y_{s}, Y_{s}^{*}$ in terms of $\sigma$, of which only $\Psi_{s}^{*}$ is of interest. The simultaneous solution of these equations is also not exactly easy, but can be performed in a straightforward manner using computer techniques.

The cell boundary streamline will be given by the expression:

$$
\begin{equation*}
-\Psi *(\sigma)+\sigma Y^{2}-x-\ln \left[\frac{(x-1)^{2}+y^{2}}{(x+1)^{2}+y^{2}}\right]=0 \tag{4-6}
\end{equation*}
$$

from which the shape of the recircuicting cells may be calculated or plotted by computer. The shape, size, or area of the recirculating cells may thus be obtained as continuous functions of the wind shear, $\sigma$, within the accuracy of the computarional methods used.

### 4.4 WAKE TRANSPORT IN A STRATIFIED FIELD

If a rortex wake descends through a stably stratified atmosphere, as is usually the case, then it becomes buoyant as it descends because its density increases at a slower (adiabatic) rate than the ambient density. A wake could also conceivably acquire buoyancy at the time of its generation, from hot engine exhaust gases mixed into it. In either case, the buoyancy thus generated acts to decrease the overall wake impulse and (possibly) circulation, although any mixing between the wake and its surroundings would tend to dilute the buoyancy and to decrease these effects.

### 4.4.1 EXPERIMENTAL OBSERVATIONS

Observations of wakes generated by full-size aircraft tend to suggest that stratification may have an effect on wake transport. Specifically, a study of the dissemination of particles released from aircraft (Smith and Wolf, 1963; Smith and MacCready, 1963) included some observations of wake descent from a variety of aircraft. Tests with aircraft up to DC-3 size gave typical descent distances of 25 to 100 feet over land, with the wake descending initially at the theoretical speed, but then broadening and slowing up. On the other hand, similar measurements over the ocean sometimes, but not always, indicated wake descents of 600 feet. A possible explanation was that turbulence slowed the wake's motion and helped spread it and that a stable atmosphere further impeded descent. Both stability and turbulence are usually less over the ocean than over land, thus resulting in longer vortex life and descent.

Another study (FAA Task Force, 1971) found that the vortices generated by jumbo jet aircraft began descending at 400 to 500 feet per minute, but that fully developed wakes were generally not found more than 1000 feet below the altitude of the generating aircraft.

Leveling off, combined with start of breakup usually began after a descent of 800 to 900 feet and the vortex spacing appeared to remain constant until breakup. Occasional descents of 1100-1200 feet, with no noticeable slowing of descent speed, were measured during the same study by Andrews, et al (1972). In contrast, in one case the wake from a cruising $\mathrm{C}-5 \mathrm{~A}$ (Mach no. $=0.8$ ) was measured at 500 feet below the dircraft at 3 miles behind it, and remained at that level for more than 40 miles.

During a British study of wakes behind jet transport aircraft (Rose and Dee, 1965) wake descents of up to 800 feet were meassured during the first 150 seconds of wake life and no general leveling-off trend was observed. Quantitative measurements of vortex spacing were obtained in another British study (Bisgood, et al, 1971) where the vortex spacing behind a delta-wing aircraft was observed to grow to about 3 times the initial spacing in about 45 seconds.

The descent of the wake generated by a lightplane in a variety of meteorological conditions was measured by Tombach (1972, 1973). He found that the descent speed of the wake was constant for about 20 seconds in a stable atmosphere, after which the descent was slowed and often stopped. There seemed to be no noticeable differences in wake behavior at different atmospheric stratifications, however, but the dominant factor governing wake descent was probably a banking of the plane of the vortices (which was discussed earlier).

Tombach also attempted to measure vortex spacing as the wake descended. The same banking tendency complicated the interpretation of the data, but he noted that slight increases in vortex spacing could be observed in the few cases when the wake remained relatively level. Considerable increases in spacing were noted for wakes which had banked well out of the horizontal plane.

The observations by Tombach showed that the smokemarked vortex wake sometimes rose again after its descent had stopped. Whether the cause for this is atmospheric upcurrents or buoyancy is not known.

A few laboratory experiments have been performed to study the motion of buoyant vortex pairs. Most of these have been concerned with line thermals, where the buoyancy of the thermal is such as to increase the total circulation around each vortex. When in a homogeneous medium and well developed, i.e., when the vortex motion is well organized, these thermals are observed (Richards, 1963; T'sang, 1971) to grow in scale as $\Delta \rho^{1 / 3} t^{2 / 3}$, where $t$ is time (with $t=0$ at the extrapolated point when the vortex pair would have been of zero size) and $\Lambda \rho$ is the density difference between the thermal and the ambient fluid at the time of release. Their speed of advance has been found to slow as $\wedge \rho^{1 / 3} \mathrm{t}^{-1 / 3}$. This entrainment of the ambient fluid is observed to erode the buoyant driving force and to eventually bring the growing thermal to a halt. Extension of such results to the vortex wake problem is not straightforward, however, because (1) the buoyancy opposes the motion of the vortex wake, and (2) the impulse of the thermal is coupled to its initial buoyancy, whereas the impulse of the vortex wake is totally unrelated to its initial buoyancy (which is often zero).

One set of laboratory experiments to study the motion of a vortex pair with initial impulse in a stratified medium was performed by Tulin and Shwartz (1971) (with a related experiment by Birkhead, et al, 1969). They generated an impulsive vortex pair in a stratified water tank and observed its motion, which looked like that shown in Figure 4-10. (The figure has been redrawn to correspond to the direction of motion of an aircraft wake). The total vertical travel of the wake was about 4 vortex spacings (which is of comparable magnitude to the aircraft results quoted earlier) before the stratification dominated


Figure 4-10. Trajectory of a Vortex Pair in a Linearly Density Stratified Medium (from Tulin and Shwartr, 1971). The Upper Sketches Show the Observed Cross Section of the Wake.
the motion and caused the wake to reverse direction and collapse. This process took place in a characteristic time which is proportional to the Vaisala period $T_{1}=\left(g / \rho d_{\rho} / d y-g^{2} / a^{2}\right)^{1 / 2}$ where $\rho$ is the atmospheric density and $a$ is the speed of sound, with the proportionality factor depending on the initial strength of the vortices, which in turn is represented by another time $T_{2}=b_{v} / V$ where $V$ is the descent speed and $\mathrm{b}_{\mathrm{v}}$ is the vortex spacing. According to Tulin and Shwartz, if $\mathrm{T}_{2}<\mathrm{T}_{1}$, the vortices dominate the motion, while if $T_{2}>T_{1}$ the stratification dominates it. Since $T_{2}$ increases if the wake descends and grows, their experiments show that stratification must invariably dominate the motion unless some instability has first destroyed the vortices.

The Tulin and Shwartz data appear to be of good quality, but some reservations exist about their applicability to the aircraft wake problem. The method of vortex generation, with an impulsive motion of a piston pushing fluid through a slit, does not generate a pure vortex pair, but adds a wake behind the moving vortex pair. (This problem is a well known one in vortex ring generation also). Since the total vortex wake motion from the wall orifice is only a few times the wake size, the effects of the wall and of the wake trailing behind the vortices might be aignificant. It should be noted also, that the stratification was quite strong, with the Vaisala period $T$ ranging from $1-4$ seconds (as compared to typical atmospheric values around 100 sec ). However, the ratio $\mathrm{T}_{2} / \mathrm{T}$ is initially of order $1 / 10$ for both their experiments and for full scale wakes in the atmosphere, which indicates comparable scaling of vortex strength versus buoyancy.

In summary, the expesimental picture is somewhat fuzzy. The evidence to show that wake descent atops in a stratified atmosphere is not conclusive, but the suggestion ia there. Similarly. vortex spacing seems to vary little with descent, or possibly to increase, but again the experinmental evidance is not extromely strong. Using the
knowledge of wake behavior, limited and confused as it may be as a foundation, let us now look at efforts to approach the problem analytically.

### 4.4.2 PREVIOUS THEORETICAL MODELS

The theories which have been developed to describe the motion of a vortex wake in a stably stratified atmosphere are numerous, include various types of modeling approaches, and give a variety oi contradictory conclusions. The state-of-the-art is summarized below in Table 4-2. As shown in the table, the models suggest two main types of behavior -- a slowing of wake descent with an increase of vortex spacing, or an acceleration of wake descent accompanied by a decrease in vortex spacing.

$$
\begin{array}{ll}
\text { TABLE 4-2. } & \text { COMPARISON OF THEORETICAL MODELS } \\
& \text { FOR HESCENT OF A VCRTEX WARE } \\
& \text { IN A STARLY STRATIFIED ATMOSPHERE. }
\end{array}
$$

| Vortex | Descent | Buoyancy- <br> generated |
| :---: | :---: | :---: |
| Author | Spacing | Speed |

Costen (1972)**

Weak stability
Strong stability
Kuhn \& Nielsen (1972)
Saffman (1972)
Scorer \& Davenport (1970)
Tombach (1971)

| Strong stability | Increases* | Stops ${ }^{\text {a }}$ | Entraired |
| :---: | :---: | :---: | :---: |
| Weak tiability | Decteases | Increzses | Entrained |
| ulin be Shwarta (1971) | Increases | Stops | Entrained \& annihilased |

- See discussion in text.
* Costen's modei is a buryant core model. All others are models for a buoyznt wake oval.

| Decreases Increabes (?)* | Increases Stops (2) | No effect |
| :---: | :---: | :---: |
| Decreases | increases | Partly entrained |
| increases | Stopa | Entrained |
| Decreasea | increases | Detrained |
| Increases* | Stops* | Entraised |
| Decreases | Increzsen | Entrained |
| Increases | Stops | Entrained \& annihilated |

In order to assess the credibility of the various models each of them will be discussed briefly below. Since there has been a slight evolutionary trend, they will be presented in a generally chronological order.

The first two modeling approaches (Scorer and Davenport, 1970; Tombach, 1971) are both based on the assumption that the effect of buoyancy is to change the classical hydrodynamic impulse of the vortex pair and its accompanying fluid. Both assume that the buoyancy effects take place rather slowly, so that the instantaneous shape of the oval cylinder of accompanying fluid is the same as that of the classical wake in a uniform medium, and hence the wake size scales with the vortex spacing. The basic equations of motion are thus

$$
\frac{d I}{d t}=\rho_{0} F
$$

and

$$
\frac{d F}{d t}=-A v / T^{2}
$$

where $I \approx 2 \rho_{0} \Gamma b$ is the impulse per unit length of wake, $o_{0}$ is the initial density, $F=g A \Delta_{\rho} /_{O_{0}}$ gives the buoyancy, $A$ is the crosssectional area of the wake oval, the Vaisala perind $T$ gives the stability, and $v$ is the descent speed of the wake.

Scorer and Davenport assume the circulation to be constant. They fird a single solution in which the vortices converge and accelerate downward in a stable atmosphere. They compute numerically the internal streamline patterns for this case and suggest that the detrainment is indeed a stable process and that the circulation is, in fact, constant because the vorticity generated by the buoyancy at the boundary of the wake oval is continually
detrained as a curtain 3 bove the descending oval. After a time, however, they staie that some of this fluid and vorticity is mixed into the wake, which results in its eventual restabilization and destruction.

The idea of defrainment has some experimental support. Photographs of contrails riade from the side (Smith and Beesmer. 1959) show a curtain of condensed vapor extending above the wake up to the flight level of the generating aircraft. These same contrail observations, as well as low altitude measurements by Smith and MacCready (1963), suggest however that the vortex separation increases at least slightiy as the wake descends instead of decreasing as the theory predicts. This troublesome point will reappear again in luter analyses, so it should be pointed out here that buoyancy may indeed cause the vorticas to converge and accelerate even though guch behavior has not been observed experimentally. Since none of the analyees consider either turbulence or viscosity, both of which are ommipresent in the atmospherc. it is quite possible that other mechanisms ovefuhelm the buoyancy ef fect and cause the observed behavior.

Safiman (1972) points out shat the Seorer and Daveaport analysis does not properly treat the dynamial effect of the vorticity generated at the interface between the accompanyins fuid and the Getiser surronnd-
 calculation of the impulse of the wake. The erpor in neglectins it misy
 the ceaterifine by interaction with its cetanterpart of opposite gien frem
 frained vorticity.

 of its conclusions. Specificaiky, the rifot sides of their equariont (sj

$$
t-41
$$

and (11) should both be multiplied by (-1). The quantity $n$, which they call the vorticity, is really the vortex sheet strength, i. e., the circulation per unit length of sheet (which is the vorticity integrated across the sheet thickness). Their equations (17) for the "non-dimensional" vorticity" $\infty$ do not give a dimensionless value. The corrected expressions should be

$$
\frac{\partial \varphi}{\partial \xi}=\frac{k}{4 \pi R} \cdot \frac{w}{u q} \quad, \frac{\partial \varphi}{\partial \zeta}=\frac{k}{4 \pi R} \frac{1}{q}
$$

If these equations are used, then the expression in Ean. (16) is correct.

In contrast to the constant circulation postulated by Scorer and Davenport, Tombach (1971) considered the possibility that buoyancygenerated vorticity would decrease the overall circulation around each vortex, by allowing their circulation to vary at a rate proportional to the ambient density gradient. The approach used followed that of Turner (1960) for rising buoyant vortex pairs. In addition to the geometric similarity assumption made by Scorer and Davenport, Tombach also assumed that the overall entrainment and mixing process could be characterized by a single parameter which is represented as a length in the Bjerkness circulation equation, with the rate of change of the circulation of a vortex being proportional to this length. He also assumed that this length scales with all the other lengths in the wake, which then results in the rate of change of circulation being related to the wnse scale and the atmospheric stability, and suppresses any explicit display of the effect of the wake buoyancy. Although a similar assumption was used by Turner (1960) and is supported by experiments by Woodward (1959) with thermals, it is not a valid one to make initially when the vortex wake has no buoyancy and thus is not generating any vorticity. Thus the Tombach model does not properly describe either the initial behavior of the wake nor can it properly treat it at very Jong times when the oval height could be sufficiently great so that gravitational effects on its shape have to be considered.

In the intermediate time intervals, the model suggests that two distinct types of wake behavior may be possible, with the choice of which behavior will occur being determined by the initial vortex strength, the atmospheric stability, and the degree and manner of entrainment (which at present cannot be quantified). If the initial vortex strength is weak enough and/or the stability great enough, the model says buoyancy rapidly erodss the circulation, and the residual impulse can only be accommodated by an increase in the vortex separation and a slowing down of the wake descent. The motion is much like that of a vortex pair descending into ground effect.

On the other hand, if the vortex is strong or the stability weak, then there is residual circulation remaining when the impulse has been eliminated by the buoyancy, and the kinematical consequence is a rapid convergence of the vortices with an accompanying downward acceleration, similar to that predicted by Scorer and Davenport. The index determining which behavior will occur, denoted by $Q$, is given by

$$
Q \sim T \Gamma_{0} / b_{0}^{2}
$$

where $T$ is the Vaisala period, $\Gamma_{0}$ is the initial circulation of one vortex, and $b_{0}$ is the initial vortex spacing. The proportionality factor depends on the entrainment details. Small $Q(<\pi / 2)$ corresponds to the first type of motion mentioned, while large $Q\left(>_{\pi} / 2\right)$ corresponds to the second type.

As one would expect, the time scale of the motion is found to be proportional to $T$, hence the greater the stability the more rapidly the motions described above take place. The vertical length scale is proportional to $T^{2 / 3}$, thus as the atmospheric stability is increased the vertical extent of the motion will decrease.

In the same way as for the Scorer and Davenport model, Saffman states that the Tombach model is also in error because it equates the change in the wake impulse to the buoyancy force, which is correct only if the Boussinesq approximation holds (i.e., if the density difference between the wake and the environment are smodi) and if all the vorticity generated by the buoyancy is included in the computation of the impulse. He claims that the latter requirement was not satisfied in the analysis. However, the Tombach analysis does te'e into account the change in impulse due to the buoyancy generated vorticity according to the Bjerkness equation. It does assume, however, that it is possible to approximate the impalse of the wake, including the new voricity, by the impulse of a geometricaily identical wake with all of the vorticity, old and new, concentrated in the two vortex cores. This approximation will probably result in errors after the wake has acquired considerable buoyancy, although by that time the model may no longer be valid for a variety of other reasons.

Although qualitatively similar, the Tombach large $-Q$ solution carnot be matched to the Score and Davenport solution. Scorer and Davenport assume constant circulation, but permit detrainment of mass from the wake oval. On the other hand, Tombach requires that any mais transfer must affect the circulation, hence the assumption of constant circulation (which requires that his entrainment parameter $s=0$ ) would automatically require no entiainment or detrainment.

Costen (1972) points out that the portion of the Tombach model for which the circulation decreases more rapidly then the impulse is impossible. This conclusion is derived by taking the ratio of Tombach's equations (3) and (9), which give (in his notation)

$$
\frac{d \log }{d \log } \frac{\Gamma}{M}=\frac{1}{\Gamma} \frac{d \Gamma}{d t} \cdot\left(\frac{1}{M} \frac{d M}{d t}\right)^{-1}=\frac{2 h}{q} \cdot \frac{\rho-\rho^{\prime \prime}}{\rho-\rho^{\prime}}
$$

where $h b$ is the height of the wake oval ( $b$ is the vortex semi-span for his analysis), the wake cross-sectional area is $q b^{2}, \rho$ is the ambient density, $\rho^{\prime}$ is the zuerage wake density, and $\rho^{\prime \prime}$ is the average wake density along the axis of symmetry between the vortices. Numerical values of the constants are $h=3.46$ and $q \approx 11.62$, giving

$$
\frac{d \log \Gamma}{d \operatorname{lng} M}=.6 \frac{\rho-\rho^{\prime \prime}}{\rho-\rho^{\prime}}
$$

and since one would expect $\rho \geq \rho^{\prime \prime} \geq \rho^{\prime}$, this gives

$$
\frac{d \log \Gamma}{d \log M} \leq .6
$$

Thus the only self-consistent solutions of Tombach's equations are those for which $Q<\pi / 2$, i. e., the converging and accelerating solutions.

Another theory was developed by Tulin and Shwartz (1971) in which they modeled the vortex system by separate velocity scaling of the flows internal and external to the wake oval, with a shear layer at the boundary between the two flow fields. Based on experiments they performed in a water tank, they conclude that the wake entrains the vorticity generated at this shear layer (in contrast to the detrainment postu ited by Scorer and Davenport) but that the ingested vorticity is mainly canceled out at the wake centerline through mixing with vorticity of the opposite sign from the other side of the wake. Consequently, they model a turbulent wake in a stratified medium by assuming conservation of volume, mass, and energy and neglecting vorticity and momenturn. They then find completely similar solutions which depend on four parameters - the initial buoyancy of the wake, the stability of the fluid, the dissipation of kinetic energy, and the ratio of kinetic to potential energy. The latter two parameters are assigned values based on their experiments, which were dircussed earlier.

Tulin and Shwartz give formulas for vortex wake motion in a stratified medium, which are too complex to present here, and present excellent corielations between their formulas and experiments. Their model predi=*s that a wake will slow down and stop its descent in a stratified fluid and that the vortex separation will increase as the wake descends. The descent in a homogeneous fluid has the behavior $z \sim t^{1 / 2}$ and $b \sim t^{1 / 2}$. In a stably stratified fluid, the descent slows down more quickly and the spreading is more rapid, but similarity no longer holds when the wake has come to rest. The nature of the motions they observed experimentally and their model endeavors to describe was shown earliex in Figure 4-10.

In the Tulin and Shwartz formulation, as well as in those discussed above, the time scale of the motion is proportional to the Vaisala period ' f . For those models which predict a stopping of wake descent, this then means that the time required for the wake to descend to its lowest level is proportional to $T$.

Saffman (1972) has presented a model which indicates that a vortex system in a stable environment descends with the vortex spacing remaining constant if there is no entrainment, and increasing if entrainment is assumed proportional to the density difference between the wake and environment. in either case, the wake descent is stopped by the stability and it subsequently oscillates verii:ally. To find these solutions, Saffman solves the Laplace equation for the potentials inside and outside the wake volume, with th. constraint that there is no change in wake volume (except due to entrainment, which he treats separately). His criticisms of the Scorer and Davenport, and Tombach, models have already been reviewed, and his formulation was designed to overcome the deficiencies he described.

Saffman points out that it is "by no means certain" that turbulent mixing, caused by the Kelvin-Helmholtz instability of the wake-atmospliere


#### Abstract

interface, must occur when the motion is generated aarodynamically and buoyancy effects are (at least initially) a small perturbation. As was pointed cut carlier, this is a different situation from the line thermal case, where the motion is generatad by buoyancy. Since the experimental evidence obtained to date does allow not allow one to assert with any confidence that significant turbulent mixing does take place, he considered both non-entraining and entraining cases for completeness.


Work by Kuhn and Nielsen (1972) attempts to model the entrainment differently than Tombach or Scorer and Davenport. One shortcorning of the Tombach model was the assumption that the rate of change of circulation was proportional to the scale of the wake. This is erroneous at small times if the wake has no initial buoyancy, since then one should have no generation of vorticity. Kuhn and Nielsen use a variable entrainment length, proportional to the density difference between the entrained fluid and the ambient fluid, with the entrained fluid density being a weighted average of the ambient density, the density in the vicinity of the vortex, and the density in the region into which the outer fluid is entrained. The effect is to make Tombach's entrainment parameter $s$ into a variable which has value zero initially.

Kuhn and Nielson thus model the entrainment so that part of the buoyancy generated can be entrained and part of it detrained, with the proportions of each being governed by an unknown wake mixing parameter. Their analysis shows the wake accelerates as it descends and the vortex spacing decreases. Increased values of the wake mixing parameter reduce the rates of descent and convergence, while the addition of heated air to the wake, but outside the vortices, is found to cause a leveling-off and spreading of the wake. Saffman's criticisms about the proper formulation of the impulse to include the entrained vorticity apply here also.

Yet another model by Costen (1972) finds that buoyant vortex cores (different from the buoyant wake considered in the other models) will accelerate and close together as they descend in a neutral atmosphere. He concludes that stable atmospheric stratification will accelerate the effect. Costen also sliggests that a strong inversion layer might retard the vortex motion and cause their spacing to increase, but does not present any support for this inference.

In summary, the analytical situation with respect to the effects of stratification is not clearer than the experimental one. In fact, if one arrives at a consensus based solely on the number of models which predict the same behavior, one would conclude that buoyancy causes the wake to accelerate downward and the vortex spacing to decrease. This appears to be in contradiction to the experimental observations, which suggest the oppusite behavior. Both the theories and the experiments may be correct however, with the differences being due to factors such as entrainment (due to, say, ambient turbulence) which were not considered in the models. In fact, all of the models which predict the slowing down and spreading out behavior do include entrainment as a factor.

There are fundamental questions which arise about the formulations of these models, however, some of which were raised during the discussion. All of the models contain simplifying assumptions about the distribution of vorticity and of buoyancy, in order to make a tractable problem. In many cases the effect of such assumptions is not really known, nor is there empirical evidence to support or refute their validity.

Aa part of the current study, $A V$ has developed a more rigorous description of the effects of buoyancy on a wake descending through
an inviscid, non-tu.bulent, incompressible, stratified fluid. It also concludes that the effect of buoyancy is to contract the vortex spacing and accelerate the wake motion. The analytical derivation of this conclusion is presented in the next subsection, in which the model is described in detail. Extension of the derivation to a compressible stratified medium is straightforward, and requires only that the potential temperature or potential density gradient be used in the computation of the Vaisala period.

### 4.4.3 NEW THEORY FOR TRAILING VORTICES IN A STABLY STRATIFIED ATMOSPHERE

Nature of the Theory

The new theory is based on two approximations, best illustrated by dimensional reasoning. We focus attention on the $x-y$ plane and imagine that an aircraft passes through the origin of the coordinates at time $t=0$. The aircraft generates a pair of vort es of circulation $\pm \Gamma$ and initial separation $2 s_{0}$, and the vortices descend along the $y$-axis according to some displacement history $y(t)<0$. The displace, nent $y(t)$ can be transformed into a spatial trajectory $y(z / U)$, where $U$ is the speed of the aircraft and $z$ is distance behind, but nonsteady motion in the $x-y$ plane is a more convenient point of view for the present analysis.

If the atmosphere were of uniform density, then the vortex spacing would remain constant, and the pair would descend according to the formula

$$
y=-\frac{\Gamma}{4 \pi_{0}^{s}} t
$$

But we are interested in the case where the ambient density $\rho_{o}$ is a function of $y$, say

$$
\rho_{o}(y)=\rho_{0}+y \frac{d_{0}}{d y}+\cdots \cdot
$$

where $d_{f_{0}} / d y$ is negative in a stable atmosphere, and the first two terms of the Taylor series describe the density variation with sufficient accuracy near the vortices. Stratification implies that $y$ may depend on
$\rho_{o}$, on $d_{\rho_{0}} / d y$, and on the acceleration of gravity $g$, as well as on the vortex parameters introduced already. Thus, $y=y\left(t, \Gamma, s_{0}, \rho_{o}, d \rho_{0} / d y, g\right)$, which can be simplified by dimensional analysis as follows:

$$
y=\frac{\Gamma}{4 \pi N s_{0}} y^{*}\left(N t, N T, s_{0} / \mathrm{g} \mathrm{~T}^{3}\right)
$$

$N$ is the Brunt-Vaisala frequency for atmospheric gravity waves,

$$
N=\sqrt{-\frac{g}{\rho_{0}} \frac{d \rho_{o}}{d y}}
$$

and $T$ is the time required for the vortex pair to induct itself downward a distance $s_{o}$ in the absence of stratification:

$$
T=\frac{4 \pi s_{o}^{2}}{\Gamma}
$$

$y^{*}$ is the dimensionless trajectory of the vortex pair, depending on the dimensionless time $N t$, and on two dimensionless parameters NT and $s_{o} / \mathrm{gT}^{2}$. The possibility must also be admitted that the spacing 2 s between the vortices changes as the pair descends into a stably stratified atmosphere, so

$$
s=s_{o} s^{*}\left(N t, N T, s_{o} / g T^{3}\right)
$$

by similar dimensional reasoning. Note that the kinematic viscosity $v$ does not appear in the arguments of $y^{*}$ and $s^{*}$. The Reynolds number $\Gamma / v$ is enormous in practical situations, say $10^{8}$, and the theory accounts for the effect of stratification without reference to viscous or turbulent diffusion.

The parameter NT is a measure of the strength of the stratification relative to the internal dynamics of the vortex pair. If NT is small, then the vortices travel downward many times their own separation within a single period of the gravity waves that they generate. The parameter $s_{o} / \mathrm{gT}^{8}$ measures the ratio of fluid acceleration incuced by the vortices to the acceleration of gravity. Local acceleration of a fluid particle operating on the density gradient within it generates vor$t$ icity. If $s_{o} / \mathrm{gT}^{2}$ is small, the vorticity arising in the flow surrounding the original vortex pair is mainly the result of gravity acting on density gradients. To estimate $N T$ and $s_{0} / g^{2}$, recall that

$$
\mathrm{N}=\sqrt{Y} \frac{g}{c}=0.035 \mathrm{sec}^{-1}
$$

for an isothermal atmosphere, where $y$ is the ratio of specific heats, $c$ is the adiabatic speed of sound, and the numerical estimate is based on $\gamma=1.4$ and the sea-level value $c=1100 \mathrm{ft} / \mathrm{sec}$. An isothermal atmosphere is untypically stable, so $0,035 \mathrm{sec}^{-1}$ can be taken as the practical upper limit of $N$. A Boeing 747 generates vortices of strength $\Gamma=9000 \mathrm{ft}^{3} / \mathrm{sec}$ and spacing $2 \mathrm{~s}{ }_{\mathrm{o}}=110 \mathrm{ft}$ during approach to landing (Crow and Olsen, 196\%), so the parameters entering the stratification problem are of order

$$
\mathrm{NT}=0.15
$$

at most, and

$$
\frac{\mathrm{s}_{\mathrm{o}}}{\mathrm{gT}^{3}}=0.10
$$

The theory is designed to exploit the fact that both parameters are much smaller than unity. In fact, the acceleration parameter $g_{o} / \mathrm{gT}^{2}$ is much

$$
4-52
$$

less important to the structure of the theory than NT, so the trajectory and vortex separation will emerge as a one-parameter family

$$
\begin{aligned}
& y=\frac{\Gamma}{4 \pi N s_{0}} y^{*}(\mathrm{Nt}, \mathrm{NT}) \\
& s=s_{0} s^{*}(\mathrm{Nt}, \mathrm{NT}),
\end{aligned}
$$

with

$$
\text { NT } \ll 1
$$

The idea of weak stratification is acceptable enough, but it leads to mathematical conclusions that seem to defy intuition. The point of view taken here is that the conventional intuition is appropria'e, if at all, to the limit NT $\gg 1$ opposite to the limi+ of aeronautical interest. Most previous theories imply that the recirculation cell surrounding the vortices becomes buoyant as it descends into heavier fluid and bounces upward after a time of order $N^{-1}$, as though the cell were a gas-filled balloon (cf. Saffman 1972). Those theories originate from the work of Turner (1960), who was concerned with vortices generated from rest by buoyancy. But here the wake is assumed to descend too fast for buoyancy to generate significant local vorticity. The vorticity responsible for altering the trajectory of the wake arises slowly, primarily in a region above the wate called the buoyant upwash. The flow induced by the upwash draws the vortices together and diminishes the volume of their recirculating fluid. Ulimately the vortices diffuse into one another, while the upwash relaxes into Boussinesq gravity waves, which rary the impulse of the aircraft wake to infinity.

## Hydrodynamic Driát

The passage of a vortex pair leaves no residual velocity disturbance in an ideal fluid. Fluid particles, however, suffer a ret downward displacement during the passage, called drift ir. the hydrodynamic literature. Figure 4-11.illustrates the drift of fluid particles above the closed recirculation cell of the trailing vortices. The heavy lines in the drawing can be regarded as filameats of ink, inscribed horizontally in the $x-y$ plane prior to the start of vortex motion. The filament BB below the recirculation cell as yet is hardly disturbed, but the filament $A A$ has been stretched around the cell and sucked downward behind it. As time passes the filament approaches an asymptotic displacement $h(x)$, which is permanent in an unstratified fluid. The drift $h(x)$ was introduced by Darwin (1953), who showed that the integral

$$
0_{0} \int_{-\infty}^{\infty} h(x) d x
$$

is an apparent mass, in this case the apparent mass of the fluid outside the recirculation cell. The drift $h(x)$ diverges logarithmacally at $x=0$ owing to the stagnation point of the recirculation cell. but the integral is finite.

Ne analytical selution for the drift of wortex pair is known, but $h(x)$ can be compnted numerically from the construction of Figure $1-12$. For the moment the cuordinates $x-y$ are fixed to the vortices, so the flow outside the recirculation cell is steacis. with an uphard velocity $V=r / f-s$ at infinity. Fluid particies $p$ and $O$ move on scparate streamlines. $P$ passing close to the recirculation cellatid Q remaining indefinitely far away. Suppose pand $Q$ come from the same stratum y far beiow the vortices asd ate presenily


Figure 4-1l. Formation of Drift Behind a Vortex Pair


separated a distance $y$ vertically. Then

$$
\frac{d \Delta y}{d t}=v-v(\Psi, y) \quad,
$$

where $v$ is the vertical component of velocity at $P$, $\psi$ is the stream function identifying the streamline on which $P$ is moving, and $y$ is the current location of $P$. Moreover,

$$
\frac{d \Delta y}{d y}=\frac{d \Delta y}{d t} \frac{d t}{d y}=\frac{V-v}{v}
$$

so

$$
\begin{equation*}
h=\int_{-\infty}^{\infty}\left[\frac{v}{v(\Psi, y)}-1\right] d y \tag{4-7}
\end{equation*}
$$

which gives the ultimate displacement of a particle on the streamline $\Psi$. The stream function for the flow around a pair of vortices located at $x= \pm s, y=0$ is given by the formula

$$
-y(x, y)=-V\left\{\ln \left[\frac{y^{2}+(x-s)^{2}}{y^{2}+(x+s)^{2}}\right]+x\right\}
$$

which can be solved for $x(\Psi, y)$ by iteration. The resulting value of $x$ can be substituted into the formula

$$
v=-\frac{\partial \Psi}{\partial x}=V\left[\frac{2(x-s) s}{y^{2}+(x-s)^{2}}-\frac{2(x+s) s}{y^{2}+(x+s)^{2}}+1\right]
$$

to determine $v(\psi, y)$. The procedure is repeated for each step of the integration 4-7, resulting in numerical values of the drift $h(x)=h(-\Psi / V)$.

The drift is best presented as a universal function $\eta(\xi)$, where

$$
\begin{equation*}
n=s(t) \pi(\xi), \xi=\frac{x}{s(t)} \tag{4-3}
\end{equation*}
$$

Here we take account of the fact that the separation $2 s(t)$ batween the vortices may change slowly with tinie, but the shage of theis recirculation cell and the shape of the drift they leave behind will be invariant.

Figure 4-13 is a plot of the dimensionless drift $n(\varepsilon)$. The function wa; obtained numerically by the procedure outlined above. Semi-logarithmic cosrdinates are used to emphasize the asymptotic behavior oin $\eta$ tor small $\varepsilon$, ard in fact the asymptotic expression

$$
n=1.073 \quad 7.298 \mathrm{en}|z|
$$

is rather accurre throughout the range $0<|\varepsilon|<1$. A grantity of major importance in the theory of wake descent is the integral

$$
\begin{equation*}
D=\int_{-1}^{1} \eta(\xi) d \xi=8.184 \tag{4-9}
\end{equation*}
$$

which also was evaluated numerically. It is a curiosity of the theory that the influcace of stratification depends upon the part of the drift between the vortex cores, as further analysis will show, rather than upon the integral.

$$
\int_{-\infty}^{\infty} \eta(\xi) d \xi
$$

proportional to the apparent mass of the recirculation cell.


## Buoyant Upwash

Drift would have no dynamical consequences in a neutrally stable atmosphere. The flow velocity would relax to zero after the passage of the wake, and fluid particles would remain permanently depressed. When the atmosphere is stratified, however, the drift results in a density distribution that cannot persist. The density distribution immediately above the wake is illustrated in Figure 4-14, where the darker bands represent the heavier strata of fluid. The wake leaves behind a density field

$$
\begin{equation*}
\rho(x, y)=\rho_{0}(y)+h(x) \frac{d \rho_{0}(y)}{d y} \tag{4-10}
\end{equation*}
$$

By our sign convention $h$ is positive, and $d \rho_{o} / d y$ is negative in a stably stratified atmosphere, so the density is reduced behind the vortex pair.

The next dynamical stage is that the depressed light fluid $\because$ ors upward. The initial buoyant motion is best treated in terms of tisi vorticity equation, which can be written as

$$
\begin{equation*}
\frac{\mathrm{D} \underset{\sim}{\underset{\sim}{2}}}{\mathrm{Dt}}=\frac{\nabla^{\rho}}{\rho} \times\left(\underset{\sim}{\mathrm{g}}-\frac{\mathrm{D} \underset{\sim}{\mathrm{D}}}{\overline{\mathrm{Dt}}}\right) \tag{4-11}
\end{equation*}
$$

for two-dimensional flow (vortex stretching is absent). Here $\underset{\sim}{〔}$ is the
 of vorticity following a fluid particle. The right-hand side is the Bjerkness force, coupling the density gradient $\nabla \rho$ to the acceleration of gravity $g$ diminished by the particle acceleration $D u / D t$. The ratio of $\mathrm{Du} / \mathrm{Dt}$ to $\underset{\sim}{\mathrm{g}}$ is of order $\mathrm{s}_{\mathrm{o}} / \mathrm{gT}^{3}$, a small parameter according to Section 1, 30 the particle acceleration can be neglected compared with gravity in Eqn. (4-11).


Figure 4-14. Density Field Above the Wake.

Suppose the current time is $t$, and the trailing vortices passed level $y(\tau)$ at some previous time $\tau$. If the time difference $(t-T)$ is short compared with the inverse of the Brunt-Vaisala frequency $N$, then the vorticity $S$ is in the initial stage of growth described by the linearized version of Eqn. (4-11), namely

$$
\frac{\partial \zeta}{\partial t}=-\frac{g}{\rho_{0}} \frac{\partial \rho}{\partial x}
$$

where the cross-product has been performed, and $\rho_{0}$ has been substituted for $\rho$ with sufficient accuracy in the denominator. By virtue of Eqn. (4-10).

$$
\frac{\partial^{\psi}}{\partial t}=-\frac{g}{\rho_{0}} \frac{d \rho_{o}}{d y} \frac{d h}{d x}=N^{2} \frac{d h}{d x},
$$

and from Eqn. (4-8),

$$
\frac{\partial \zeta}{\partial t}=N^{2} \eta^{\prime}\left[\frac{x}{s(\tau)}\right],
$$

where the prime denotes differentiation with respect to the dimensionless argument of $\eta$. The vorticity equation can be integrated from the time of wake passage $\tau$ to the current time $t$, with the result that

$$
\begin{equation*}
\zeta[x, y(\tau)]=N^{2}(t-\tau) \eta^{\prime}\left[\frac{x}{s(\tau)}\right] \tag{4-12}
\end{equation*}
$$

as long as

$$
N(t-\tau) \ll 1 .
$$

The ordinate $y$ is given the argument $\tau$ as a reminder that we are dealing with the stratum $y$ penetrated by the vortex pair at time $\tau$, when the vortex cores were separated a distance $2 s(\tau)$.

The drift $\eta(\xi)$ is positive and decreases symmetrica ${ }^{1} l y$ around $q=0$, so the vorticity $S$ is antisymmetric and negative for positive $x$. The vorticity distribution (Eqn. (4-12) thus represents an upwelling of gathering strength behind the original vortex pair. Figure 4-15 depicts the vertical component of velocity with.n the upweling flow. From the vantage point of the vortex pair, the flow appears as a concentrated upwash whose strength increases linearly with distance upward, a buoyant upwash. Figure 4-16 shows instantaneous streamlines induced by the buoyant upwash. The induced flow tends to draw the trailing vortices upward and together.

## Induction by the Buoyant Upwash

The trailing vortices move under their own induction and the induced field of the buoyant upwash. Suppose the righthand vortex is located at $x=s(t)$ and $y=y(t)$. Then it moves according to the equations

$$
\begin{equation*}
\frac{d y}{d t}=-\frac{\Gamma}{4 \pi s(t)}+v \tag{4-13a}
\end{equation*}
$$

$$
\begin{equation*}
\frac{d s}{d t}=u \tag{4-13b}
\end{equation*}
$$

where $\Gamma / 4 \pi s(t)$ is the downward speed induced by the left-hand vortex, $u$ is the $x$-component of velocity induced by the upwash, and $v$ is the $y$-component. Both $u$ and $v$ are to be evaluated at $x=s(t), y=y(t)$.


Figure 4-15. Profiles of the Buoyant Upwach.


Figure 4-16. Induced Field of the Buoyant Upwash.

The vorticity distribution $\delta$ determines $u$ and $v$ according to the construction shown in Figure 4-1.7 An area element $d A$ is located at $x, y(T)$ a distance

$$
\begin{equation*}
r=\sqrt{[s(t)-x]^{3}+[y(r)-y(t)]^{3}} \tag{4-14}
\end{equation*}
$$

away from the trailing vortex. A vorticity element 6 dA has evolved since time $T$ and currently induces a velocity

$$
\begin{align*}
& d u=\zeta \mathrm{dA} \frac{[y(\tau)-y(t)]}{2 \pi r^{a}}  \tag{4-15a}\\
& d v=\zeta d A \frac{[s(t)-x]}{2 \pi r^{3}} \tag{4-15b}
\end{align*}
$$

The induced velocity $u$, $v$ follows by integrating Eqns. (4-15) over the $x-y$ plane. It is convenient to introduce the transformation

$$
d \Lambda=d x \operatorname{dy}(\tau)=d x\left|\frac{d y}{d T}\right| d T
$$

Li $N(t-\tau)$ is everywhere small, the vorticity field (4-12) can be combined with Eqns. (4-15) to produce the following results:

$$
\begin{aligned}
& u=\int_{0}^{t} d \tau \int_{-\infty}^{\infty} d x N^{2}(t-\tau) \eta^{\prime}\left[\frac{x}{s(\tau)}\right]\left|\frac{d y}{d \tau}\right| \frac{[s(\tau)-y(t)]}{2 \pi^{3}}, \\
& v=\int_{0}^{t} d \tau \int_{-\infty}^{\infty} d x N^{a}(t-\tau) \eta^{\prime}\left[\frac{x}{s(\tau)}\right]\left|\frac{d y}{d \tau}\right| \frac{[s(t)-x]}{2 \pi r^{2}},
\end{aligned}
$$

where $r$ is given by Eqn. (4-14) A partial integration over $x$ re. moves the derivative on $\eta$, and finally


Figure 4-17. Framenork for Calculatine the RuoyancyInduced Field at a Vortex Core.

$$
\begin{align*}
& u=\int_{0}^{t} d \tau \int_{-\infty}^{\infty} d x N^{2}(t-\tau) s(\tau) \eta\left[\frac{x}{s(\tau)}\right]\left|\frac{d y}{d t}\right| \frac{2[y(\tau)-y(t)][x-s(t)]}{2 \pi r^{4}}  \tag{4-16a}\\
& v=\int_{0}^{t} d \tau \int_{-\infty}^{\infty} d x N^{2}(t-\tau) s(\tau) \eta\left[\frac{x}{s(\tau)}\right]\left|\frac{d y}{d t}\right|\left\{\frac{[y(t)-y(\tau)]^{2}-[x-s(t)]^{2}}{2 \pi r^{4}}\right\} \tag{4-16b}
\end{align*}
$$

Now consider the structure of Eqns. (4-13) and (4-16). The integrations over $x$ in Eqns. (4-16) could be carried out numerically if the current value $s(t)$ and its previous values $s(\tau)$ were known. The remaining integrals over $\tau$ involve the prior history of the vortex trajectory $s(\tau), y(\tau)$ weighted by the memory function $N^{2}(t-\tau)$. Equations $(4-13)$ and (4-16) are thus a set of integro-differential equations for the trajectory $s(t), y(t)$. They could be solved step-by-sted numerically up to a time when Nt ceases to be small. Beyond that time the memory function $N^{a}(t-T)$ becomes suspect over earlier times $T$. We should expect the memory function to become oscillatory in the manner of an atmospheric gravity wave.

To pose the correct gravity-wave problem. we must examine the integrand of Eqn. (4-16b) under the simultancous conditions

$$
\begin{gather*}
N(t-T) \ll 1, \\
\Delta y=y(T)-y(t) \gg s(T) \tag{4-1b}
\end{gather*}
$$

$$
(4-1: 3)
$$

Both inequalities can be satisfied if the parameter NT is sufficiently small, because

$$
\Delta y \sim \frac{(t-T)}{T} s .
$$

Inequality Eqn. (4-17a) means that we are dealing with a time lapse $(t-\tau)$ such that the simple vorticity formula Eqn. (4-12) is still valid, and $\angle \mathrm{qn}$. (4-17b) means that

$$
\frac{(\Delta y)^{2}-[x-g(t)]^{2}}{2 \pi r^{4}} \approx \frac{1}{2 \pi(\Delta y)^{2}}
$$

for any $x$ such that $\eta[x / s(\tau)]$ is appreciable. Then the integrand of the integral over $T$ is approximately

$$
\begin{equation*}
\frac{N^{2}(t-\tau)}{2 \pi(\Delta y)^{2}}\left[\left|\frac{d y}{d \tau}\right| s^{2}(\tau) \int_{-\infty}^{\infty} \eta(\xi) d \xi\right] \tag{4-18}
\end{equation*}
$$

where the term in square brackets is the total rate of volumetric displacement per unit time. The exact shape of the drift $n(f)$ is immaterial at large $\wedge y$, which suggests that we look for gravity waves generated by a concentrated impulsive displacement of fuid.

## Boussinesq Far Field

A concentrated impulsive displacement can be represented as a product of delfa functions:

$$
\begin{equation*}
\vec{v}(x, y, t)=-Q \delta(x) \delta(y) \delta(t) \tag{y}
\end{equation*}
$$

The meaning of Eqn. (4-19) is that an element of fluid of vanishing area $A$ is impelled downward an arbitrarily large distance $D$ in such a way that the product $A D$ equals a finite volume $Q$. The concentrated displacement imparts an infinite potential energy to the fluid, which serves as an infinite reservoir for the growth of kinetic energy later on. The problem is similar to the impulsive displacement of a free water surface, as treated at length by Lamb (1932).

Equation (4-19) is valid only near $t=0$.
Throughout the course of its evolution, the vertical component of velocity $\tilde{v}$ satisfies the Boussinesq wave equation

$$
\begin{equation*}
\frac{\partial^{2}}{\partial t^{2}} \nabla^{2} \tilde{v}+N^{2} \frac{\partial^{2} \tilde{v}}{\partial x^{2}}=\frac{\partial^{2}}{\partial t^{2}} \nabla^{2}[Q \delta(x) \delta(y) \delta(t)] \tag{4-20}
\end{equation*}
$$

where the inhomogeneous forcing function on the right-hand side insures that Eqn. ( $4-19$ ) is satisfied near $t=0$. The wave Eqn. (4-20) is based on the essumption that $\tilde{\mathbf{v}}$ is small enough to render nonlinear convection unimportant (cf. Yih 1969). We are interested in the response of the stratified atmosphere to drift imposed many vortex spacings away (inequality Eqn. (4-17b), so the linearization implicit in Eqn. (4-20) should be all right. Once Eqn. (4-20) is solved in the form

$$
\tilde{v}(x, y, t)=Q f(x, y, t)
$$

for quiescent conditions prior to time zero, then the solution for a displacement source moving down the $y$-axis can be written as a Duhamel's integral:

$$
\begin{equation*}
v(x, y, t)=\int_{0}^{t} f[x, y-y(\tau), t-\tau] \frac{d Q}{d \tau} d \tau \tag{4-21}
\end{equation*}
$$

If the theory is correctly formulated, then the integrand of Eqn. (4-21) at $x=0$ should match Eqn. (4-18) in the limit $N(t-r) \ll 1$. A condition for matching ought to be that

$$
\begin{equation*}
\frac{d Q}{d \tau}=\left|\frac{d y}{d \tau}\right| s^{2}(\tau) \int_{-\infty}^{\infty} \eta(\xi) d \xi, \tag{4-22}
\end{equation*}
$$

whereby $d Q / d T$ can be identified as the rate of creation of drift. The temporal behavior of $f(0, \Delta y, t-T)$ should yield a uniformly valid memory function, matching the function $N^{2}(t-\tau)$ of Eqn. (4-18) for small values of $N(t-\tau)$, and displaying oscillations for large $N(t-\tau)$. The approach resembles a singular-perturbation theory, with the vorticity distribution (4-12) as the inner solution, and the Boussinesq wave field as the outer solution.

The method for solving dispersive wave equations like ( $4-20$ ) was developed in full generality by Lighthill (1965). The solution is assumed to have the form of double Fourier integral,

$$
\begin{equation*}
\tilde{v}=\int_{-\infty}^{\infty} d k \int_{-\infty}^{\infty} d \ell e^{i(k x+l y)}\left[F_{1}(k, l) e^{i \omega t}+F_{2}(k, l) e^{-i \omega t}\right] \tag{4-23}
\end{equation*}
$$

in which $\omega(k, \ell)$ is chosen to satisfy the homogeneous wave equations for $t>0$, and $F_{1}$ and $F_{a}$ are chosen to mode! the particular excitation that takes place at $t=0$. In the case of Eqn. (4-20)

$$
\omega=\frac{N \mathbf{k}}{\mathbf{K}}
$$

where

$$
K=\sqrt{k^{2}+\ell^{2}}
$$

The phase velocity ${\underset{\sim}{c}}^{p}$ of the waves and their group velocity $\underset{\sim}{c} \underset{\sim}{c}$ play an : important role in evaluating Eqn, (4-23) in the limit of large Nt:

$$
\mathcal{E}_{p}=\frac{\omega}{K}\left(\frac{k}{K} e_{x}+\frac{l}{K} \underset{\sim}{E_{y}}\right)=\frac{N k}{K^{3}}\left(k \underset{\sim}{e} \underset{x}{ }+2{\underset{\sim}{y}}_{y}\right)
$$

and

$$
\begin{equation*}
\mathcal{E}_{g}=\frac{\partial \omega}{\partial k} e_{x}+\frac{\hat{\partial}}{\partial L} e_{y}=\frac{N \ell}{K^{3}}\left(\ell_{e_{x}}-k e_{y}\right) \tag{4-24}
\end{equation*}
$$

where $\underset{\sim}{e} x$ and $\underset{\sim}{e}$ arc unit vectors in the $x$ and $y$ directions. The phase velocity and group velocity are perpendicular to each other,

$$
\mathfrak{E}_{\mathrm{p}} \cdot \mathfrak{L}_{\mathfrak{g}}=0,
$$

which means the disturbance radiates outward as a fan, with radii of constant angle $\theta=\tan ^{-1}(x / y)$ being lines of constant phase. Mowbray and Rarity (1967) evaluated Eqn. (4-23) in the limit of large Nt. by the method of stationary phase and presented some beautiful photographs of the Boussinesq wave field around a point disturbance.

The methad of stations ry phase, however, is not applicable to the present analysis. In the first place, we need the behavior of Eqn. (4-23) in the limit of small Nt in order to match the inner solution (4-18) and secondly, we must fine $\overline{\mathrm{v}}$ on the y axis, where the method of stationary phase happens to break down. The origin of the breakdown is evident in Eqn. (4-24); waves propagating directly down the $y$-axis must have $\ell=0$, in which case $\underset{\sim}{c} g=0$. We must turn io the full solution (4-23), which takes the form

$$
\begin{equation*}
\tilde{v}=\frac{\partial^{2}}{\partial t^{2}}\left[-H(t) \frac{Q}{\pi^{2} N} \cdot \int_{0}^{\infty} d k \int_{0}^{\infty} d \ell \frac{K}{k} \sin \frac{k N t}{K} \cos k x \cos \ell y\right] \tag{4-25}
\end{equation*}
$$

for the problem posed in Eqn. (4-20). $\mathrm{H}(+)$ is the Heaviside unit step function, and Eqn. (4-25) is valld throughout all time $-\infty<t<\infty$. For $t>0$, the time derivatives can be carried through the integrations

$$
\begin{aligned}
& \tilde{v}=\frac{Q N}{\hbar} \int_{0}^{\infty} d k \int_{0}^{\infty} d i \frac{k}{K} \sin \frac{k N t}{K} \operatorname{coskx} \cos \ell y, \\
& t>0 .
\end{aligned}
$$

In parcicular

$$
\begin{equation*}
f(u, y, t)=\frac{N}{2} \int_{0}^{\infty} d k \int_{0}^{\infty} d \ell \frac{k}{K} \sin \frac{k N t}{K} \cos \ell y \tag{4-26}
\end{equation*}
$$

which gives the reoponse along the $y$-axis to a unit volumetric displacement at the origin.

Expression (4. 26) can be evalu3ted by irans* forming the variables of integration from $k, \&$ to $K, \ell$. Thus

$$
f(0, y, t)=\frac{N}{\pi} \int_{0}^{\infty} d K \int_{0}^{K} d \ell \sin \left(\frac{N t}{K} \sqrt{K^{2}-l^{2}}\right) \cos 2 y
$$

According to entry 3.876.7 in the tables of Gradshteyn and Ryzhik (1965),

$$
\int_{0}^{K} \sin \left(P \sqrt{K^{2}-R^{2}}\right) \cos b y d i=-\frac{\pi}{2} \frac{\partial}{\partial p} J_{0}\left(K \sqrt{y^{2}+p^{2}}\right)
$$

Thus

$$
f(0, y, t)=-\frac{N}{2 \pi} \int_{0}^{\infty} \frac{J_{0}^{\prime}\left[\sqrt{(K y)^{2}+(\lambda t)^{2}}\right]}{\sqrt{(K y)^{2}+(N t)^{2}}} N t K d K
$$

## Substitute

$$
K=\sqrt{(\mathrm{Ky})^{2}+(\mathrm{Nt})^{2}}
$$

with the result that

$$
\begin{align*}
f(0, y, t) & =-\frac{N^{2} t}{2 \pi y^{2}} \int_{N t}^{\infty} J_{0}^{1}(k) d k \\
& =\frac{N^{2} t J_{0}(N t)}{2 \pi y^{2}} \tag{4-27}
\end{align*}
$$

Equation (4-27) meets the objective of the Boussinesq-wave analysis. $J_{0}(\mathrm{Nt})$ is a Bessel function of the first kind, with asymptotic behavior such that

$$
f(0, y, t) \rightarrow \sqrt{\frac{N^{3} t}{2 \pi^{3}}} \frac{\cos (N t-\pi / 4)}{y^{2}}
$$

$N t \rightarrow \infty$.

The impulse response $f(0, y, t)$ oscillates in time but not in space, in accord with the disappearance of group velocity in the $y$-direction. The amplitude of the oscillation grows parabolically with time, but an infinite reservoir of potential energy is available to feed the disturbance as in the case of the analogous water-wave problem (Lamb 1932).

Combining equations ( $4-27$ ) and (4-21),
we find the Duhamel's-integral representation for the vertical component of velocity beneath a moving source of displacement:

$$
\begin{equation*}
v(0, y, i)=\int_{0}^{t} \frac{N^{2}(t-\tau) J_{Q}[N(t-\tau)]}{2 \pi(\Delta y)^{2}} \frac{d Q}{d \tau} d \tau \tag{4-28}
\end{equation*}
$$

where

$$
\Delta y=y-y(\tau)
$$

Since

$$
J_{0}(0)=1,
$$

the integrand of (4-28) does indeed approach (4-18) in the limit of small $N(t-T)$, provided $d Q / d$ is chosen to satisfy (4-22). Equation (4-16b), moreover, preserves its character in the limit (4-17a) and approaches (4-28) in the limit (4-17b), if the approximate memory function $N^{2}(t-\tau)$ is replaced with

$$
\begin{equation*}
\mathbb{N}^{2}(t-\tau) J_{0}[\mathbb{N}(t-\tau)] \tag{4-29}
\end{equation*}
$$

## Wake Trajectory

The trailing vortices move according to Eqn.
(4-13), in which $u$ and $v$ represent the indication of the buoyant upwasl. Equations (4-16) specify $u$ and $v$ for small values of $N(t-r)$, and the equations can be generalized for arbitrary times by replacing $\mathrm{N}^{2}(\mathrm{t}-\mathrm{r})$ with the more general memory function (4-29). The result is a uniformly valid set of integro-differential equations for the wake trajectory:

$$
\begin{align*}
& \frac{d y}{d t}=- \frac{r}{4 \pi s(t)}+\int_{0}^{t} d \tau \int_{-\infty}^{\infty} d x\left\{N^{2}(t-\tau) J_{0}[N(t-\tau)]\right.  \tag{4-30a}\\
& s(\tau) \eta\left[\frac{x}{s(\tau)}\right]\left.\left|\frac{d y}{d \tau}\right| \frac{[y(\tau)-y(t)]^{2}-[x-s(t)]^{2}}{2 \pi r^{4}}\right\} \\
& \frac{d s}{d t}=\int_{a}^{t} d \tau \int_{-\infty}^{\infty} d x\left\{N^{2}(t-\tau) J_{0}[N(t-\tau)] s(\tau) \eta\left[\frac{x}{s(\tau)}\right]\right.  \tag{4-30b}\\
&\left.\left|\frac{d y}{d \tau}\right| \frac{2[y(\tau)-y(t)][x-s(t)]}{2 \pi r^{4}}\right\},
\end{align*}
$$

with $r$ still given by Eqn. (4-14). The equations could be solved on a computer for given values of $\mathrm{r} . \mathrm{N}$, and the initial conditions

$$
y(0)=0, \quad s(0)=s_{0},
$$

and the exercise might produce some interesting results. Implicit in Eqns. (4-30), however, is the assumption that

$$
\delta=N T \ll 1 \text {, }
$$

where $T$ is the inner time scale $4 \pi_{0}^{s}{ }_{0}^{2} / \Gamma$ discussed on page 4-51 otherwise the buoyant upwash would commence before the drift had approached its asymptotic value. So far the full power of the assumptior: $\delta \ll 1$ has not been exploited.

To appreciate the simplification that takes place for very small values of $\delta$, we should scale the variables appearing in Eqns. (4-30) as follows:

$$
\begin{aligned}
& t=\frac{t^{*}}{N}, \\
& x=s_{0} x^{*}, \\
& s=s_{0} s^{*}, \\
& y=\frac{\Gamma}{4 \pi N s_{0}^{*}} y^{*} \\
& r=\frac{\Gamma}{4 \pi s_{0}}
\end{aligned}
$$

The asterisk denotes a dimensionless quantity. To avoid a proliferation of asterises, they will be omitted from the trajectory equations in favor of a siagle asterisk beside the equation number. Thus

$$
\begin{aligned}
\frac{d y}{d t}= & -\frac{1}{s(t)}+\frac{\delta^{2}}{2 \pi} \int_{0}^{t} d \tau \int_{-\infty}^{\infty} d \xi\left\{(t-\tau) J_{0}(t-\tau) s^{2}(\tau) \eta(\xi)\right. \\
& \left.\left|\frac{d y}{d \tau}\right| \frac{[y(\tau)-y(t)]^{2}-\delta^{2}[\xi s(\tau)-s(\xi)]^{2}}{r^{2}}\right\} \\
\frac{d s}{d t}= & \frac{\delta^{2}}{\pi} \int_{0}^{t} d \tau \int_{-\infty}^{\infty} d \xi\left\{(t-\tau) \cdot \tau_{0}(t-\tau) s^{2}(\tau) \eta(\xi)\left|\frac{d y}{d \tau}\right|\right. \\
& \left.\frac{[y(\tau)-y(t)][\xi s(\tau)-s(t)]}{4}\right\}
\end{aligned}
$$

where

$$
\begin{equation*}
r^{2}=[y(r)-y(t)]^{2}+\delta^{2}\left[\xi_{B}(r)-s(t)\right]^{2} \tag{4-32*}
\end{equation*}
$$

and the integration over $x$ has been transformed into an integration over $\&$. An examination of the order of magnitude of the various terms in Eqne. (IV-31) reveals that

$$
\begin{aligned}
& \frac{d y}{d t} \rightarrow-\frac{1}{s(t)}+O\left(\delta^{2} \ln \delta\right) \\
& \frac{d s}{d t} \rightarrow O(\delta)
\end{aligned}
$$

as $\delta \rightarrow 0$. The quantity $\delta^{8}$ en $\delta$ is very much smaller than $\delta$, so the buoyant upwash moves the vortices horizontally much more effectively than it retards their descent vertically. For the purpose of calculating the trajectory for very small $\delta$, it is sufficient to retain the righthand side of ( $4-31 * b$ ) as the only buoyant effect.

If $\delta$ were set equal to zero in the integrand of $(4-31 * b)$, then the integral would diverge near $(t-r)=0$. The effect of finite $\delta$ is to prevent the denominator $r^{4}$ from going to zero at $(t-\tau)=0$. The functions $s(T)$ and $y(T)$ influence the value of the integral only near $\tau=t$. With sufficient accuracy for the purpose of integrating in the limit of small $\delta$.

$$
\begin{gathered}
y(t)-y(\tau) \approx(t-T) \frac{d y}{d t}, \\
r^{4} \approx\left(\frac{d y}{d t}\right)^{4}\left[(t-\tau)^{2}+\frac{\delta^{2} s^{2}(t)}{(d y / d t)^{2}}(1-\xi)^{2}\right]^{2},
\end{gathered}
$$

and so forth. The outcome of such approximations is that

$$
\begin{gathered}
\frac{d s}{d t}=\frac{\delta^{2}}{\pi} s^{3}(t) \frac{|d y / d t|}{(d y / d t)^{3}} \int_{-\infty}^{\infty} d \xi \int_{0}^{t} d \tau \\
\frac{(t-\tau)^{2} \eta(\xi)(1-\xi)}{\left\{(t-\tau)^{2}+\left[\frac{\delta s(1-\xi)}{d y / d t}\right]^{2}\right\}^{2}}
\end{gathered}
$$

where the order of integration has been reversed. The integration over $\tau$ can be performed explicitly. Set

$$
a=\frac{s(l-\xi)}{d y / d t}
$$

Then

$$
\int_{0}^{t} \frac{(t-\tau)^{2} d \tau}{\left[(t-\tau)^{2}+\delta^{2} a^{2}\right]^{2}}=\frac{1}{2 \delta a^{-1}} \tan ^{-1}\left(\frac{t}{\delta a}\right)-\frac{t}{2\left(t^{2}+\delta^{2} a^{2}\right)} .
$$

which approaches

$$
\frac{\pi}{4 \delta|a|}=\frac{\pi}{4 \delta s} \frac{\left|\frac{d y}{d t}\right|}{|1-\xi|}
$$

in the limit of small $\delta$. Thus

$$
\begin{aligned}
\frac{d s}{d t} & =\frac{\delta}{4} a^{2}(t) \frac{|d y / d t|^{2}}{(d y / d t)^{3}} \int_{-\infty}^{\infty} \eta(\xi) \frac{(1-\xi)}{|1-\xi|} d \xi \\
& =\frac{\delta}{4} \frac{s^{2}(t)}{(d y / d t)} \int_{-\infty}^{\infty} \eta(\xi) \operatorname{sgn}(1-\xi) d \xi
\end{aligned}
$$

Since $\eta(\mathrm{g})$ is a symmetric function, the factor $\mathrm{sgn}(1-\mathrm{E})$ causes contributions to the integral from the region $p>1$ to cancel those from $\mathrm{e}<1$. Thus

$$
\int_{-\infty}^{\infty} \eta(\xi) \operatorname{sgn}(1-\xi) d \xi=\int_{-1}^{1} \eta(\xi) d \xi=D .
$$

where D is the constant evaluated in Eqn. (4-9). Evidently the buoyant upwash outside the span of the two vortices contributes nothing to their rate of separation. The phenomenon is analogous to a spherical gravitating shell, which exerts no force on a particle inude.

The trajectory equations have reduced to a very simple form in the limit of small $\delta$ :

$$
\begin{align*}
& \frac{d y}{d t}=-\frac{1}{s},  \tag{4-33*a}\\
& \frac{d s}{d t}=\frac{\delta D}{4} \frac{s^{2}}{d y / d t} . \tag{4-33*b}
\end{align*}
$$

to be solved in conjunction with the initial conditi jns

$$
\begin{equation*}
y(0)=0, \quad y(0)=1 . \tag{4-34*}
\end{equation*}
$$

The solution is also simple. Take

$$
\frac{d^{2} y}{d t^{2}}=\frac{1}{a^{2}} \frac{d s}{d t}=\frac{\partial D}{4 d y / d t} .
$$

or

$$
\frac{d y}{d t} \frac{d^{2} y}{d t^{2}}=\frac{1}{2} \frac{d}{d t}\left(\frac{d y}{d t}\right)^{2}=\frac{6 D}{4} .
$$

Integrate from 0 to $t$, with $d y(o) / d t=-1$ according to ( $4-33 * a$ ) and (4-34*):

$$
\left(\frac{d y}{d t}\right)^{2}+1=\frac{\delta D t}{2}
$$

or

$$
\frac{d y}{d t}=-\left(1+\frac{6 D t}{2}\right)^{\frac{1}{2}} .
$$

The vortex spacing follows at once from $(4-3.3 * a)$,

$$
\begin{equation*}
s(t)=\left(1+\frac{6 D t}{2}\right)^{-\frac{1}{2}} \tag{4-35*}
\end{equation*}
$$

and after a further integration with respect to time,

$$
\begin{equation*}
y(t)=-\frac{\frac{4}{5}}{3 \delta D}\left(1+\frac{6 D t^{3 / 2}}{2}\right)^{3 / 1} \tag{4-36*}
\end{equation*}
$$

The trailing vorticea gradually draw together under the induction of the buoyant upwash, and there travel downward under their own induct ion at an ever increasing spsed.

## Concluting Remarks

Solutions ( $4-35$ ) and (4-36*) of the trajactory eçux tiont can be written in cerme of dimensional quanition a followe

$$
\begin{align*}
& t(t)=a_{0}\left(1+\frac{D}{2} N^{2}+t\right)^{-\frac{1}{2}} \text {. } \\
& F(t)=-\frac{4{ }^{4} 0}{3 X(A)^{2}}\left[11+\frac{D}{2} N^{2}+t^{3 / 2}-1\right]
\end{align*}
$$

where

$$
T=\frac{4 \pi s_{0}^{2}}{T^{2}} \quad \text { and } \quad D=8.184 .
$$

Figure IV-18 is a plo. of the vortex separation $28(t)$ and displacement $y(t)$ for the example mentioned on page IV-52, a Boeing 747 during approach to landing ia an isothermal atmosphere. The effect of stratification is not very atrong usider such conditions. The vortex separation is halved, for example, only after 140 sec , by which time the mutual. induction instability (Crow 1970 ) would probably have destroyed the vortices anyway. The leveling of trailing vortices observed by Condit and Tracy (1971) should probably be attributed to instability rather than busyaney.

The really interesting coasequence of the solution (4-37) is qualitative. Intuitively it would acem that stabie stratification ought to have somewhat the samo effect as a rigid barior, which would sase the vortices to decelergte and spread apart. The simplest
 of a blob of fipid bouncine tader the action of buoyancy. It is hard to abudon the rigia bereier asui bouncing blobay intuitive modely. but that if what Eqs (6-37) ack te to do. Throughout the developricaf we have as.

 The rigidabaryfer mindogy might work for NT $\gg 1$, and phe texiboak
 by definition. Eertapl it thothit nat be to0 strpprising that the now is cualitatively difercix whon NT $\ll 1$.

The pliynicai model implied fy the incony sian be
 scesther.


Figure 4-18. Wake Trajectory for the Case $\Gamma=9000 \mathrm{ft}^{2} / \mathrm{sec}$. , $2 \mathrm{~s}{ }_{\mathrm{o}}=110 \mathrm{ft} ., \mathrm{N}=0.035 \mathrm{sec}^{-1}$. The dashed line applies for zero stratification.

Figure 4-19 illustrates the density field and atreamlines near the contracting recirculation cell. The white area represents fluid captured at the instant of wake formation, and the shaded area represents the denser fluid at the present wake location $y(t)$. The buoyant upwash is not represented (cf, Figs. 4-14 through -16) but is squeezing the recirculation cell, forcing light fluid to drain upward near the rear stagnation point. Vorticity arises at the interface between light and heavy fluid, but the trailing vortices sweep the vorticity back toward the rear stagnation point. The interfacial vorticity becomes the boundaries of the drainage filament, as described by Scorer and Davenport (1970). As far as the present theo $o_{i} y$ is concerned, the drainage filament is an interior detail of the buoyant upwash that eliminates the integrable singularity $\eta(0)=\infty$ of the drift. The wake acceierates downward but loses captured fluid and impulse to the buoyant upwash. Ultimately the impulse due to the aircraft radiates away as gravity waves, while the trailing vortex cores overlap and annihilate each other.

A preliminary study suggests that the interfacial vorticity may contribute significantly to the induced velocity at the trailing vortex cores. If so, the contribution should be added to the induced field of the buoy ant upwash treated in this report. The quantitative predictions of Figure 4-18 may change, but the qualitative conclusion that the vorLices draw together and accelerate downward will be reinforced. The effect of interfacial vorticity will be incorporated in a further report.


Figure 4-19. Flow Within the Contracting Recirculation Cell.


We now discuss the other factor which may be expected to affect the general core dissipation $\sim$ this is the presence of organized vorticity in the external field. Such a situation occurs in a uniformly sheared crosswind. It would be expected that the presence of this ambient vorticity (of one sign) would have opposite effects on the two core vorticity distribution; causing the cores to develop differently, and possibly causing the pair to tilt. This phenomenon has been fully discussed in the section on vortex transport. It is believed that this is the effect causing the appearance of the solitary vortex which is frequently e\%ceptionally long lived.

The uner process by which organized vorticity is developed in the dmbient fluid is buoyancy. Although the effects of buayancy are highly controversıal, it is clear that the buoyancy field will generate vorticicy both in the outur hlow due to ron-uniform vertical displacements of the stratified density layers, and on the cell boundary itself due to the density discontinuities between the cell flow and the external flow. A new approach $t$, this problem has been described and analyzed in Seciion 4. 4.

As postulated ir the encounter hazard section, it is believed that, with the exception of the solitary vortex, the regular vortex dissipation, by whatever process, is sufficientlv gradual that its effecis on alleviating the encuunter hazayd are not important. In most cases the vorter pair retains its danger potential until it id dc.troyed oy one of the two major instabilities. These - stabilities are discusssd in detail in the following sections.

### 5.2 LINKING INSSTABILITY

## Introduction

A theory for the stability of a pair of trailing vortices has been developed by Crow ( $19: 0$ ). This theory examines the time dependent displacemente of vortex ines under an initial hypothetical perturbation in the wake of an aircrafi. Under the mutual influence of the velcity fields assuciated with the perturbed vortex lines, the displacements become unstable and grow with time. Sinuous oscillations develop on the vortex lines, which, due to the instability, can grow sufficiently large to cause the two vortex lines to touch and subsequently link together. This linking occurs periodicaliy in the downstream cirection, $z$, at distances corresponding to the wavelength of the sinuour ascillation of the vortex pair. At each link point, each original vortex line is severed, the two originally parallel vortices join together, and a closed vortex loop or ring is formed. Such a vortex ring is formed between each iink point, resulting in a train of vortex rings as the wake of the aircraft continues to develop. The coherent flow in tise wake, characteristic of the energetic flow around the original organized trailing vortices, abruptly changes in character once the vortex rings have formed.

It is shown by Crow that several modes of instabilities can occur, although it is argued that the structure of atmospheric turbulence, acting in the capacity of providing forced turbulent excitation of the instability, imposes an overwhalming bias in favor of long waves, which are unstable only in the symmetric mode. Astual observations have confirmed this long wave interaction behavior, where, from the Crow theory, the wavelength is appioximately 8.6 times the original, undisturbed vortex spacing (Smith and Beesmer, 1959; Chevalier, 1973).

Vortex lines move in accordance with the velocity field of the surround-: ing fluid. Atmospheric turbulence provides a continuous disturbance of the vortex lines and convective turbulent excitation of the resultant instability. Once initially distorted, the vortex lines also move in accordance to their own mutual and self-induced velocity fields, which was the condition originally treated by Crow in the analysis of the instability. No attempt was made to account in detail for the effects of atmospheric turbulent excitation, except to state that such turbulence is the probable factor which excites the long wave instability.

The present analysis examines the effects of atmospheric lurbulence as an input forcing function for the instability. The time for the instability to grow to the point at which vortex linking occurs is found as a function of the vortex strength, $\Gamma_{o^{\prime}}$ and spacing, $b$, and the atmospheric turbulence dissipation rate, $\varepsilon$.

## Structure of the Atmospheric Turbulence

The vortices from the generating aircraft are originally undisturbed, but under the influence of atmospheric turbulence they are continuously deformed. The instability is thus initiated by turbulent convection, and further, the turbulence acts contimuously as a source of input exergy to drive the instability.

The structure of the atmospheric turbulence is assumed to be stationary with respect to the coordinates of the downward moving vortex pair. That is, the turbulent eddies are assumed large enough that any time variation of turbulence is small as the vortices move downward a distance $\ell=V^{T} T_{\ell}$ where $v=\Gamma / 2 \pi b$, and $T_{\ell}$ equals the time to linking. Typically, $v=.5-3 \mathrm{~m} / \mathrm{sec}$, ant $\mathrm{r}_{\ell}$ is of the order 50 sec for light to moderate turbulence. Thin requires a turbulent "cell" (in which the time variation of turbulence is small) of dimensions on the
order of 25-150m. It is also agoumed that within this cell the turbulence is isotropic. The turbulent input to the instability is thus only a function of the energy spectrum, $E(k, \varepsilon)$, where $k$ is the wave number.

The resulting behavior of the vortex motion under the influence of the turbulent field can be analyzed as a time-invariant linear filter interacting with the turbulent energy input spectrum integrated over all wave numbers, with $\epsilon$ constant. The frequency response function of the filter is obtained from the dynamics of the vortex instability, with a peak at $k_{0}$, the wave number associated with the maximum amplification rate of the instability.

Fromn Crow, $k_{o}=\beta_{\max } / b$, where $b$ is the undisturbed vortexpair spacing, and $\beta$ is the dimensionless wave number. $\beta_{\max }$ is of the order one so that $.02 \leq k \leq .10$ radians/meter, which is within the inertial subrange commonly encountered in geophysical flows (MacCready, 1962). Although turbulence supplies energy to the instability at all wave numbers, the filter action of the dynamics of the vortex instability responds only to the energy contained in the spectrum at ur near $k_{0}$. Wave numbers not in the inertial subrange will be sufficiently far to either side of $k_{o}$ that the effects of the filter function will have attenuated their contribution. Their exact analytical form is thus not critical, and the assumption of an energy spectrum described by that of the inertial subrange will be adequate over all wave space.

With these assumptions, the input turbulent forcing function for the instability is taken as the one-dimensional, transverse energy spectrum in the Komolgorov inertial subrange. This energy spectrum is givon (page 273, Tennekes and Lumley, 1972) as:

$$
\begin{equation*}
F_{22}(k)=\frac{12}{55} \bar{\alpha} \varepsilon^{2 / 3} k^{-5 / 3} \tag{5-1}
\end{equation*}
$$

with $\bar{\alpha}=1.5$. The actual turbulent forcing function that drives the instability, $F(\mathrm{k})$, will be some percentage of $\mathrm{F}_{22}(\mathrm{k})$, given by:

$$
\begin{equation*}
F(k)=\gamma F_{22}(k) \tag{5-2}
\end{equation*}
$$

$\gamma$ will be determined by the geometrical relationship which couples $F_{22}(\mathrm{k})$ with the dynamics of the instability.

Geometry of Vortex Interaction with the Turbulent Field

In order to determine the ratio, $\gamma$, between the total available one-dimensional, transverse energy, $F_{22}(k)$, and that which actually couples with the instability, $F(k)$, the relationship of the available turbulent energy to that contained in the oscillating vortices must be established with respect to the interaction between the geometrics of these two flows.

With reference to Figure $5-1$, the instability for the gymmetrical mode develops at each vortex in a plane inclined at $\theta_{\mathrm{g}}=48^{\circ}$ to the horizontal. At any downstream position, $z$, the perturbation velocities develop in the $x-y$ plane with velocity components $u$, and $v$ respectively.

Let:

$$
\begin{align*}
& U_{1}(x, y, z, t)=U_{1}(x, y, z, 0)=\tilde{e}_{x} u_{1}(z, 0)+\tilde{e}_{y} v_{1}(z, 0)  \tag{5-3}\\
& U_{2}(x, y, z, t)=U_{2}(x, y, z, 0)=\tilde{e}_{x} u_{2}(z, 0)+\tilde{e}_{y} v_{2}(z, 0) \tag{5-4}
\end{align*}
$$

be the turbulent velocities at the positions of vortices 1 and 2 , reapectively. The velocities at these two positions have the dame direction in the $x-y$ plane, so that $u_{1} / v_{1}=u_{2} / v_{2}$, although they may have a differ ent magnitude.


If a velocity in wave space is defined as:

$$
\begin{equation*}
\int_{-\infty}^{\infty} \hat{U}_{n}\left(k, x_{y} v, 0\right) e^{i k z} d k=U_{n}(x, y, z, 0) \tag{5-5}
\end{equation*}
$$

( $n$ equals 1 and 2), then, from Eqns. (5-3 and (5-4) follows the definition of $\hat{v}_{n}(k, o)$ and $\hat{u}_{,}(k, o)$ in a like manner.

Since th: ratios $\hat{v}_{n} / v_{n}$ and $\dot{i}_{n} / u_{n}$ for any conatant, arbitrary wave number: $k$, and at the same distance, $z$, remain constant for $n=1$ and $n=2$, it is possible to use tiese transformed velocities as forcing functions to obtain the relative geometrical response of the two vortices in the $x$ and $y$-directions. Differentiating Eqns. (9) of Crow for the symmetrical instability, and substituting Eqne. (8), with the transformed velocities added as forcing functions, the relationshipe between the components of velocity resulting from the turbulence, and the velocity components associated with the symmetrical instability mode are established:

$$
\begin{align*}
& \frac{\partial \hat{x}_{s}}{\partial t}=\frac{\Gamma_{o}}{2 \pi b^{2}}\left(1-\psi+\beta_{\omega}^{2}\right) \hat{y}_{8}+\left(\hat{U}_{2}-\hat{U}_{1}\right)  \tag{5-6}\\
& \frac{\partial y_{t}}{\partial t}=\frac{\Gamma_{o}}{2 r b^{z}}\left(1+x-\beta_{\omega}^{2}\right) \hat{x}_{s}+\left(\hat{v}_{1}+\hat{v}_{2}\right) \tag{5-7}
\end{align*}
$$

where the dimensionless interaction functions $Y(\bar{Y}), X(\beta)$, and w( 5 ) are defined by Crow.

The ingtablity develops in a piane inclinad at an angle $\theta_{\mathrm{g}}$ to the horisontal,and the vertical and horizaatal velocities of the developing inctabiltty, $v_{s}$ and $u_{k}$, are coupled by $v_{s} / u_{s}=\operatorname{kn} \theta_{j}$. Therefore, the equation of motion of the system ia alinear combination of Equs. ( $5-6$ and ( $5 \times 7$ ) differentiated with reapect to time. Because of the linear roiationship, it is euficient to deal with only $z$ single componeut
of the equation of motion for the purposes of establishing geometrical relationships between velocity components.

Differentiating Eqn. (5-6) and subatituting Eqn. (5-7) results in the $x$-component of the equation of motion:

$$
\begin{align*}
& \frac{\theta^{2} \hat{x}^{2}}{\theta^{2}}=\left(\frac{r_{0}}{2 \mathbf{x b}^{2}}\right)^{2}\left(1-\psi+\beta_{\omega}^{2}\right)\left(1+X-\beta_{\omega}^{2} \hat{x}_{B}\right. \\
& +\frac{\Gamma_{0}}{2 \pi 0^{2}}\left(1-\psi+\beta_{\omega}^{2}\right)\left(\hat{v}_{1}+\hat{v}_{2}\right),  \tag{5-8}\\
& \hat{x}_{g}(k, 0)=0, \frac{\hat{x}_{c}(k, 0)}{\partial t}=\left(a_{2}-\hat{a}_{1}\right), \tag{5-9}
\end{align*}
$$

with:
and

$$
a^{2}(k)=\left(\frac{\Gamma_{0}}{2 a b^{2}}\right)\left(1-\psi+\beta_{w}^{2}\right)\left(\alpha+x-\beta^{2} \omega\right)
$$

A solution existe of the form:

$$
\begin{equation*}
\hat{x}_{B}(k, t)=A+B e^{a(k) t}+C e^{-a(k) t} \tag{5-10}
\end{equation*}
$$

Substitution of Ę̧n. (5-10) in Eqn. (5-8) with the initial conditions (5-9), and from Crow, Eqn. (1E), the expreasion for ©,

$$
\tan \theta_{9}=\sqrt{\frac{1+\beta^{2} w}{1-\dot{\psi}+\beta_{w}^{2}}} .
$$

Fetulas in the fetermiation of the coesfictente $A, B$, ard $C$.

$$
A=-\frac{\left(\hat{v}_{1}+\hat{\theta}_{2}\right)}{\frac{T}{2 b^{2}}\left(1+\lambda-\beta^{2} w\right)} .
$$

$$
\mathrm{B}=\frac{\left(\hat{\theta}_{2}-\hat{\theta}_{1}\right)+\frac{\left(\theta_{1}+\theta_{2}\right)}{\tan \theta_{g}}}{2(k)}
$$

and

$$
C=\frac{\left(\hat{v}_{1}+\hat{v}_{2}\right)}{\frac{F_{0}}{2 \sigma^{2}}\left(1+x-\beta_{\omega}^{2}\right)}-\frac{\left(\hat{u}_{2}-\hat{u}_{1}\right)+\frac{\left(\hat{v}_{1}+\hat{v}_{2}\right)}{\tan }{ }_{a}}{2 a(k)}
$$

At large timea Eqn. (5-10) becomeg:

$$
\begin{align*}
& \hat{X}_{9}(k, t)-B e^{a(k) t}=\frac{\hat{U}_{f}(k, t)}{2(k)} e^{a(k) t}  \tag{5-11}\\
& B=\frac{\hat{u}_{g}(k, t)}{a(k)} .
\end{align*}
$$

from which $\quad B=\frac{{\hat{\hat{u}_{s}}(k, t)}_{3(k)} .}{\text {. }}$
and

$$
\hat{u}_{n}(k, t)=a(k) B=\frac{\left(\hat{U}_{2}-\hat{u}_{1}\right)+\frac{\left.\hat{v}_{1}+\hat{v}_{2}\right)}{t_{1}+n t}}{2}
$$

If Eaw the fidial dipplacement of ene of the vorticel from its


from which,

$$
\hat{v}_{s}=\frac{\hat{U}_{5}}{\cos \theta_{0}}=\frac{\left(\hat{u}_{2}-\hat{\mathrm{Q}}_{1}\right)+\frac{\left(\hat{v}_{1}+\hat{v}_{2}\right)}{\tan \theta_{3}}}{\operatorname{ing} \theta_{\theta}}
$$

$\tan \theta_{2}=\tan 48^{\circ}=1.11$, so that:

$$
\begin{equation*}
\hat{v}_{t}=\frac{\left(\hat{u}_{2}-\hat{u}_{1}\right)+\frac{\left(\dot{v}_{1}+\hat{t}_{2}\right)}{I_{n} 1_{1}}}{4 \sin \theta_{3}} \tag{5-14}
\end{equation*}
$$

If long wavet are considerad, that is, $L \gg y \gg b$. the furoulont velocity field within the "cell" of dimension $\&$ would resuh in velocity componente at each vortex position ensh that

$$
\hat{u}_{2}-\dot{u_{1}} \ll \hat{v}_{1}+\hat{v}_{2}
$$

and, from Ean. (5-14)

Thun, a copling boweet tie vertical vortex mozian asa the at-

 is that avilizhe Erow the turbutent field is:


$$
\begin{equation*}
F(x)=\gamma_{22^{2}}(k)=\frac{2,43 \bar{a}}{55 \sin ^{2} \theta_{8}} e^{2 / 3} k^{-5 / 3} \tag{5-16}
\end{equation*}
$$

## Dynamica of the Vortex Notion

The radial displacement of a vortex line from its original, undisturbed position is given by $r(z, t)$. Let:

$$
x(x, t)=\int_{-\infty}^{\infty} \hat{r}(k, t) e^{i k z} d \xi
$$

An exponential solution in time for $\hat{F}(k, t)$, similar to that for Eqn. (5-10) exists:

$$
\begin{equation*}
\hat{r}(k, t)=A^{\prime}+B^{\prime} e^{-1 k k}+c^{\prime} e^{-a(k) k} \tag{5-17}
\end{equation*}
$$

which ar large times becomes:

$$
\begin{equation*}
\hat{r}^{(k, l)}-B^{\prime} e^{a(i k) t}=\frac{e^{a(k) t}}{z(k)} \hat{V}_{f}(k) \tag{5-18}
\end{equation*}
$$

The mean quare rimplacement of vortax element in defined by:
$3 \pi 9$

Define a function $\mathrm{R}(\mathrm{k}, \mathrm{t})$ :

$$
\begin{equation*}
\left\langle r^{2}\right\rangle=\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} R(k, t) \delta\left(k+k^{\prime}\right) e^{\left.i!k+k^{\prime}\right) z} d k d k^{\prime} \tag{5-21}
\end{equation*}
$$

which has the property that:

$$
\begin{equation*}
\left\langle r^{2}\right\rangle=\int_{-\infty}^{\infty} R(k, t) d k, \quad \text { at } k^{\prime}=k \tag{5-21a}
\end{equation*}
$$

Thus, from Equs. (5.:9) and (5-23)

$$
\begin{equation*}
R(k, t) \delta\left(k+k^{\prime}\right)=\left\langle\hat{r}(k, t) \hat{r}\left(k^{\prime}, t\right)\right\rangle \tag{5-22}
\end{equation*}
$$

Alsc define:

$$
\begin{equation*}
F(k) \delta\left(k+k^{\prime}\right)=\left\langle\hat{V}_{r}(k) \hat{V}_{\mathbf{r}}\left(k^{\prime}\right)\right\rangle \tag{5-23}
\end{equation*}
$$

Combining Eqras. (5-20), (5-22), and (5-23) gives:

$$
R(k, t)=\left[\frac{e^{a(k) t}}{a(k)}\right]\left[\frac{e^{a\left(k^{i}\right) t}}{a\left(k^{\prime}\right)}\right] \quad F(k)
$$

end for $k^{\prime}=-k$ :

$$
\begin{equation*}
R(k, t)=\left[\frac{e^{a(k) t}}{a(k)}\right]\left[\frac{e^{a(-k) t}}{a(-k)}\right] F(k) \tag{5-24}
\end{equation*}
$$

The wavelength, $\lambda$, is related to the wave number, $k$, by:

$$
\lambda=-\frac{2 \pi}{|k|}
$$

Therefore,

$$
\begin{gather*}
a(k)=a(-k) \\
R(k, t)=\left[\frac{e^{a(k) t}}{a(k)}\right]^{2} F(k) \tag{5.25}
\end{gather*}
$$

Substikuting Eqn. (5-25) in Eqn. (5-21a) gives:

$$
\begin{equation*}
\left\langle r^{2}\right\rangle=\int_{-\infty}^{\infty}\left[\frac{e^{a(k) t}}{a(k)}\right]^{2} F(k) d k \tag{5-26}
\end{equation*}
$$

This integral is significant oris near the wave number of maximum amplification rate, $k_{o}$, for the symmetrical mode of instability. The term in brackets represents the filter function of the forced oscillation, and $F(\underline{y})$ represents the turbulent forcing function.

Equation (5-26) becomes: upon squaring the term in brackets:

$$
\begin{equation*}
\left\langle r^{2}\right\rangle=\int_{-\infty}^{\infty} \frac{F(k)}{a^{2}(k)} e^{2 a(k) t} d k \tag{5-27}
\end{equation*}
$$

Expanding $a(k)$ in the exponential term in a Taylor series and substituting in Eqn. (5-27),

$$
a(k)=a\left(k_{0}\right)+\left(k-k_{0}\right) a^{\prime}\left(k_{0}\right)+\frac{\left(k-k_{0}\right)^{2}}{2} a^{\prime \prime}\left(k_{0}\right)+\cdots
$$

with:

$$
a^{\prime}\left(k_{0}\right) \equiv 0
$$

so that:

$$
\left\langle I^{2}\right\rangle=\frac{2 F\left(k_{0}\right)}{a^{2}\left(k_{0}\right)} e^{2 t a\left(k_{0}\right)} \int_{0}^{\infty} e^{-t / a^{: 1}\left(k_{0}\right) /\left(k-k_{0}\right)^{2}} d k \cdot(5-28)
$$

$a^{\prime \prime}\left(k_{0}\right)$ is negative so that $t a^{\prime \prime}\left(k_{0}\right)\left(k-k_{0}\right)^{2}=-t\left|a n\left(k_{0}\right)\right|\left(k-k_{0}\right)^{2}$.
Equation (5-28) has a solution of the form:

$$
\begin{equation*}
\left\langle x^{2}\right\rangle=\frac{2 F\left(k_{0}\right)}{a^{2}\left\langle k_{0}\right)} e^{2 \operatorname{ta}\left(k_{0}\right)} \sqrt{\frac{\pi}{t \mid a^{\prime \prime}\left(k_{0}\right)!}} \tag{5-29}
\end{equation*}
$$

From the definition from Crow:

$$
a(k)=\frac{0}{2 \pi b^{2}} a(\beta), \quad \beta=k b,
$$

and

$$
\begin{aligned}
& \alpha\left(k_{0}\right)=\frac{\Gamma_{0}}{2 \pi b^{2}} a\left(\beta_{\max }\right)=\frac{\Gamma_{0}}{2 \pi b^{2}} a_{\max } \\
& a^{\prime \prime}\left(k_{0}\right)=\frac{\Gamma_{0}}{2 \pi} a^{\prime \prime}\left(\beta_{\max }\right)=\frac{\Gamma_{0}}{2 \pi} a_{\max }^{\prime \prime}
\end{aligned}
$$

The "lifetime" of a vortex pair is that time $T_{\ell}$ at which $\left.<r^{2}\right\rangle=$ (b/2 sin 的 ${ }^{2}$. Substituting this condition in Eqn. (5-29), and recalling that:

$$
\begin{aligned}
& F\left(k_{0}\right)=F\left(\frac{\beta_{\text {max }}}{b}\right)=\frac{2.43 \bar{a}^{2}}{55 \sin ^{2} \theta_{s}} \varepsilon^{2 / 3}\left(\frac{\beta_{\text {max }}}{b}\right)^{-5 / 3},
\end{aligned}
$$

Introducing non-dimensional time and turbulence parameters as:

$$
\tau=\frac{T_{2} \Gamma_{0}}{2 \pi b^{2}}
$$

and

$$
\eta=\frac{\varepsilon b^{4}}{\Gamma_{0}^{3}}
$$

Eqn. (5-30) reduces to:

$$
\begin{equation*}
\eta^{2 / 3} \frac{e^{2 \tau a_{\max }}}{\sqrt{\tau}}=\frac{a_{\max }^{2}\left(\beta_{\max }\right)^{5 / 3} \sqrt{\left|a_{\max }^{11}\right|}}{32(\pi)^{5 / 2} \frac{2.43}{55} \bar{a}} . \tag{5-31}
\end{equation*}
$$

Thus the condition of time to linking for the sinuous instability of a vortex pair has been stated in terms of the vortex parameters $\Gamma_{0}$ and $b$, and the turbulent dissipation rate, $\epsilon$. The relationship between these parameters involves the dimensionless amplification rate $\alpha$ ( $\beta_{\max }$ ) and wave number $\beta_{\max }$. From Crow, for the symmetrical mode, $\beta_{\max }=0.73$ and $\alpha\left(\beta_{\max }\right)=0.83$. A parabolic approximation of $\alpha$ vs $\beta$ gives $\alpha^{\prime \prime}\left(\beta_{\text {max }}\right)=-3.12$. With $\bar{\alpha}=1.5$, Eqn. (5-31) becomes:

$$
\begin{equation*}
\eta=0,00271 \frac{(r)^{3 / 4}}{e^{2.497}}, \tag{5-32}
\end{equation*}
$$

from which a "universal lifetime" niot may be constructed in terms of the dimensionless parameters $\eta$ and $T$. Such a plot is shown in Fig. (5-2). Figure (5-3) shows in dimensional form the time to vortex linking, $T_{\ell}$, as a function of turbulence dissipation rate, $\epsilon^{1 / 3}$ for various aircraft.

Figure 5-2. Non Dimenaional Vortex Pair "Lifetime" to Lunking, Eqn. 5-32.

Figure 5-3. Time to Vortex Pair Linking as a Function of Atmospheric Turbulent Dissipation Rate, $\varepsilon$, for take off.

## Correlation of Present Theory with Experiment

The predictiong of time to linking for various turbulence levels iollow straight from the theory, and involve no fitted constants. There is scant experimental information with which to compare the present theory. However, a series of experiments in which motion pictures were taken of the trailing vortices from a light plane were performed (Tombach, 1972). The linking phenomenon was observed in come of these experiments, and the time to linking was measured from the motion pictures for various lev is of atmospheric curbulence. The results of these experiments are presented in Figure (5-4) along with the present theory in which the vortex parameters, $\Gamma_{0}$ and $b$, characteristic of the light plane used (Cessna 170) were included.

The agreement between experiment and the present $t_{i}$. ory is quite good. The scatter in the data are well within reason, considering that the experiments were performed under field conditions, with random effects undoubtedly influencing the limited amount of data obtained.

When future experimental data becomes available with respect to the vortex linking phenomena for different classes of aircraft subjected to varying degrees of atmospheric turbulence, the constants in Eq̣. ( $5-32$ ) can be modified (if required) in a rational and consistent manner to account for such effects as non-elliptical wing loading, flap denections, wing/non-wing mounted engines, etc.

The present theory was developed for assumed "large" times. For this reason, the behavior of the predictions of the theory for amall times, say less than $10-20 \mathrm{sec}$, is no 18 reliable as for larger times. Modifications to the theory can be performed that will extend the range of the theoretical predictions inso the small time region. However, examination of Fig. 5-1 showe that for the clase of large airpianes that

would present a wake hazard, this small time region occurs in the area of atmospheric turbulence that is described as moderate-to-severe. It is expected that for such levels of turbulence, other vortex transport and decay mechanisms may dominate the wake behavior. In any event extreme accuracy in predicting wake behavior at the short lifetime end is not required since, from an operational viewpoint, practical factors other than wake turbulence require aircraft spacings of greater than 10 to 20 seconds.

## Range of Validity of Theory

Three assumptions were made in the development of the present theory which bear directly on the validity of the resulta. These assumptions were that the disturbance energy spectrum could be represented analytically as the two dimensional, transverse turbulent energy spectrum in the inertial subrange, that the turbulence is isotropic, and that vortex linking is a "long time" phenomenon.

The wave number of maximum amplification rate, $K_{0}$, is defined by $K_{0}=\beta_{\text {max }} / b$ where $9_{\text {max }}$ the dimensionless wave number, is of the order one, and $b$ is the vortex spacing. For the cless of airplanes that would present a wake hazard, $.02 \leq \mathrm{K}_{\mathrm{o}} \leqslant .10$ radians/meter. This range in wave number is within the atmospheric inertial subrange. Turbuient excitation of the inatability at wave numbers cutside of this range would be greatly attenuated fue to the filtering effect of the vortex dynamics. Thus. the inertial eubrange model of the turbulent energy spectrum over all wave numbers is valid since only those wave numbers at or near $K_{0}$ (which are actually within the inertial subrange) contribute significantly to the growth of the instability, and the model is accurate here.

The value of the turbulent digsipation rate, $e$. is only a scaling parameter and does not affect the analysia. The value of $\varepsilon$ enters into the ond result and in this respect it must be chosen proper!y. The figures
shown assume that $e$ has been chosen for the case of isotropic turbulence. This is realistic for situations where the aircratit wake existe in an unbounded medium (at altitude). Near the ground, the turbulence is not isotropic, and an effective $\epsilon$ should be used. The section on Meteorology gives guidelines for appropriate corrections for near the ground. However, near the ground, the vortices interact with the ground plane itself. This interaction produces velccity fields at the vortices which alters the nature of the dynamics assumed in the present analysis. Thus, near the ground the theory may not be directly applicable or at best the constants may differ from those quoted.
However, for the isotropic turbulence situation, the theory agrees very well with the limited amount of experimental data available. For the case of $\varepsilon \rightarrow 0$, the theory predicts an infinite time to linking. The figures of link time vs. $\varepsilon$ are dixawn with logarithmic axes so that this limit is not immediately obvious from the graphs.
The theory was developed for assumed large times in which the first order exponential term in the expression fer the radial displacement at the vortex lines dominates. Exactly what constitutes "long times" depends upon the type of airplane which is generating the wake. This is so because the amplification rate, a(k), contained in the exponential term depends upon the ajrcraft parameters $r_{0}$ and $b$. As an example, for a 727 aircraft, it is estimated that the error in the theory is labs than $10 \%$ for a time gfeater than about 30 soconds. Thus. for thit paricular aiperaft. the theory dots not give reshisic resulte for tintes less finan $\mathbf{j 0}$ seconds.
The theory can and will be medified to account for the "pmall tione" region. However, with respect to the eacoukter hazard. the fimall fiste region is of little practical ingaficance becauso factors cther than wate furbulence dictate aircraft spacian much greater timan watd exibt at "small fimes."

### 5.3 CORE BIORSTHNG

## Genexpl

One najor form of vortex instability is core burating or vortex breakdown. This mode concists of a sudden abrupt widening of the vortex core and in the case of a smoke-marked aircraft vortex, the diaappearance of the tracer elements. Thig phenomenca has been observed in the laboratory (Sarpiaya 1970) and in figight teets (Tombach 1972).

The qualitative details of the breakdown are still guite obscure. In the experiments of Sarpkaya, the firgt effect $\quad$ semed to be an axisymmetric disturbance under which the care expanded and cantra-ied smoothly. Downatreara of this was a sistinct apiral disturtance which was then followed by 3 dieorganized but roughly axisymmetric core widening, the final breakdown. In Tombach's experimenta the amene sequence of events seems to take place, although the inifal emooth axisymmatric bulging is not as digeinct.

Several explanasions of voriex byeakdown have been propozed
 tion of the axial velocity fiain, 196t; Boweol, 1969 , the conjugate
 Nonc of thes io entirely setimactory, yot all contain eommon demente and all stem parially sopportec by experiment. It apacirs agrees fat the sapid eniargenter of the vartex core is accompnotst by axial presture pradewte, amithat thin breaxdown can only oceur when the



 developroint with cave the core to approzch the sritical siate so that ia mont cases the sore cill eventully cavelop to a fiate capable of breakiown.

It apsears likely that this critical atate can be deacribed crudely an a function of the axial velocity on the centeg Line and the swirl ratio. This evirit ratio is the ratio of the maximum tangential velocity to the ireestream speed. A recent papar by Mager (1972) gives the critical conditions in fais agreement with Benjamin's results. Qualitative'y. both of theae papers claim that as the axial velocity reduces so dous the critical owirl ratio.

If the critical axial/tangential flow combinaion could be determined, then the braakdown prediction problem would reciuce to teafing for the critical state after computing the core development.

It appears that the core bursting-phenomenon is mat related so inferaction between the left and right elements of the horaesto vortex pair, but rather to the development of the core itfeli and this development is a funtion of the kinematic viscosity controling the corz development. Obecruations by Tombach (!gis) have shown that, withe feale of his Aight tens, in which vortices were generated by a single fnfine light aimplame (Cesona 170) core burging occurred first in light atmospheric turbulence. while in bigh ambient torixieace the Crow Ingeabilisy invariably terminated the Fortex isfe.

Thus it it passible that, at the ame turbutonce lewel, the fime scaies for core burstian and Crow Instahisty ara selated to wame thation






 fintous :nstability.

Thus it is possible (although not definitely substantiated) that Crow Instability is the most significant mode of decay for vortices characteristic of large tuansport aircraft. However under unusual, or artifically perturbed circumstances, core bursting may be important.

The development of the core is principally a function of time, vortex strength, ambient turbulence, viscosity and wing drag. Many attempts have been made to analytically determine axial and circumferential velocities in the core. The significance of such a calculation is that if core bursting is related only to local details of core velocity; then breakdown could be predicted if the criteria on velocity profiles were known and the profiles could be analytically determined. Consequently we discuss some of the global aspects of the core developinent problem, and specifically state the invariants involved.

## Force and Moment In Global Terms

Considering the steady flow behind a lifting ving, it is clear that in any plane normal to the flight path that the total cirag and lift must be conserved. For this analysis we will ignore the propulsive system, noting that usually this does not contribute to the lift and that its axial contribution will be a thrust associated with the slipstream due to the jet engines or propellors. Thus we will find that considering the wing only, application of the momentum theorem in an axial direction gives a drag on the wing associated with its profile and induced drag.

If we consider a large circuit, along the centerline and around one side of the flow system (Fig. 5-5) we see that on the outer bound. aries there can be no viscous torque since the flow is that of a potential dipole having velocity gradients decaying like the inverse cube of the distance. On the center line, even if vorticity is present, the shear

is zero by symmetry, so there can be no force due to lamirar viscosity. Thus no torque can be applied to the circuit $C$, so the impulse moment (loosely called the "moment of momentum") is conserved, implying that the angular momentum on eacis side is conserved. This observation seems first to have been made by Betz (1933) by considering a set of free trailing vortices. We digress temporarily to show that the polar moment of vorticity referred to by Betz is directly related to the impulse moment.

## Relationship Between Impulse Moment and Polar Vorticity Moment

It is shown in Lamb's Hydrodynamics that all potential motions can be described by an impulsive wrench applied to the fluid. This wrench consists of both force and couple constituents. The force term is most familiar and is usually called the impulse. We note that the impulse is not strictly equal to the momentum of the system, which in many cases is mathematically indeterminate. However, the impulse is frequently loosely referred to as the "momentum." In our analysis we refer to the moment of the impulse as the "moment of momentum," recognizing again that this term is mathematically indeterminate, but that the impulse moment is a well-defined quantity. Here we show how this is connected to the polar vorticity moment developed by Betz (1933).

Take a two dimensional field, with a curved line of vorticity, representing the shed vortex wake. The wake vorticity, $Y$, is given by $d \Gamma / d s$ where $\Gamma$ is the bound vorticity on the wing and ds the element of length in the wake plane. The potential $\phi$ in the wake plane is directly proportionsl to the circulation $\Gamma$, with scale constants which do not concern us. We assume the wake is symmetrical about the vertical centerline and consider only one side as shown in Fig. 5-6.


Figure 5-6. Shed Vorticity in Transverse Plane
Taking the impulsive moment about $0, M_{0}$, and referring to the figure, we get

$$
\mathrm{M}_{0}=\int \phi \mathrm{Rd} s
$$

Now, according $t$, Betz $_{3}$ the polar moment of the vorticity about $0 . M_{p}$, is given by

$$
M_{p}=\int y r^{2} d s
$$

We note that $\gamma=\mathrm{d} \mathrm{\Gamma} / \mathrm{ds} \sim \mathrm{d} \phi / \mathrm{ds}$ so get

$$
M_{p}=\int\left(\frac{d g}{d \theta}\right) r^{2} d s
$$

Integrating by parts gives

$$
\left.M_{p}=\phi r^{2}\right]-2 \int \phi r \frac{d r}{d \theta} d a
$$

The first term vanishes at the limits of integration since at the centerline $r=0$ and at the tip $\phi=0$. Now we note that $d r / d s=$ $\cos (\theta-\theta *)$, so obtain

$$
M_{0}=-2 M_{p}
$$

This shows that the Betz formulation, which is frequently simpler analytically, does in fact express the impulsive moment. The impulsive force and moment can be transferred to other centers. When taken about the centroid of shed vorticity the inpulsive moment corresponds to the angular "momentum" of the flow about that point and is a measure of the "swirl" in the vortex core and surrounding flow.

## Global Invariants

We have discussed the determination of lift, drag, and moment of momentum for a viscous vortex system. These three quantities must be invariznt. We note further, that if vorticity has not yet reached the centerline, as is the case during the early core development, then this total vorticity on each side is conserved even though it diffuses radially under viscons and turbulent influences.

If the total vorticity on one side is conserved, it can be shown by arguments in the Treffiz Plane (Munk, 1924) that the centroid of this vorticity (or the mean vortex span) must be fixed for the lift to remain constant. Thus, we cen summarize the results to state that the following four properties must be congerved along the wake on ne.ch side of the centerline at least within a few hundred spans downstream of the wing:
a. the total vorticity,
3. the centroid of vorticity,
c. the moment of momentum,
d. the axial force.

| As |  |
| :---: | :---: |
| As explained previously, after mixing or interaction between the |  |

It is of interest to note that as vorticity diffuses tbrough the cell, it is annihilated at the centerline. Then, for lift to be conserved, the centroid of the remainiag vorticity must move outboard. We see that this is consistent with the requirement of conserving vorticity polar moment, since aithough the total vorticity is reduced, it occupies a larger area.

In general terms the drag (or axial force) can be determined by applying the momentum equation to a normal plane ard the impulse moment calculated by taking the second moment of vorticity, To determine the axial force both pressure and velocity must be known. However, since the flow is in general not homoenergetic \{because of viscous dissipation) the pressure cannot be inferred from the velocity. Thus both pressure and velocity must, in principle, be known for the momentum theorem to be applied.

This difficulty is avoided $D ;$ a conventent approximation, valid when the core is atill fairly compact and circular in shape.

## Rupesentation of the Druk abu Mument Intagrals <br> for Vortex Cores

The standard sechnique of simplification ts to assume the region to be divided into two portions, an outer portion where the few is iryotational so that the total head is connerved and where the axias perturbation is negligible, and an inner core region utere if is assumed that the flow feld is axisymmeiric, conciating of a tangential and an



Figure j.7. Pregsure and Velocities Near Representative Vortex.

Superimposed on this flow is the wake-like flow associated with the profile drag of the wing, and represented by the boundary layer shed from the trailing edge. It is assumed that this spanwise vorticity and the streamwise lifting vorticity are "attached" to the same particles in the flow leaving the wing, and is thus swept into the core during the rollup process. Consequently, reduction in total head is present even before there is any dissipation due to motion in the core. Thus the profile drag of the wing (a wake-like component) will reduce or even eliminate the jet-like flow at the center of the vortex core.

In general terms, then, the swirling flow due to lift, and represented by induced drag, develops the initial jet-like flow; while the profile drag develops the wake-like flow. As the core develops, dissipation produces reductions in head, causing the axial flow to become more and more wake-like.

When the core becomes very large, the centrifugal pressure gradients redre and the final flow state is one of a constant pressure wake, with both the profile and the induced drag appearing as a velocity defect wake.

Of course, the magnitude and rate of change of axial flows depends on the ratio of induced to profile drag and the dissipative process. It is not certain whether the latter is laminar or turbulent. Howover, we see that this can qualitatively account for the observations of a jet-like flow near the wing developing late: into a normal wake-like axial flow.

This model so far conserves lift by conserving circulation and vortex span. We note now that if we assume vorticity and total head lose are uniform in the core, then the additional two global relations of impulse moment and drag are sufficient to determine the core radius and the head loss. Thus, in a crude fashion, one can obtain the major characteristics of the viscous vortex field, whichare the swirl number
and scale of the axial flow. Unfortunately, such an analysis can not
give any streamwise variation in the core parameters. To obtain an
insight into the core development with time (or axial distance), more
parameters must be introduced to relax the core assumptions of uni-
form vorticity and uniforns head loss.
An interesting approach to this was made by Mager (1972). He
considered a single vortex only, and ignored all but the core flow.
Then he assumed a normalized distribution for the tangential velocity
with no free parameters, and a normalized distribution for the axial
velocity with one free parameter, $a$, scaling this velocity. In this
model, of course, angular momentum of the core is not conserved,
but is continuously reduced by viscous torques on the circumference,
taken by the author to be laminar. Thus a core development with ax-
ial distance can be predicted.
Mager showed that, according to his assumptions, for given
initial conditions of angular momentum and axial drag, there were
in general two possible flow states, provided certain conditions were
met. The one state was characterized by a larger core ano lower ax-
ial flow speed than the other. Moving downstream, subject to viscous
attenuation, these states converged to a single solution. Beyond this
point no solution was possible. He defined the point at which no solu-
tions (in his assumed similarity form) existed as the critical point,
and derived a curve of critical axial velocity as a function ef swirl
number. This curve has a very similar character, and is also quite
close numerically, to that given by Benjamin.
In order to continue the flow beyond this critical point fat which
regular solutions vanished), Mager intróuced a further parameter
into the tangential velocity profile. With this parametor adied, it
was found that the flow beyond the critical point exhibited a very large
increase in core size (about doubling) and a severe reduction in axial
velocity, such that the flow near the centerline was ciose to stagnation.

Mager postulated that this model acccunted for the obeerved vortex Ereakdown phenomena, the two solutions in the regular laminar state representing the bulging and contraction of the axisymetric bubble, the vanishing of the axisymmetric solution representing the spiral instability, and the final axioymmetric solution (with the large core and large axial flow deficit) representing the vortex breakdown.

It is clear that the actual critical values obtained by Mager are dictated by the form of the axial and tangential profiles assumed. However, the forms used are properly continuous and appear falriy reasonable, thus other selections would probably not greatly affect the numerical results. However, the vanishing of the solutions, a key element in hypothesizing the opiral instability followed by breakdown, apparently occur because of the limited number of parameters. Additional profile parameters could be introduced, which would continue the real eolutions beyond Mager's critical value. It is believed that it would be valuable to make an extension of Mager's technique, extending the analysis irom a single axisymmetric vortex core to the vortex pair associated with an aircraft vortex. Hers one mould take into account the force and moment contributions of the outer inviscid fow and investigate numerically the signifisance of further parameters on the disappearasce of solutions. It shot!d be noted, foo, thaf Mager's solution does not congider unsteady hows so ary wave-like instatilities are excluded.

In this light we refer to $\neq$ new resulf by Bilanin and Widnall (l973) in which the unsteady core instabilisy is treated. Fiere the authors state that the Crow instability induces unsteany asial pressure perfurbationt Which in turn cause the poriex core to undergo breatdoun. Aciording Po this theory the breakdown will occur whete the two vortices afe furtheat agari.

Thio phenomenwathas been obecrved in lab texts and alwo sppents
 ancher undeady way of exciting vortex breaksown and in thus not inconsistent with the ofiker cheories guoted.

We note that thin does not imply that sinuous instability ie a necessary condition for core bursting. In fact, the movies of Tombach (1973) show many cases where core bursting is clearly occurring on a trailing vortex whichis oseentially rectilinear.

## Conclusion

The nature and precise mechanism of vatex kreakdown is still controversial. It is generally agrued that the braokdoun in always associated with adverse pressust gradicnit, appazently these may be either cause or effect. The conditions for breakdown to occur are related principally to the swirl ratio and the magnitude of the axial flow. No general agreement on this critical function has been reached but both Benjamin and Mager give similar results which can be expresed as Figure 5-8. Here we have expressed the axdal velocity as the ratio of the mean core axial velocity to the freestream flow, We stress that although scales are given this is oniy a representative sketch.

It should be noted that both supercritical and subcritical states are driven by dissipation towards the critical condition so that in general most cores will appreach the breakdown siate. We note that Mager defines a further diriding line in the $R, \bar{W}_{c}$ W diagram above which no breakdown is possible.

If proper condition for breakdown wers knoun, then it wotald ztill be necesany to be able to calculate core development to determine the core state.

One of the comtroversial factory of thif calculation in the tiecoue tranofer conziant and thether it is surbulent or laminar. Thim is certainly extremeiy important in any precisction scheme ance turbulent difsjpation woule catise breakdown $2 f$ orier aif magariwde mere rapidiy than lamias foce.


Figure 5-8. Entimated Vortex Breakcoun Diagram.
6. MESEOROLOCZCA ASPECTS
6.1 GENESAL

In thia section of the repart wa describe now, stability, and turbuiknce characteristics in the layer up to about 70m. Formulas interselating factors in thif layor are presenfed and diocisosed. Then a data syotem ie preaented winich $c: r$ be a starting point for the design of th; stian to monitor the atmoaphere for the evaluation of techaiques id monitior and forc ast wake behayior. An evential operational eyatem will se simpler than the research/evanation zyatem.

The 70 m height ar sag from consicaring that the grestest dager irom voriex-wake encounters eccurs at low altitisdes where there is too little he ight (time) for the ancount ering a ictaft to recover from an uncousi statude. Fos she standard medi m and large jet sransports. it is estimpted that thip raitical heigh: can reach us to about 70m, aind of courge anm langef geto freatar as the airn raft height abovaground decroases. Any ezact assentment w the hozati is romrlex and depends on many quajitice, Thif heighe als. concides apt reximateiy with the height along the glide slope $z$ : the mis "而maper of ar ins sytum. Concoatration on such an area as do tal ran ie use ut beczuse many vortex sourtes are ava lab' $\cdot$ amithe area bri ..it th, hasafious segion in the landing approach when the vortices from low-fyun and slow
 Ayine aircraft in high.

Eip Atmosphe: Enverasmant
 100m, have teen inensive' y sindief isa smach as this covars the ze-

"hore lhe vertical fluxes of momentum, heat, and moistire are strontost. Oul of the studies there heve evolved various formulas for the mean profiles and turbulent cnaracteristics, formulas usually bascol on a thoorotical treatment tailored by ampirical results. To summarize the situation, the overall understanding of the characteristics of these lowest layers can be deemed "good". In the simplest situations of hnmepencous terrain, and constant meteorological conditions, which do not have stringly stable lapse rates, there are satisfactory semi-empirical formulas available to represent conditions. In the more common complex conditions the application of the formulas is more difficult and the apparent success in applying the formulas orobably rel.te:s more to the fact that great accuracy is not required by the prohlem nor are the data availaule to show what accuracy was attained.

## A General View

The planetary boundary layer is the layer near the surface where momentum and heat flux can be large, i.e., where there is a flux lint between the upper flow and the earth's surface. The boundary layer height can be under 0.1 km or over 10 km , with 1 km being a common height in daytime. In most typical form, the boundary layer is divided into two parts: (1) the surface boundary layer, across which the vertical fluxes are constant (this constancy provides simplifications which yield rather simple equations for mean and turbulent properties), and (2) the total boundary layer, whose top is found where the fluxes become small.

The height (cm) of the surface boundary layer below which the stress magnitude varies by less than $20 \%$ is put by Lumley and Panofsky (1964) at $2000 T$ (where $T$ is the surface stress in synes $/ \mathrm{cm}^{2}$ ). Since $T$ is typically of the order $1-10$ dynes $/ \mathrm{cm}, \mathrm{h}$ is typically between $2 \cdot 10^{3}$ $2 \cdot 10^{4} \mathrm{~cm}$, or 20 and 200 m .


#### Abstract

In analogous ashion, the vertical beat flux, $H$, varies little with height near the ground. Above the first meter radiation can be neglected. Then with typical values of heat flux on a clear day the height $h^{\prime}$ of the surface layer comes out at about 50 m . However, on nights with little wind, $h^{\prime}$ will be much lower. The top of th- total boundary layer is where the link between upper flow and grourd becomes small -- say the momentum flux (shearing stress) drops to $1 \%$ of its surface value.

Turbulence is an inherent characteristic of the boundary layer. since it is turbulence (cr convection, where the turbulence is organized) which carries the fluxes. The difficult cases to handle are those where the lapse rate is strongly stable and the wind is light . . a common situation at night. Then the layers with appreciable vertical fluxes may be quite lo: -. a few tens of meters or less -- and there is no way to inier characteristics dloft from direct measurements near the ground.


## 6. 2 THE QUANTITIES OF INTEREST

## The Various Quantities

Present thinking suggests that the primary meteorological variables of importance in the transport and decay of a vortex-wake system are: the mean horizontal wind field which causes horizontal drifi, and the turbulence and stability fields which play a role in determining descent and decay of the vortex-wake. These wind, turbulence, and stability conditions need to be known throug out the region occupied by the wake during ite whole life until it has decayed to the point where it cannot const...ute a hazard. Another meteorological quantity will in some cases also prove to be of primary importance -- the mean vertical motion of the local atmosphere. Cne final property, the vertical shear of the horizontal wind, across the aircraft filight path, may be of importance.

Evidence is accumulating that stability of the ctmosphere may have only secondary importance in either vortex decay or vertical descent. Thus we do not consider stability a primary variable, but meas.. ure it because of its effect on determining turbulence, drift, or the location of vertical shears.

When we talk about the mean wind, horizontal or vertical, we are thinking of "mean" as referring to events taking place for a minute or two, the lifetime of the wake. Viewed on a longer scale, the same motion might well be considered ar turbulence. Such ronsiderations of scale and averaging are important in the design of an operational system.


Turbulence covers all wavelengths, and multiple correlations, in the three directional components, so one major problerr nere is to decide what wavelength range anci component is most appropriste for the problem of vortex-wake decay. One powerfui simplifying assuription which has been used by AeroVironment Inc. is that the turbulence of significance is in the inertial subrange of eddy size. Kolrnos orov's similarity hypothesis results in the conclusion that all statistical properties of turbulence within the inertial subrange relate only to $\epsilon$, the equilibrium rate of curbulent eddy dissipation. This picture of turbulence is so simple it permits many important formulas to be derived easily by dimensional analysis and it also permits useful empirical relations to be made even when all the physical connecting links may not be understood. Kolmogorov's similarity hypothesis starts from a very simple picture of turbulence mechanisms. It pictures that turbuience enters a system cotinually, primarily at large wavelengths. The large addies break down to smalier eddies, which eventually break down to molecular motions through. vis yosity effects. The smaller eddies are sc decayed from the large "inpu:" eddies that the smalier eddies have "forgotten" their an firy and sc cannot be aware af direction -- statistically their energy must be isotropic. There is an "inertial subrange" within the range of eddies which exhibit isolropy; the inertial subrange covess all tho ${ }^{\circ}$ except the tiny ones which are stroygly affected by viscosity. In the inertial subrange the only quantity to which any statistical property of the flow can be related is $s$, the equilibrium rate at which energy enters the system, cascades through the inertial suivange of eddy sizes, and is removed (as heat) by viscous effects. In the atmosphere well above the ground the inertial subrange of eddy sizes extends typically from a few centirneters to many hundreds of meters. Even though it is an idealized concept, the relations derived using it ate found to be surprisingly ycod repres antations of the atmosphere.

|  | The atmospheric wavelengths under discussion for vortex-wake oreakup are probabiy no larger than the primary wavelength for the Crow looping insiability ( 300 m for a large jet), and may even be considerably smaller. Thus they fat into the inertial subrange and ef is the appropriate turbulent quantity to use. The Crow-Kolmogorov quantitative instability theory considers that vortex motions in a $V$ shaped (approximately $90^{\circ}$ ) trough along the line of flight are the primary instability initiators, and assumes the strength of such motions to be related to $\varepsilon$. The theory assumes that symmetrical perturbations along this specific trough surface are as likely as motion of a single line in any single plane. It is certainly posaible that otter modes of motion, including correlations between vortex lines, are fundamental to the initiation of Crow instability. Fiowever, all modes are within the inerial subrange, and so $\varepsilon$ is the turbulent quantity to observe and to fit into thecries about vortex lifespan as related to atmospheric turbulence. <br> If we were concerned with wakes only at allitudes of, say 300 m or more, thus we could probably comfortably use $e$ as the most appropriate turbulent quantity and avail ourselves of the simplicity inherent in inertial subrange concepts. However, we are actually interested in wakes at $70 \mathrm{~m}, 50 \mathrm{~m}$. and often even lower. Studies have shown (MacCready, 1962, and the discussion by Lumley and 2anofaky, 1904. pg. 166-167) that near the ground the spuctrum lawi seem reasonably valid even at wavelengths about twice the height of observation fand still longer in unstable cases). It may seem surprising that wayelengthe greater than $z$ can wctually show statistical energy isotropy and agree with the simplified spectrum laws. but it should be remembered that in our turbulence formulat we are dealing with energies and velocities, not final displacements, and so inotropy of encrgy for the three directional components for 100 m horizontal wavelengths may De reazonable at a height of ordy 30 m . |
| :---: | :---: |

In summary, the strength of the inertial subrange conc. pt is so great in the wake area that the monitoring or predicting of its sole vas $1-$ able, $\in$, may be quite useful for any evaluation program of monitoring/ predicting oystems. However, we must realize the possible increasing inadequacy of the inertial subrange concept for wake decay prer ztions as we approach the ground, and therefore also consider other formulations for tnrbulence there, especially the vertical turbulence component at wavelengths near the Crow looping stability.

## 6. 3 BOUNDARY LAYER RELATIONSHIPS

## The Mnnin-Obukhov Similarity Theory

The Monin-Obukhov similarity theory develops uaeful relationships in the surface boundary layer based on certain quantities which are constant. Lumley and Panofsky (1964) provides a critical review of the concepts.

Throughout the surface boundary layer the momentum flux (and hence the shearing stress $T$ ) is constant, and the heat flux (H) is constant. Using these, together with $\frac{L_{T}}{T}$ (where $g$ is gravity and $T$ is temperature) which is another dimensional "constant", it is possible to develop a velocity ( $u^{*}$ ), a length (L), and a temperature ( $\mathrm{T}^{*}$ ) which are also consiant and are three convenient dimensions for scaling. When the principal variables, such as temperature, wind, and height, are expressed non-dimensionally as fractions of these quantitites, a series of non-dimensional equations results that are of general validity in the surface boundary layer for conditions not ioo far from neutral. For simplicity, we neglect humidity in this discussion since its effects are orditarily small.

The equations reduce to simple laws for the neutral stability case ( $H=0$ ), but turn out to be rather complex and awkward to apply when $H \neq 0$. Much attention has been paid to empirical approximations to the basic equations. With the final equations having a general form which relater well to the fundamental theory.

The friction velocity. $u^{*}$. is related th the shearing stress ? and density 0 by

$$
\begin{equation*}
s *=(T i p)^{\frac{1}{2}} \tag{6-1}
\end{equation*}
$$

Since the are in the "constant" stress region, T is the same an © . the surface stress. The wind direction is constant in the surface botndary layer.

## The scaling length $L$ is given by

$$
\begin{equation*}
L=u *^{3} C_{p} p T /(\mathrm{kgH}) \tag{6-2}
\end{equation*}
$$

where $C_{p}$ is the specific heat at constant pressure, and $k$ is a dimensionless coefficient (the vou Karman constant, $K=0,4$ ). The dimensionlese height $\frac{z}{L}$ turns out to be especially useful for quantifying the effects of non-neutral lapse rates. L, a length which can be negative as well as positive, is basically a comenient lapse rate parameter. $\left|\frac{1}{30}\right|$ defines the height below which mechanical turbulence is dominant.

$$
\begin{equation*}
\frac{z}{L}=\frac{-p r o d u c t i o n ~ r a t e ~ o f ~ c o n v e c t i v e ~ e n e r g y ~}{\text { production rate of mechanical energy }}=\frac{z k g H}{u * C_{p} p T} \tag{6-3}
\end{equation*}
$$

Because $H$ is usually not measured, another quantity $L^{\dagger}$ is often used rather than L.

$$
\begin{equation*}
L^{\prime}=L K_{h} / K_{m} \tag{6-4}
\end{equation*}
$$

where $K_{m}$ is the eddy viscosity (erschange coefficient for momenturn) and $K_{h}$ that for heat flux. $K_{h} / K_{m}$ is near uni.y, being larger in unstable conditions and amaller in stable conditione.

$$
\begin{equation*}
L^{\prime}=\frac{u I^{\prime} \frac{\partial v}{\partial z}}{k g \frac{\partial \theta}{\partial z}} \tag{6-5}
\end{equation*}
$$

where $\theta^{\prime}$ is potential temperature. Ald the quantities in Eqn. (6-5) are fairly readily measurable.

Another measure oi "non-neutralness" is $\mathbb{R}_{f}$, the flux Richardson Number, which is sometimes involved in derivations. It if defined
as the ratio of the production of turbulent energy due to buoyancy forces to that due to Reynolds stresses (mechanical forces).

A gradient Richardson Number $R_{i}$ is defined by

$$
\begin{equation*}
R_{i}=R_{g} K_{m} / K_{n} \tag{6-6}
\end{equation*}
$$

$R_{i}$ is the ratio of buoyancy to inertia forces.

$$
\begin{equation*}
R_{i}=\frac{g}{T} \frac{\frac{\partial \theta^{\prime}}{\partial z}}{\left(\frac{\partial u}{\partial z}\right)^{2}} \tag{6-7}
\end{equation*}
$$

$R_{i}$ can be ascertained from mean wind and temperature measurements at two levels of a tower.

By various derivations, the logarithmic wind profile is found for neutral conditions:

$$
\begin{equation*}
\frac{\partial u}{\partial z}=\frac{u^{*}}{k\left(z+z_{0}\right)} \tag{6-8}
\end{equation*}
$$

where $z_{0}$ is the roughness height. For $z \gg z_{0}$, substituting Eqn. (6-8) in (6-7) and comparing with Eqn. (6-5) shows us

$$
\begin{equation*}
z / L^{\prime}=R_{i} \tag{6-9}
\end{equation*}
$$

for cases very close to neutral. The relation is more complex as we move away from neutral. The more complete relations are shuwn in Table 6-1. For $R_{i}$ greater than (more stable than) about 0.1 the ex-
change between layers diminighes so much that similarity laws are not applicable. For $R_{i}$ leas than (more unstable than) about -1 the region of free convection is reached and L'no longer is a useful scaling parameter.

TABLE 6-1, $R_{i}-z / L^{\prime}$ REI ATIONSHIP


The Wind Profile

The gradient for n of the wind profile valid for neutral conditions has already been given in Eqn. (6-8). Integrated, this becomes

$$
\begin{equation*}
U=\frac{u^{*}}{k} \ln \left(\frac{z^{+} z_{o}}{z_{o}}\right) \tag{6.10}
\end{equation*}
$$

which is weually compromised to

$$
\begin{equation*}
U=\frac{u^{*}}{k} k \frac{z}{z_{0}} \tag{6-11}
\end{equation*}
$$

since $z \gg 2_{0}$ where the formula is usually applied.

Fiedler and Panofsky $\left.{ }^{\prime} 1972\right)$ summarize information ow the validity of the logarithmic wind profile. They point out that Eqn. (6-11) may be valid only up to 30 m or so, and nffer a more accurate form which should serve up to 100 m :

$$
\begin{equation*}
U=\frac{14}{k} \hat{m} \frac{z}{z_{0}}+144 \mathrm{iz} \tag{b-11a}
\end{equation*}
$$

where $\{$ is the Coriolis parameter (i $=2 X$ radians per aecond rotation of earth $X$ sine oi latitude). Nrw evidence suggests the constant ist may be large. In any case, in the present review, for simplicity we will use Eqn. (6-11) rather than (6-11a).

The suriace from which $a$ is meagured is the grownd, for $\xi_{0}$ small, but ahould be elevaped when $z_{o}$ is large. For example, in $z$ forest the "surface" nay be set at about $2 / 3$ the beight of the treet.. abous where obstruction density is maximum.

If $U$ ve $a$ is known at two heights, wa ant $z_{0}$ can both be foumd from Eqn. (6-11). For accuracy, the calculation ia often made from a more complete wind protile. $J_{0}$ is a property of the local suriace.a constant for all condition (although at a giver point it will vary whth wind direction since it relates to upwind conditions). With $z_{0}$ known. Eign. (6-11) gives the whole U we aprofile tinoughout the iayer correspondiag to one velocity $U$ ai a particuiar masasuremeat meighs $\#$.

For nen-neutral conditions there are various approaches to modigisations to Egn. (6-11) based on the Monin-Obuktiov concepts, yieiding a lag-linear curve. The most complete form of the equation it
wher the univergal function $Y\left(\frac{2}{L_{0}}\right)$ is presented graphically by Lumley and Panoteky (1964) in Figuxe 6-1 reproduced on the following page.

Eq̧uation (b-12) coyefs adequately from the moat stable condisions far which the Moxin-Obukhov theory can hold iabout $R_{i}=0$.1) to atrongly unsmbie cases, say $\mathrm{F}_{\mathrm{i}}<-5$.

Note that for conditions very close to neutral $\mid{ }_{i}<0.01$ ) is given explicitiy in the legrend of Fig. 6-1. For greater atability, up to $R_{i}=0.1$, the explicit equation for $Y\left(\frac{L^{2}}{L}\right)$ is $\operatorname{till} 4.5 \frac{2}{L^{7}}$. although there is some controverey about the coofficient which may be at large ae 7.0 inftead of 4.5.

The bove relationd are dsived fer forced convection, there the vertical tranaier of heat and momentum ia more irnm mechanical sustulence linan by heat convection. Othey resulte are obtained lor free convection, when the relative esfect of heat convection t larger. eay $R_{i}$ - -1 , when mechanicai and keat energy are prodizect at about
 condtu $n$. Heat energy resuity in much more fitichent vertical trander of fropertise than it the case for mechamical onesty; he convective cidies are in ryer. with trore verliczl controity. For free convec-
 yosal argumart (mixing feath theary) cives:


Figure 6-1. Universal Function * for the Integrated Wind Profile. Main Graph for Unstable Stratification; Insert, for Stable Conditions. For $\left|z / L^{\prime}\right|$ or $\left|R_{i}\right| \leqq 0.01$, $\because=4.5 \mathrm{z} / \mathrm{L}^{\prime}=4.5 \mathrm{R}_{\mathrm{i}}$.


If $\mathrm{K}_{\mathrm{h}} / \mathrm{K}_{\mathrm{m}}$ is conntant with height, Eqn. (6-13) shows and $U \sim z^{-1 / 3}$, predictions which are reported to fit wiad observations on convective daye with moderate wind speecs quate well. In atrong convection with little wind, the $U$ profile is rather indeterminant.

Another approach, the power law method, has been found convenient for describing wind profiles, especially for fable conditions. It has no theoretical basis, wut has a flexibilisy and simplicity which makes it useful for talloring to empirical data.

$$
\begin{equation*}
\frac{u}{U_{i}}=\left(\frac{z_{i}}{z_{i}}\right)^{p} \tag{6-14}
\end{equation*}
$$

Where the absacript (1) refer to the conditions at a reference level and $p$ is an exponene waryiff between a and 3 . Panofiky. Biactadar. and Mryanil (1760) showed $p$ could berined from $=$. L', and $E_{0}$, and that the resultiog vabues on were consistent with numerous observed valuec. $p$ ie not a contian winh beyts of ehz layes concerned:
 stabie cases. If is found to depen. tidite upon wime poest: it in-



 cefund in Table $6-2$ on the followisg paçe.

TABLE 6-2. RELATION OF TURBULENCE TYPES TO WEATHER CONDIT!ONS

| A- Extremely unstable conditions <br> B- Minderately unstable conditions <br> C-Slightly unstable conditions |  |  | D- Neutral conditiona*E-Slightly stable conditionsF- Moderately stable conditions |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Surfare wind speed. m /srec | Daytime insolation |  |  | Nighttime conditions |  |
|  |  |  |  | Thin overcast or $21 / 8$ | s 9 |
|  | Strong | Morlarate | Sllght | cloudinesst | cloudiness |
| $<2$ | A . | A-B | B |  |  |
| 2 | $A \rightarrow B$ | B | C | E | $F \cdot$ |
| 4 | B | $\mathrm{B}-\mathrm{C}$ | C | D | E |
| 6 | C | $C-D$ | D | D | D |
| $>6$ | C | D | D | D | D |

* Applicable to heavy nvercist. day or nieht.
$t$ The degree of cloudiness is defined as that fraction of the sky above the local appaient horicon which is covermed by clouds.

These classes can be used to summarize data as follows:

TABLE 6-3. DATA SUMMARY

| Pasquill Class | $\delta \mathrm{T} / \mathrm{S}_{\mathrm{z}}$ | ${ }_{8}^{80}$ | $p$ |
| :---: | :---: | :---: | :---: |
| A | $\therefore-0.6{ }^{\circ} \mathrm{C} / 100^{\prime}$ | $25^{\circ}$ | . 15 |
| B | -0. 5 | $20^{\circ}$ | . 17 |
| C | -0.4 | $15^{\circ}$ | . 20 |
| D | -0.2 to -0.3 | $10^{\circ}$ | . 26 |
| E | -0.1 to 0.3 | $5^{\circ}$ | . 39 |
| $F$ | 0.4 to 1.1 | $2.5{ }^{\circ}$ | . 48 |
| (G) | (>1.2) | (1.70) |  |

The $\delta 7$ /es criteria come from USAEC Saiety Guide 23 (1972), and are ralated to Pasquill classes for molerate winds. They are measured between 10 and 45 m . The $\delta_{\varphi}$ (vertical-axis vane $R M S$ ) vaiues are $f$ from Safety Guide 23 and Gifford (1968), and Slade (1965). The $p$ values come from DeMarrais (1959) and apply to the rough Brookhaven site ( $z_{o} \sim 0.5$ to 1.0 m ).

Golder (1972) provides a more rational way ef relatitig Pasquil] classes to the quantitics on which the Monin-Obikhov laws are based. He shows how the Pasquill classes are functions of $z_{0}$ and $L$ (read L' or rather $1 / L^{\prime}$ for convenient plotting). Figure $6-2$ is taken from Golder's report.


Figure 6-2. $1 / L$ as a Function of Pasquill Classes and $z_{0}$

Since use of $z_{o}$ and $L^{\prime}$ permits one to draw conveniently on the rational and versatile equations baged o. Mcnin-Uubkhov, it geema preferable to use them instead of the empirical power laws which vary with site and altitude.

## $\varepsilon$ vs Height

The dissipation rate $\epsilon$, under steady state conditions, is equal to the mechanical production of energy $u^{*} \partial U / \partial z$ plus the buoyant production $g H /\left\langle C_{p} \rho T\right\rangle$ less the ilux divergence $1 / \rho\left(\delta F_{r_{m}} / \lambda z\right)$ where $F_{E_{m}}$ is the upward flux of total tu:bulent kinetic energy per unitaiter.

There is consicerable controversy about magnitude of the flux divesgence terr. The observations and theory cited by various authors for unstable conditions vary from the concept that the flux divergence term fully cancels the buoyant production term to the concept that the flux divergence term is negligible. Thus all formulations fit:

$$
\begin{equation*}
\varepsilon=u *^{2} \frac{\partial U}{\partial z}+C_{1} \frac{g H}{C_{p} \rho T} \tag{6-15}
\end{equation*}
$$

with $t_{i}$ value of the coefficient being between 0 and 1. G. Briggs (at NOAA Erivironmental Research Laboratories, Oak Ridge, Tennessee) has recently reviewed all the available evidence and concluded that $C_{1}=1 / 3$ gives the best fit with observations, and that with this vatue Eqn. (6-15) works reasonably up to $2 / 3$ of the total depth of the mixed regicn.

Using the methods discrssed by Hanna, Hutchison, and Gifford (1969), and setting $C_{1}=1 / 3$, Eqn. $(6-15)$ can be transformed to

$$
\begin{equation*}
\varepsilon=\frac{u *^{3}}{k z} \frac{1-\frac{1}{3} R_{i}\left(1-18 R_{i}\right)^{\frac{1}{4}}}{\left(1-18 R_{i}\right)^{\frac{1}{4}}} \tag{6-16}
\end{equation*}
$$

It should be noted that the turbulence intensity function commonly of interest is $\varepsilon^{1 / 3}$, not $\epsilon$, which minimizes the accuracy demands on the $\varepsilon$ vs $z$ equations. Taking the neutral approximation, with $R_{i}=0$. Eqn. (6-16) becomes

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=u * k^{-\frac{1}{3}} z^{-\frac{1}{3}} \tag{6-17}
\end{equation*}
$$

If we have already ascertained $z_{0}$ and we measure $U_{1}$ at height $z_{1}$, then from Eqn. (6-11)
and so

$$
\begin{equation*}
u^{*}=U_{1} k /\left(\ln z_{1} / z_{0}\right) \tag{6-18}
\end{equation*}
$$

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=\frac{U_{1} k^{\frac{2}{3}}}{z^{\frac{7}{3}} \ln z_{1} / z_{0}} \tag{6-19}
\end{equation*}
$$

The inverse relation between $\epsilon$ and $z$ is not inconsistent with the $\varepsilon$ observations from many sources summarized by Ball (1961). Ball examined data from 10 cm to 10 km , covering various roughnesses, winds, and stabilities, and although his summary plot shows a wide scatter the ground-related turbulence points , ave a $z^{-1}$ tendancy.
. In the next section, where we consider total vertical energy, there is further discussion about obtaining $\varepsilon^{1 / 3}$ in non-neutral conditions. The conclusion is that Eqn, ( 6 -17) is probably adequate in conditions well away from neutral.

An excellent, recent review of $\varepsilon \mathrm{va} z$ is given by Pasquill (1972). One of his main points is that surface roughness variability in the horizontal, which makes even many field research sites "non-ideal", can have a bigger effect on the $\varepsilon$ vs $z$ relationship than the difference between specific formulas.

## Energy vs Height

Lumley and Panofsky (1964) state that for neutral conditions in the surface boundary layer, the total turbulent energy $\overline{\mathrm{e}}$ should be proportional io $u^{* 2}$ and thus be given by

$$
\begin{equation*}
\overline{\mathrm{e}}=C U^{2} /\left(\ln \mathrm{z} / \mathrm{z}_{0}\right)^{2} \tag{6-20}
\end{equation*}
$$

where the dimensionless constant $C$ is of order one. The few available data are not in exact accord with Eqn. (6-20), and the authors note that although ihe nearby $z_{0}$ may determine the vertical motions, the $z_{o}$ far upstream can have an important bearing on the other components.

As to the vertical component of turbulent energy, $o_{w}$, containing all frequencies, the Monin-Obukhov theory indicates for a range of stabilities that

$$
\begin{equation*}
\sigma_{w}=A u^{*} \tag{6-21}
\end{equation*}
$$

where $A$ has a value probably near 1.3 in the neutral case, but slightly different values for other $\mathrm{R}_{\mathrm{i}}$ or $\mathrm{z} / \mathrm{L}$. Combining Eqns. (6-8) and


In stable air $\sigma_{w}$ decreases with height; in unstable air it slowly increases.

The value of $A$ found by Monin (1962), corrected by $\times 1.5$ as discussed by Lumley and Panofsky (1964), is given for approximately various stabilities in terms of $\mathrm{z} / \mathrm{L}^{\prime}$. The form shown by Panofsky (1970) is somewhet different, but since it represents more data we will use it. Putting it in equation form:

$$
\begin{equation*}
A=1.3-0.3 \mathrm{z} / \mathrm{L}^{\prime} \tag{6-23}
\end{equation*}
$$

This is a fair approximation for $-1.2<z / L^{\prime}<0.3$.

If one uses a vertical angle vane (horizontal axis) to measure turbulence, its RMS vaiue $\sigma_{\theta}$ in radians will be related to the RMS of vertical velocity $\sigma_{w}$ by

$$
\begin{equation*}
\sigma_{\theta}=\sigma_{w} / U \tag{6-24}
\end{equation*}
$$

ii we asoume angles are small so the sine is the same as the angle. Then from Eqns. $(6-22)$ and $(6-24)$ we see, for the neutral case,

$$
\begin{equation*}
\sigma_{\theta}=0.5 / \ln z / z_{0} \tag{6-25}
\end{equation*}
$$

In neutral stability, (or any strong wind condition) $\sigma_{\theta}$ is independent of $U$. Also, note that by measuring $\sigma_{\theta}$ at a certain height $z_{1}, z_{0}$ is readily obtained from Eqn. (6-25).

Equations which require knowledge of $L^{\prime}$ may not be particularly convenient since $L^{\prime \prime}$ 's calculation from $R_{i}$ requires observing both wind and temperature at two heights and performing long-term averaging to obtain the needed gradients. In the practical case it may be more suitable to estimate $H$ or $L^{\prime}$ or $R_{i}$. This is what AeroVironment Inc. did in developing its General Concentration Model whose application is described by the AV staff (1972). They used the method suggested by Panofsky and McCormick (1960), who postulated on the basis of similarity theory as applied to the constant flux layer that $\sigma_{w}$ should be a function of height, the rate of energy supplied by mechanical turbulence, and the rate of supply of convective energy. This approach has the further advantage of covering the free convection range as well as stabilities closer to neutral. The basic equation is:

$$
\begin{equation*}
\sigma_{w}=A_{2}\left[z\left(u *^{2} \frac{\partial U}{\partial z}+\frac{\delta g H}{\rho C_{p} T}\right)\right]^{\frac{1}{3}} \tag{5-26}
\end{equation*}
$$

where $A_{2}$ and $\delta$ are constants (assumed to be $A_{2}=1.25, \delta \div 2.4$ ). Using Eqns. ( $6-12$ ) and (6-26), one than has $\sigma_{w}$ as a function of $H$, $\Psi\left(z / L^{\prime}\right), z_{0}$, and $U$. Next, various forms of stability dependent $\Psi$ as proposed by Panofsky et al (1960) and McVehil (1962) were assumed in Eqn. $(6-26)$. The scale length $L^{\prime}$ is related to heat flux and friction velocity (eqn. ( $6-2$ )). The friction velocity further couples the mean wind and $L^{\prime}$ through the diabatic wind equation (Eqn. 6-12)). The unique way in which these parameters are related to one another makes it possible to determine $\sigma_{w}$ as a function of wind for given values of $z_{0}$ and $H$, although not in closed form. The results were plotted by AV in the form of $\sigma_{w}$ vs $U$ at 10 m for families of $z_{0}$ and $H$ values. The
roughness dependence was removed by using a roughness reduced velacity $U_{r}=U_{0}\left(\ln 10 / z_{0}\right) /\left(\ln 10 / z_{r}\right)$ where $U_{r}$ is the wind speed at 10 m corresponding to the roughness $z_{r}$. The "universal roughness re.. duced curves" provided a simple summary for computer storage giving quantitatively $\sigma_{W}=D U_{Y}$ for the appropriate strong winds, $\sigma_{W}=\mathrm{CH}^{2 / 3}$ for $H>0$ and $U_{r}<B H^{1 / 3}$ (positive $H$ but very low winds), and $\sigma_{w}=0$ for $H<0$ and $U_{r}^{r}<-A H^{1 / 3}$ (negative $H$ and very low winds). For $U$ and $\sigma_{w}$ in $\mathrm{m} / \mathrm{sec}, \mathrm{z}_{\mathbf{r}}$ in m , and H in cals $/ \mathrm{cm}^{3} / \mathrm{min}$, the constants are $\mathrm{A}=13.2, \mathrm{~B}=5.7, \mathrm{C}=0.87$, and $\mathrm{D}=0.125$.

The similarity in the derivations of Eqns, $(6-21)$ and (6-26), (and the uncertainty in each) suggests that one will not have much larger errors if one sets

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=\frac{A_{3} u^{*}}{k^{\frac{1}{3}} g^{\frac{1}{3}}} \tag{6-27}
\end{equation*}
$$

where the coefficient $A_{3}$ is assumed to be a weak function oi $z / L^{\prime}$. The variation of $A_{3}$ should be less than the cube root of the variation of $A$ with $z / L^{\prime}(E q n .(6-23))$ since the coefficient $C_{j}$ in Eqn. (6-15) is so much smaller than the $\delta$ in Eqn. (6-26). $\quad A_{3}=1$ at $z / L^{\prime}=0$, the neutral case. The implication here is that $A_{3}$ will be within $5 \%$ or unity even for $z / L^{\prime}=-1$. Since the controversy about the magnitude of constants involves indeterminancies much greater than $5 \%$, the temporal variability of $\epsilon^{1 / 3}$ certainly exceeds $10 \%$, and $z_{o}$ has local variations, for practical purposes well into the free convection range we can probably set $A_{3}=1$. Eqn. (6-17), the neutral approximation, is deemed adequate in non-neutral cases; it may be as suitable as any other simple equation.

For the RMS lateral velocity $\pi_{v}$, or the RMS angular variations of a direction vane with a vertical axis, $\sigma_{\phi}$, the situation becomes more
complex because of the increasingly large energies at increasingly large wavelengths (or periods) which are suppressed in the vertical turbulence covered above.

For the neutral case, various data suggest we can use as a first approximation

$$
\begin{equation*}
\sigma_{v}=2 u^{*} \tag{6-28}
\end{equation*}
$$

in analogy to Eqn. (6-21), but even in smooth terrain we should use a large $z_{o}$ to infer an "effective" $u *$ such as 1 m because large eddies dominate the quantity. There is aiways the problem that $\sigma_{v}$ depends on the sampling time used to derive it - . if longer wavelengths (times) are included, measured $\sigma_{v}$ is larger. In non-neutral conditions everything gets very complex. There is no comp!ete theory relating $\sigma_{w}$ to $U, z_{o}, R_{i}$, primarily because the instability aloft above the surface layer can have a large effect on the large scale lateral motions down low. Table III tries to summarize some of the data (note $\sigma_{\varphi} \sim \sigma_{V} / U$ ), and many other compilations are available (see, for example, Slade 1968, Pasquill, 1962, and Lumley and Panofsky 1964).

To summarize from the last reference: " $\sigma_{v}$ increases with increasing wind speed, at constant stability, particularly in stable air. It further is generally much larger in unstable air for the same wind speed, with the exception of low-wind speed inversions in which a gradual drift of wind direction may produce large standard deviations."

For the longitudinal component, $\sigma_{u}$, in anology to Eqns. (6-21) and ( $6-28$ ), we find in the neutral case $\sigma_{u} \sim u^{*}$, with the constent of proportionality empirically set at 2.5 . The relative variation of $\sigma_{u}$ with stability is a bit less than given for $\sigma_{w}$ in Table III. Panofsky ( $1: 70$ ), provides the latest summary plot of the constant of proportionality for $\sigma_{u} \sim u^{*}$ and $\sigma_{v} \sim H^{*}$, which we summarize in a later section.

## Spectra and Scales of Turbulence

The high frequency end of the atmospheric turbulence horizontal spectrum, the inertial subrange, in neutral and turbulent conditions, fits the "-5/3 law":

$$
\begin{equation*}
S\left(k_{1}\right)=b \varepsilon^{\frac{2}{3}} k_{1}^{-5 / 3} \tag{6-29}
\end{equation*}
$$

where $k_{l}$ is wave number in radians per unit length, $\epsilon$ is the dissipation rate, $S\left(k_{1}\right)$ is the energy per unit wave number, and $b$ is 0.50 for longitudinal spectra and 0.66 for lateral spectra. If $k_{1}$ is the wave number in cycles per unit length, $b$ is 0.15 for longitudinal spectra. (See the review by Panofsky, 1970, and Pasquill, 1972). There is still controversy abcut the value of constants. Franzen (1973) considers $b_{2}=0.55$ as the test estimate, and suggests that the relation between $b_{2}$ and von Karman's $k$ makes the latter take the value 0.35 .

With the ahape of the high frequency end of the spectrum known, it is possible to derive the intensity factor, $\varepsilon$, from any measurement which depends only on wavelengths within the inertial subrange. For example, we can obtain $\varepsilon^{1 / 3}$ from the $\sigma_{\theta}$ or $\sigma_{\varphi}$ signals where only high frequencies (in the inertial subrange) are used in the "aigma meter" giving the averages of $\theta$ or $\varphi$ resulting in $\sigma_{\theta}$ or $\sigma_{\varphi}$ (here we will call it $\sigma_{T_{3}}$ ). The filter characteristic time $T_{3}$ must be

$$
\begin{equation*}
T_{3}<0.44 z / \mathrm{U} \tag{6-30}
\end{equation*}
$$

(see MacCready and Jex, 1964); for accuracy it is safest for $T_{3}$ to be even considerably shorter than Eqn. ( $6-30$ ). Then the authors find

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=0.62 \mathrm{U}^{\frac{2}{3}} \tau_{3}^{-\frac{1}{3}} \sigma_{\tau_{3}} \tag{6-31}
\end{equation*}
$$

for $\varepsilon$ in $\mathrm{cm}^{2} \mathrm{sec}^{-3}, U$ in $\mathrm{m} / \mathrm{sec}, T_{3}$ in seconds, and $J_{\tau_{3}}$ in degrees. $\varepsilon^{1 / 3}$ can also be derived from high frequency wind speed fluctuations (see Franzen, 1965). The sensur must have a very short response distance to permit the measurement.

For the whole spectrum of vertical velocity up to $z=50 \mathrm{~m}$, Panofsky (1970) notes for neutral and moderately unstable cases that the spectra can be normalized to the following equation for neutral anc moderately unstable air:

$$
\begin{equation*}
\frac{k_{1} S\left(k_{1}\right)}{u *^{2}}=\frac{3.36 f}{1+10 f^{5 / 3}} \tag{6-32}
\end{equation*}
$$

where $f=k_{1} 2$, and $u^{*}{ }^{2}$ can be replaced by $J .6 \sigma_{w}{ }^{2}$. Note here $k_{1}$ is in terms of cycles (not radians) per unit length, and $f$ is a frequency normalized by height.
$k_{1} S\left(k_{1}\right)$ has a maximum vaiue of $0.43 u^{* 2}$ for $k_{1} z \sim 0.3$, that is, for a wavelergth about 3 times the height. In stable air the peak shifto to higher values.

More recently, Pasquill has examined the $S\left(k_{1}\right)$ spectrum in much greater detail. He considers the scale $\lambda_{m}$, the wavelength of the peak of a $k_{1} S\left(k_{1}\right)$ cure, where $\lambda_{m}$ is the wavelength oi maximum power input, expressible also in corresponding wave number by $\lambda_{m}=\left(K_{1 m}\right)^{-1}$. The situation can be summarized as follows:
a. In neutral flow $\lambda_{\mathrm{m}} / z$ is between 2 and 4 , and effectively constant with height in the first 20 m or so. The best evidence tends toward the value 2 .
b. The effect of thermal stratification is to increase or decrease the acale in unstable or stable conditions, and in effect to increase or derrease the height range with effectively linear increase.

Pasquill also discusses the Eulerian length scale, $L_{E}$, derived from correlation coefficents. Since this depends rather strongly on the large wavelengths, for which measurement (averaging) is diffirult, the relative variability of $\varepsilon_{E}$ is even greater than that for $\lambda_{\mathrm{m}}$. Summary"; above holds for $\ell_{E}$ as well as $\lambda_{m}$.

In final summary, for maximum accuracy the scale constanto in Eqn. (6-32) should be adjusted as a function of stability. However, we feel that for our practical application Eqn. $(6-32)$ as it stands is probably adequate.

## The Total Boundary Layer

Hanna (1969) exiamined many different methods for estimating $\mathrm{h}_{\mathrm{r}}$. the height of the mixed region conatituing the total boundary layer. He iound it ampossible to relste $h_{T}$ to quantities meaoured simply near the suriace, such as us, $U, U$, the geostrophic wind?. together with the Coriolis parameter. The beat iormula for $h_{T}$ larger than about 150 m was founs so be that of Laikhtman (isel). with the constant of proportionality altered to fit observations:

$$
\begin{equation*}
H_{T}\left(U_{z} \cdot \Delta \theta^{\prime}\right)=0.75 U_{=}\left(\frac{\operatorname{gat} \theta^{\prime}}{\Delta \Delta z}\right)^{-\frac{1}{2}} \tag{6-33}
\end{equation*}
$$

where $3 Q^{\prime} / \mathrm{A} \mathrm{z}$ is the average vertical gradieni of potential temperature through the boundary layer. Quoting from Hannt, when a complete vertical timperature sounding is available, "h $T$ is the lnweat level whien the vertical gradiens of temperature exhibity a diecontinaify.

For example, during well-mixed afternoon periods when an adiabatic layer near the ground is capped by a relatively stable inversion layer, ${ }^{h} T$ corresponds to the level of the base of the inversion layer. During a clear, calm night-time period where there is a ground based inversion, $h_{T}$ corresponds to the top of the strong inversion layer." For the alti:udes studied, $\Delta \theta^{\prime}$ was always positive (conditions always stable over this layer).

The conclusion is that temperature measurements aloft are needed to establish $h_{T}$ with any reliability. With such measurements, $h_{T}$ can be taken directly from the sounding. Eqn. (6-33) may be of some slight help if a partial $\Delta \theta^{\prime} / \wedge z$ is given.

## Empirical Values of $z_{0}$

The surface boundary layer equations are based on a picture of flat terrain with the roughness everywhere the same. The equations, however, are used for approximating reality in actual terrain, which can be considered as an assemblage of areas each with a different representative $z_{o}$, all being averaged together to give a final net effect. One effect of a change of roughness can be seen from the analysis by Panofsky (1968). He found that if wind blows over a surface whose roughness length abruptly changes to $z_{0}($ at $x=0)$, then below the height $z_{1}$ where

$$
\begin{equation*}
z_{1}=0.8 z_{0}\left(x / z_{0}\right)^{0.8} \tag{6-34}
\end{equation*}
$$

the boundary layer similarity laws can be applied based on the new $z_{o}$. This whole subject of changing and patchy roughness is extensively reviewed by Pasquill (1972). Where small $z_{0}$ is involved, it
takes about 1 kia fis $\therefore$ instuence to cover the beight of concern to us ( 70 m ) for neutrat conditions, ati much moze for atable conditions. For larger $x_{0}$, of for unstable cases. the influence establishes itself more quickly. Fop putchy roughness, the situation is obviously complex, and the resuer ia referred to Pasquill's paper.

Some representative ialues of roughess are given below:

| Smooth snow | 1.005 to 0.1 |
| :---: | :---: |
| Mown grass | c. 0 |
| Long grass (60-70m) | 6 |
| Open Country ( $0^{\circ} \mathrm{Neill}$ ) | 7,8 |
| Cotion Fiold | 50 |
| San Jose, suburb | 6 |
| Loncon, urbas | 75 |
| Pliladelphia |  os wind direction |
| Brookhaven tower (in woods) | 100 |

### 6.4 AREVIEW OF OPERATIONAL EQUATIONS

Pxeliminaty. Find $z_{0}, R_{i}, L^{\prime}, u \neq$
Assume one measures $U$ and $T$ at two heights, the subscript (1) referting to the lower at $z_{1}$ and the subscript (2) referring to the higher at $z_{2} \cdot U$ and $T$ designate averages, averaged over about 10 minutes.

1. In near neutral conditions, preferaily with strong winds, calculate $z_{0}$.

From $U_{1}$ at $z_{1}$ and $U_{2}$ at $z_{2}$ :

$$
\begin{equation*}
U=\frac{u^{*}}{k} t a z / z_{0} \tag{a}
\end{equation*}
$$

leada to

$$
\operatorname{mn} z_{0}=\frac{\frac{U_{1}}{U_{2}} \ln s_{2}-\operatorname{ta} z_{1}}{U_{1} / N_{2}-1}
$$

(b)

Second mett-s:

From of 28 any height $=$

$$
\begin{equation*}
\tan x_{0}=\operatorname{tn}=-\frac{0.5 v}{\sigma_{w}} \tag{c}
\end{equation*}
$$

## From $\sigma_{w}$ at ary $z$ and $U$

$$
\begin{equation*}
\ln z_{0}=\ln z-\frac{0.5 U}{\sigma_{w}} \tag{d}
\end{equation*}
$$

Third method:
Obtain $\epsilon^{1 / 3}$ from $U$ and $\sigma_{\theta}$ (or $\sigma_{\varphi}$ ) with a very short filter characteristic $\tau_{3}$ where $T_{3}<0.44 z / \mathbb{U}$. Then

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=0.62 \mathrm{U}^{\frac{2}{3}} \tau^{\frac{i}{3}} \sigma_{\theta} \tag{e}
\end{equation*}
$$

Obtain $z_{o}$ from

$$
\begin{equation*}
\ln z_{0}=\ln z-\frac{U k^{\frac{2}{3}}}{\varepsilon^{\frac{1}{3}} z^{\frac{1}{3}}} \tag{f}
\end{equation*}
$$

Use all three methods, and note agreement. Also note gysfematic variation of $z_{o}$ with azimuth angle $\varphi$.
2. Calculate $\mathrm{R}_{\mathrm{i}}(\overline{\mathrm{z}})$

$$
\begin{equation*}
R_{i}(\bar{z})=(g / T) \frac{\theta_{2}^{1}-\theta_{1}^{1}}{\left(U_{2}-U_{1}\right)^{2}} \bar{z} \ln z_{2} / z_{1} \tag{g}
\end{equation*}
$$

which gives the best finite differance approximation to the formula for Richardson Number where $\bar{z}=\left(z_{1} z_{2}\right)^{1 / 2}$ is the geometric mean of the heights $z_{1}$ and $z_{2}$.
3. Calculate $L^{\prime}$ at $\bar{z}$ using the appropriate equation shown or Table 6-I.
4. With this $L^{\prime}$, and using the ${ }^{\prime}\left(2 / L^{\prime}\right)$ relation from Figure $6-1$, calculate $u^{*}$ from the height data usiag

$$
\begin{equation*}
I=\frac{u^{i} k}{k}\left[\ell n z / z_{o}-\psi\left(z / L^{\prime}\right)\right] \tag{h}
\end{equation*}
$$

Then calcuiate $u^{*}$ from the height 2 data. The two $u^{*}$ values should be fairly close. Take their average as the best estimate of $u^{*}$. One could weight the calculation from the lower height more strongly since the $z / L$ effect is less at low heights, but the calculation from the higher height may be more suitable as a basis for extrapolation.

## Uvs z

1. Eqn. (h) gives $U$ vs $z$ for all heights in the surface boundary layer.

Evs $z$

1. Calculate $R_{i}$ for all the heights of interest from $z / L^{\prime}$ from the Table 6-I equations
2. At these heights, calculate $\varepsilon^{1 / 3}$ by

$$
\begin{equation*}
\epsilon^{1 / 3}=\frac{u^{*}}{k^{1 / 3} z^{1 / 3}} \frac{\left[1-1 / 3 R_{i}\left(1-18 R_{i}\right)^{1 / 4}\right]^{1 / 3}}{\left(1-18 R_{i}\right)^{1 / 12}} \tag{i}
\end{equation*}
$$

3. As an operational alternative, ignore the $R_{i}$ effects (which tend to be smail) and just assume

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=\frac{u^{*}}{k^{\frac{1}{3}} z^{\frac{1}{3}}} \tag{j}
\end{equation*}
$$

which is the equivalent of

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=\frac{1}{2^{\frac{1}{3}}} \frac{U_{1} k^{\frac{2}{3}}}{\ln z_{1} / z_{0}} \tag{k}
\end{equation*}
$$

If $U_{1}$ is measured at $z_{1}$, and $z_{o}$ is known, Eqn. (k) gives $\epsilon$ vs $z$. If only $\varepsilon^{1 / 3}$ is measured at $z_{1}$, Eqn. (j) then shows

$$
\begin{equation*}
\varepsilon^{\frac{1}{3}}=\varepsilon_{1}^{\frac{1}{3}}\left(\frac{z_{1}}{z}\right)^{\frac{1}{3}} \tag{1}
\end{equation*}
$$

| $\sigma_{\mathbf{W}}$ vs z |
| :---: |

1. Use $\sigma_{w}=\left(1.3-0.3 \mathrm{z} / L^{\prime}\right) u^{*}$
(m)
2. For neutral cases, thus $\quad \sigma_{w}=\frac{0.5 \mathrm{U}}{\ln z / z_{o}}$
and is constant with height.
3. Equivalentiy,

$$
\begin{equation*}
\sigma_{\theta}=0.5 / \operatorname{ln~z} / z_{0} \tag{0}
\end{equation*}
$$

4. If $H$ can be estimated rather than $R_{i}$ or $L^{\prime}$, use the algorithm provided in the discussion below Eqn. (6-26).
5. Consider $\sigma_{v} \sim u^{*}$ and $\sigma_{u} \sim u^{*}$, with the constants of proportionality being functions of stability as summarized on Figure! of Panofsky (1970). The constant is about 2.5 for all stabilities for $\sigma_{u}$, while for $\sigma_{v}$ it is about 1.5 for stable cases and exceeds 3 for unstable. The relations for any given case are found to be poor because the energy yielding $\sigma_{u}$ and $\sigma_{v}$ is not closely tied to local small-scale $z_{o}$ and so similarity laws are really inapplicable. If accuracy of $\sigma_{v}$ and $\sigma_{u}$ is required, they should be measured directly at one point and extrapolated upward and downward by physical reasoning and empirical results from various prior studies.

## Vertical Turbulence Spectra

1. For neutral and moderately unstable cases,

$$
\begin{equation*}
\frac{k_{1} S\left(k_{1}\right)}{u *^{2}}=\frac{3.36 f}{1+10 f^{5 / 3}} \tag{p}
\end{equation*}
$$

where $k_{1}$ is wave number in radians per $\mathrm{cm}, i=k_{1} z$, $S\left(k_{1}\right)$ is the power spectral density, and $u^{*}{ }^{2}$ can be replaced by $0.60_{w}^{2} \cdot k_{l} S\left(k_{1}\right)$ has maximum of $0.43 u *^{2}$ at $k_{1} z \sim 0.3$. In stable air this peak shifts to higher vailues of $k_{1}$.

### 6.5 AN APPROACH TO AN OPERATIONAL DATA SYSTEM

The system must be fully automatic, requiring no manual inputs. For eventual operational application it should provide information on mean air flow profiles and turbulence throughout a "target" volume about 1 km long, 300 m wide, and of thickness tapering from $\sim 70 \mathrm{~m}$ at one end to 50 m at the other. For initial research purposes where one is evaluating monitoring and forecast schemes, the volume of interest can be a bit smaller (especially as regards length).

For the system which is to aid in wake forecanting, the atmospheric characteristics it presents are estimates of future conditions where "future" means a few seconds up to several hundred seconds. Conceptually, the forecast could be rather accurate if an extensive observational system were set up and an associated extensive computational system. The observational setup would have to cover all the air which, for the period of prediction, would move into the target volume, For a strong wind of $20 \mathrm{~m} /$ sec, a forecast period of 200 secunds, and considering the vertical and horizontal variations of the wind, for one wind direction observations as far as 2000 m from the target would be needed, and from heights up to about 500 m and crosswind dimensions of 1 km or so. To cover all wind directions and speeds, obvicusly a dense three-dimensional network extending out several km in all directions would be required. Even with such a network and an elegant computational scheme, small-scale turbulence could not be forecast in detail.

For simplicity, atmospheric motions are usually divided into two scales of motion: (1) mean motion, explicitly described, and (2) furbulent motion, described anly in the form of statistical averages. Actually, the two scales form parts of one continuum. For vortex-wake transport and decay, it is customary to consider that the air's mean motion provides the medium for specifying drift, while the turbulence determines the decay mechanism and speed.

However, the whole motion of the wake evolves in some tens of seconds or a hundred seconds or so, which can be a time down in the turbulent eddy scale rather than the mean scale. To be specific, in unstable conditions the mean vertical wind at 50 m over a flat surface may be zero and yet an individual vortex may spend its minute of life in air rising at $2 \mathrm{~m} / \mathrm{sec}$. Similarly, there may be no crossflow component of vertical shear in the mean, but for the evolution of an individual vortex system the shear may be considerable.

Such considerations force one into some important practical conclusions with respect to forecasting vortex-wake characteristics.

1. Forecasting, to be of practical value, must be for a period of 10 minutes or rnore -- enough time so total terminal traffic flow can actually be adjusted to the anticipated conditions. Since there is no chance of forecasting individual eddies over such a time interval, the meteorological forecast is limited to providing only mean values .especially the most important limits of mean flow conditions and turbulence statistics. The forecasting data base must be at least 10 minutes, and experience at the site may show 30 minutes or even 60 min utes to be better.
2. For the monitoring of conditions in the ILS localizer beam near the middle marker, one wants to establish the mean flow, turbulence, and thermal stability throughout the Lagrangian parcel of air in which the observed vortex moves and did move. The prohibition against research airciaft and high towers in this region means direct measurements cannot be made. Indirect probing, say by doppler lidar for velocity and passive IR for temperature, are not deemed practical at the present stage of development. Thus one must do the best one can by extrapolations from surface (or low tower) measurements.

All such considerations seem to force the design of an operational system along the following lines. The suggested system:
a. is directed especially toward the evaluation of wake-vortex transport-decay forecasts,
b. involves some redundancies since the system, applied to the specific site, utilizing the not-entirely-satisfactory theoretical relations of the surface boundary layer, should in itself permit crosschecks,
c. should itself receive some evaluation and calibration by specific pibal releases and measurements from a test aircraft moving along the ILS path, and
d. is more complex than the minimum which might suffice in future years for operational airport uses.

The system consists of a network of tower-mounted meteorological sensors with appropriate data storage and real-time computing and display.

Tower $\$ 1$ will be 50 m high ( 70 m would be even better) and hopefully can be located within 2 km of the test volume where upwind surface conditions are not dissimilar from conditions upwind of the test volume. An existing airport control tower can perhaps be used, especially if a small meteorological tower is superimposed on it, but one must take account of the disturbance of a fat tower on $U$ and $\varphi$ (azimuth) measurements.

At $z=10,20,35$, and 50 m there should be measurements of $U, \varphi$ and $T$, from which 10 minute averages are derived.

Towers \#2, 3, 4, and 5 will each bo just 10 m high, instrumented only at the top, where $U, \zeta$, and $\sigma$ (elevation angle) will be monitored. The towers should be put at the corners of a square with 500 m sides straddling the landing path with the downwind side centered on the middle marker. The measurements will be processed to include 10 minute averages of $U, \zeta$, and $\sigma_{\theta}$, as well as higher frequency data to be discussed below.

Desirable, but not absolutely essential, measurements would be $\theta$ at $z=20 \mathrm{~m}$ and $\mathrm{z}=5 \mathrm{~m}$ on Tower \#1, and $T$ at $\mathrm{z}=2 \mathrm{~m}$ and 10 m on Towers \#2-5. A monostatic acoustic sounder giving information on low altitude layering during stable conditions would also be helpful. It would be located someplace within the square of Towers \#2-5.

Tower \#l is essential to give meaningful stability information, for which Towers \#2-5 are presumed low, It also can give this information aloft (near the 60 m height of the glide slope over the middle markers). This stability information tends to characterize conditions over a large area, hence Tower \#l need not be right at the test volume. In strongly stable c'nditions, vertical shears aloft in speed and direction can also be monitored from Tower \#1.

A T measurements on all the low towers may be found useful if the relative sensor accuracy is 0.1 C and careful averaging over all towers is done. This cannot give accurate $R_{i}$ or $L^{\prime}$ measurements, but can at least categorize the nore extreme stabilities and instabilities where the neutral-case equations may not suffice.

Tower \#2-5 are used first to derive mean $z_{0}$ characteristics for the area. Then, for operational forecasts, they provide $U$ and $G$ information from which meaningful average $U$ vs $S$ are derived to use as inputs to calculate $U$ vs $z, \varepsilon^{1 / 3} \mathrm{vs} z, \sigma_{w} v s z$, and vertical turbulence spectravs $z$. From the high frequency ends of $\sigma_{\theta}$ and $\sigma_{T_{3}}$ (evaluated by sigma meter) one also can calculate $\varepsilon^{1 / 3}$ at 10 m . and
from the low frequency sigma meter evaluation of $\sigma_{\theta}$ one derives "w at 10 m . These provide useful redundancy. Upcurrents and downcurrents are estimated from the convergences and divergences measured using the $U$ and $\zeta$ values of Towers \#2-5. For the area defined by Towers \#2-5, instantaneous components of wind perpendicular to the sides of the square will show the convergence-divergence situation. A similar calculation for the area defined by any three of the towers can give clues as to centering of the up or down current, if such is desired.

The research/evaluation system suggested here is presumably much more complex than an eventual operational system will be. After the "calibration" of the site which the use of this system provides, a final operational system could conceivably be as simple as (a) a single anemometer at 5 m or 10 m , (b) one measure of a quantity related to heat flux -- a temperature between 2 m (or ground) and $10 \mathrm{~m} .$. and (c) something to indicate the height of stable layers, such as an automated short-range acoustic sounder. For the critical stable cases some direct observations of stability and wind shear from a tower to at least $50-60 \mathrm{~m}$ is highly desirable, even if the tower is a few miles from the operational site. The acoustic sounder should at least be able to show the height above which turbulence is low, and thus illuminate critical hazard conditions. Radiation measurements are so poorly coupled to heat flux or stability-instability that they are not recommended.

The system detailed in this section does not give as mush detail $2 s$ one might want for checking the accuracy of monitoring system. Such detail could only come from indirect probing, as with a dopple: lidar which is considered not yet fully operational, or from a network of tall towers, which is inappropriate as long as the tests are located at an airport. Even though this system is not perfect for case enudies of monitoring devices, the system will likely be found satisfactury for that purpose.

## 7. SUMMARY

The general factors relating aircraft wing geornetry to the vorticity shed from the wing are presented. The rollup of the resulting vortex sheet into the classical wake consisting of a pair of conira-rotating trailing vortices is described, and a set of operational equations for principal wake characteristics is presented.

The wake encounter hazard with respect to a following aircraft is discussed. Axial a-d normal encounter danger factors are developed. These factors relate the aerodyramic response and control characteristics of the following aircraft to the wake naracteristics of the generating aircraft. It is recommended that the axial danger factor $D$ be used as a primary criterion for determining hazal lous en-runter situations.

The general conditions relating to the transport of a vortex pair are reviewed. The effects on the vortex pair dynamics du; to vortirity gen ration and re-distribution are discussed. Vorticity generation is caused by density variation in the fluid. Vorticiey re-distribution results from the iffusing vortex cores and the entrainment of mass and the detrainment of vorticity at the boundaries of the recirculation cella surround ing eqch vurtex. In gencral, this diffusion of the vortex cores rezults is departure from the classiesal equations describing the ..rotion of the vortex pair. He"ever if is pointed out that this is a relatively slow process, and usually, before it becomes significant, the vortex paiv has experienced on ${ }^{\circ}$ of several decay mechanisms linking, core ibstangh.

The eifectia of (crosswind) wind ahear are examined. In this case the geomatry of the vortex recirculation cells is dyastically atered. The upwind cell increases in gize and the downind cell shrinis. This effect increazes with shear, , na th probable effecss that wind shear might have on banking of the vortex pais and core bursing are discuased.

Previous experimental and theoretical works in the area of wake descent into a stratified atmosphere are discussed. The conclusion is that at the present there seems to be no motel that adequately describes the phenomenon of vorticity generation. A new theory for the descent of a wake into a weakly-stratified atmosphere is developed. For the condition of weak stratification (large Brunt Vaisala times) and fast wake descents (time for wake to descend one vortex span) - a situation whach normally occurs in practice - the theory concludes that the vortices will draw closer together with a consequent increase ir descent speed. Thus the effect of stable stratification alone (without entrainment) is to increase descent speed.

A discussion of the effects of eatrainment alone (without buoyancy' is given. If appears that entrainment will definitely retard wake descent, however the magnitude and mechanism of entrainment is controversial.

The general concepts regardirg Crow Instability are discussed. A new theory is developed as an extension of the Crow Stability Theory which assumes an atmospheric turbulent input to the grouth of the instability. The fortices are treated as a pair of linear ositlators with a fandom forcing function assumed to be that corresponding to the ont dimensional, transverse energy spectrum within the korsolgarovimertial subrange. The "tirne-to-linking" is found from the aew theory as a fun:tion of the vorter parameters, $\Gamma_{0}$ and $b$, and the furbulence dissipation rate, $\varepsilon$. This correlates very well with avaidable fight test data.

Core bursting is discusged with respec: to global invariania of core development. The nature and presiot mechanisn of vostex breakdown is stall controversial. It is generally agreed that breakdown is always associated with adverse pruasure gradients. The conditions for breakto sta are related principally to the swirl ratio and the magnitude of the axial Iow, botn peraneters being a function of the developing core flow. The pact nature of the disaipative processes in infuencing the develop-

ment of the core is still not known, but plays an important role in the core flow. and hence in tie time required for breakdown.

The final section shows how to determine the atmospheric dyamic parameters from general meteorological information, using such concepts as Kolmogorov's hypothesis and Monin-Obukhov Similariiy. Operation i equations are presented, and a projected design for an operationa: metejrological data system is given.

## 8. CONCLUSIONS

The general thrust of this report is cincerned with the still unsolved problems of vortex descent and decay in a zeal atmosphere. Considerable theoretical progress has been made in certain aspects of thio problem, in that new analytical results have been devaloped for

1) Time-to-linkage in a turbulent enviromment.
2) Descent in a stratified flow.
3) Cell shapes ir sheared flow.

These theoretical results can only be validated by experimert, and a major recommendation is that the appropriate data be obtained. The analyser developed here and the specific problems idertified, will provide tine proper basis for sound experiments designed to resolve some of these difículties.

For the important area os core bursting, the conclusion is that still no definite operational criteria usist and that more experimental and theoretical reseapeh is reguired. It is possible, however. that atiural core bursting seldom cecurs in the vortex wakes of large airezafand thus an operational criteria is not a praciacal necessity. Furthef night test, and examination of existing test data iy necessary to establimin this hypntitesis.

The problem of the zolicary vortex, and the possibinty that this is caused by a sheared erosswind is raived in this report. This appears to be an important operational consideration, more experimental dista here is badly needed. It is likely that further analyeje, continuing the shearad cell-shape developed in this report, woule afsist in ugierstanding this phenomenon.

The turbulent entrainment in the descending vortex pair is discussed at length. Here again, experiments are needed, and the report defines appropriate parameters which would be essential in designing the proper tests to unambiguously determine the turbulent mechanisms.

For the wing aerodynamics defining the vortex configuration, the aircraft ciynamics defining the hazard, and the meteorology defining the atmospheric dynamics, this report gives specific operational equations which can be used in predictive models. It is believed that in these fields the state of the art is sufficiently well developed that engineering calculations can be made with reasonable confidence.

An assessment of the situation with respect to the prediction of vortex wake descent and decay is as follows:

1) Wing aerodynamics defining vortex configuration, aircraft dynamics defining hazard, and meteorology defining atmospheric dynamics are reasonably well understood.
2) Vortex descent and decay in a stratified flow is still not properly understood.
3) Crow Instab lity is well understood, and equations presented here give a rational approach to predicting time to-linkage as a function of turbulence.
4) Corebursting is still very poorly understood and qualitative equations are lacking, however core bursting may not be as important as Crow Instability for vortices from large aircraft.
5) Unsymmetric effects causing vortex tilting and unsymmetric breakdown, resulting in a long-lived sclitary vortex have been observed; and are not understood. These may be very important for hazard predict,on.
6) Further analytical research and Alight test experimentation $\dot{\lambda} s$ very definitely required to resolve the problems above, both at altitude and in ground effect. This report defines a basis for the rational and effective design of continued researck.

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